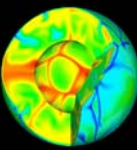




The Earth's Structure from Travel Times



Spherically symmetric structure: PREM

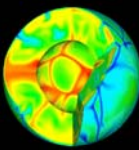
- Crustal Structure
- Upper Mantle structure
 - Phase transitions
 - Anisotropy
- Lower Mantle Structure
 - D''
- Structure of the Outer and Inner Core

3-D Structure of the Mantle from Seismic Tomography

- Upper mantle
- Mid mantle
- Lower Mantle



Spherically Symmetric Structure

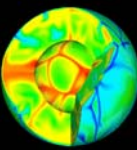


Parameters which can be determined for a **reference** model

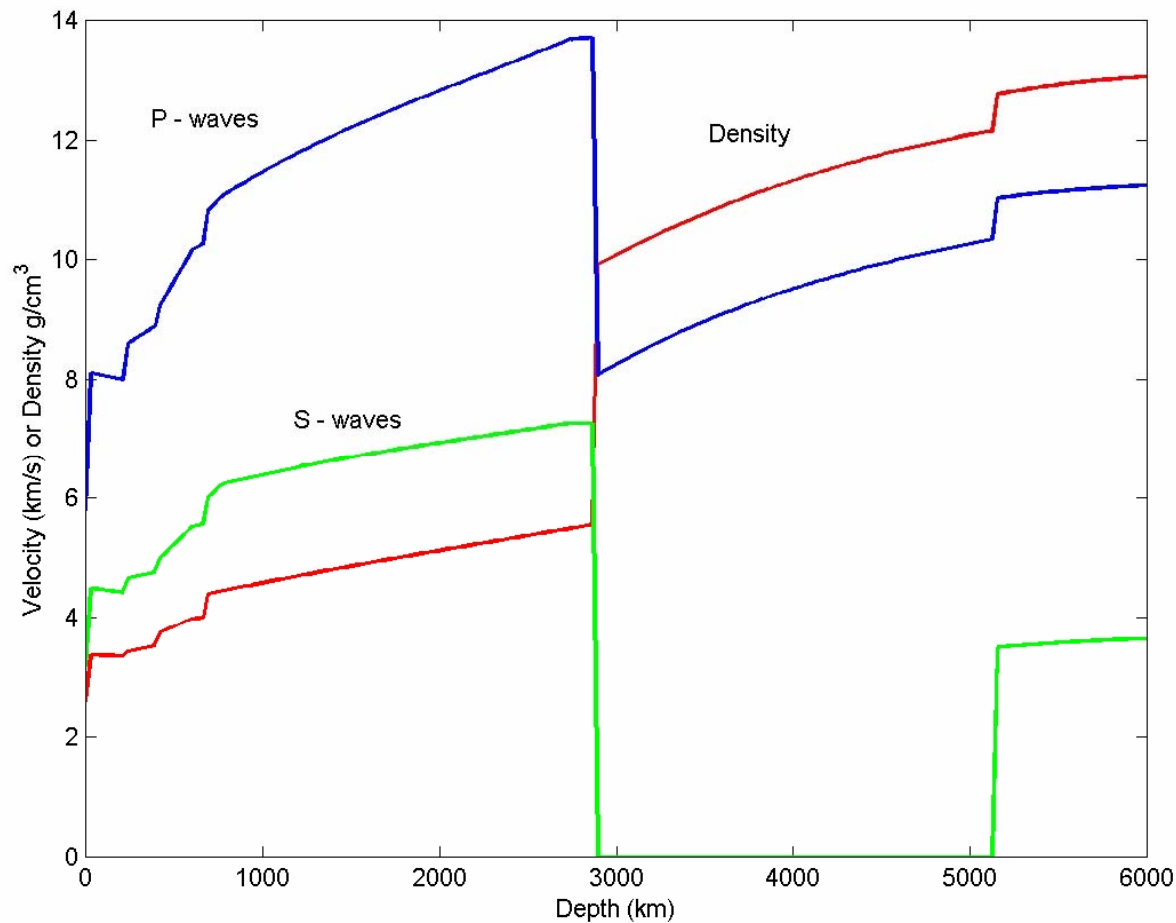
- P-wave velocity
- S-wave velocity
- Density
- Attenuation (Q)
- Anisotropic parameters
- Bulk modulus K_s
- rigidity μ
- pressure
- gravity



PREM: velocities and density

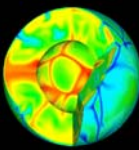


PREM: Preliminary Reference Earth Model (Dziewonski and Anderson, 1981)

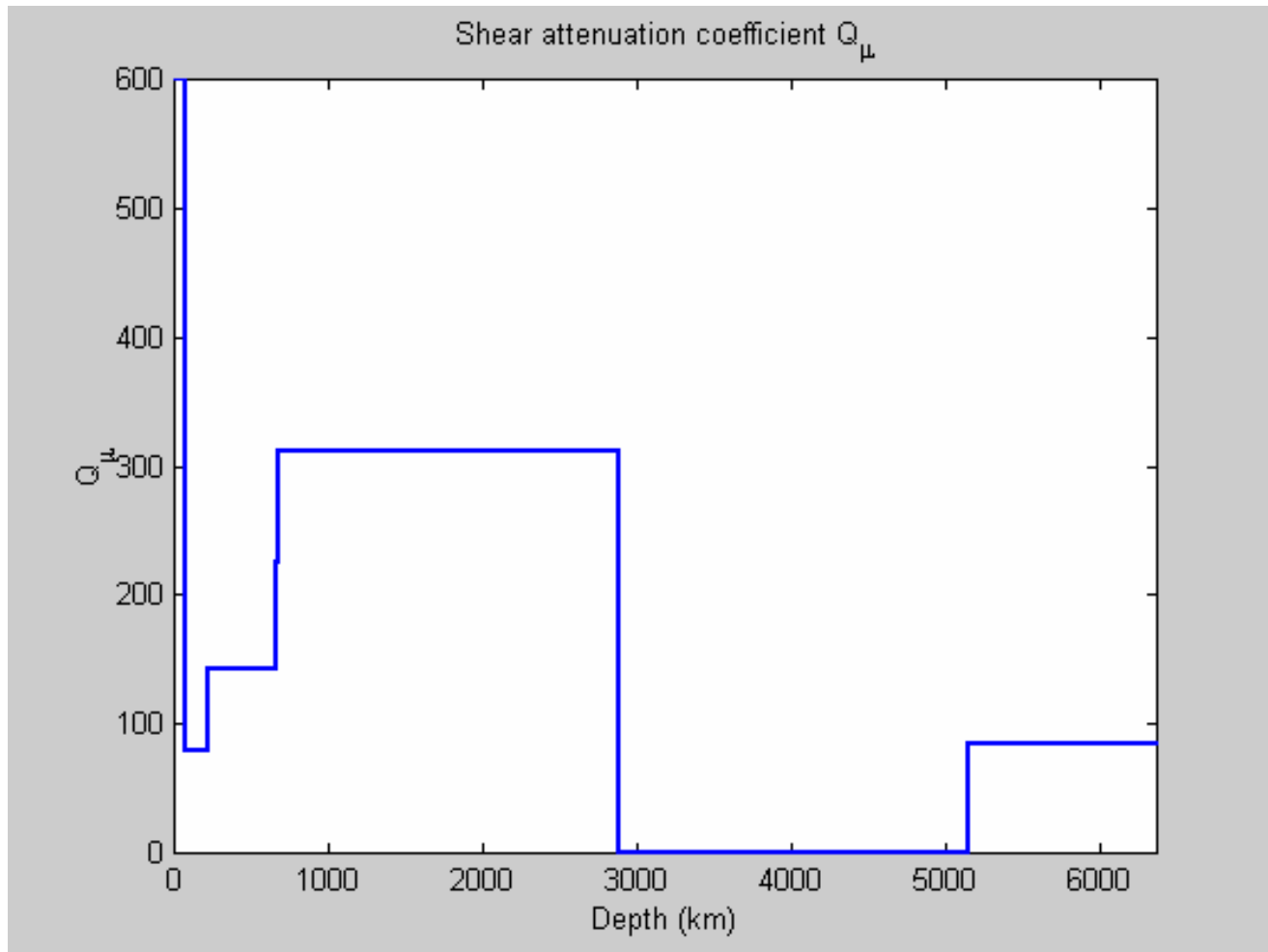




PREM: Attenuation



PREM: Preliminary Reference Earth Model (Dziewonski and Anderson, 1981)





Earth's Regions and Fractional Mass

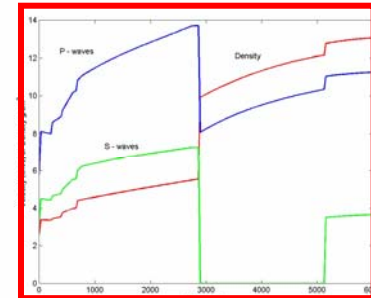
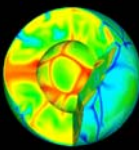
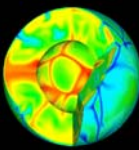


TABLE 3-1
Summary of Earth Structure

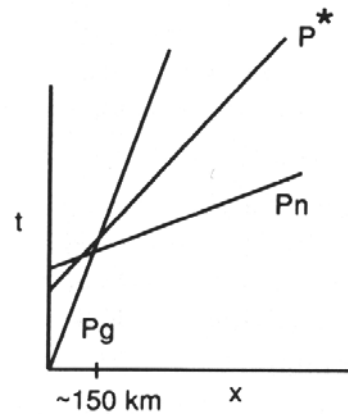
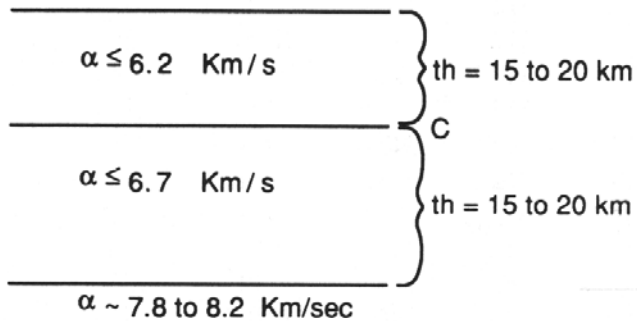
Region	Depth (km)	Fraction of Total Earth Mass	Fraction of Mantle and Crust
Continental crust	0–50	0.00374	0.00554
Oceanic crust	0–10	0.00099	0.00147
Upper mantle	10–400	0.103	0.153
Transition region	400–650	0.075	0.111
Lower mantle	650–2890	0.492	0.729
Outer core	2890–5150	0.308	—
Inner core	5150–6370	0.017	—



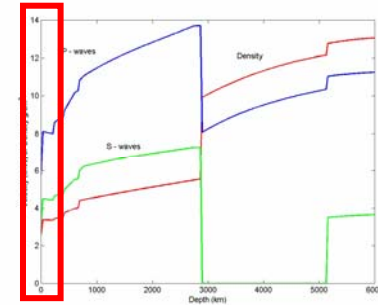
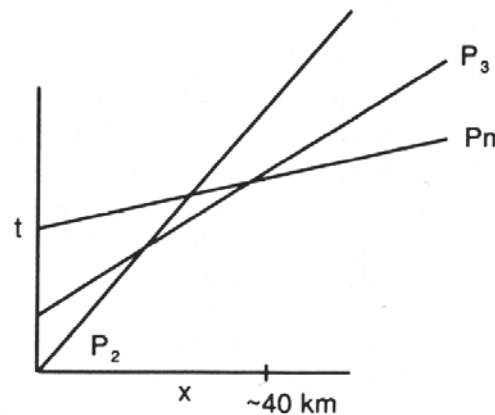
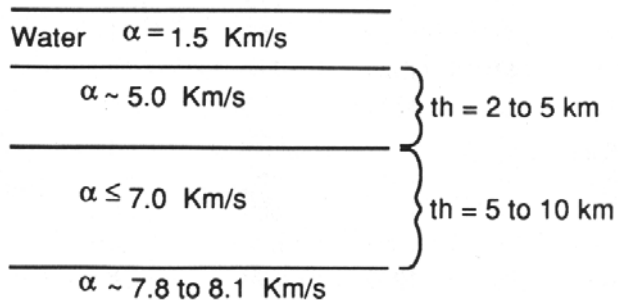
The Earth's Crust: Travel Times



a



b



Continental crust (a) and oceanic crust (b) with corresponding travel-time curves



The Earth's Crust: Minerals and Velocities

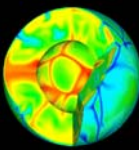
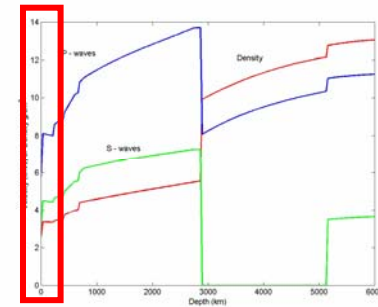


TABLE 3-3

Average Crustal Abundance, Density and Seismic Velocities of Major Crustal Minerals

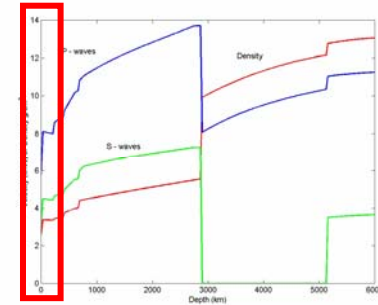
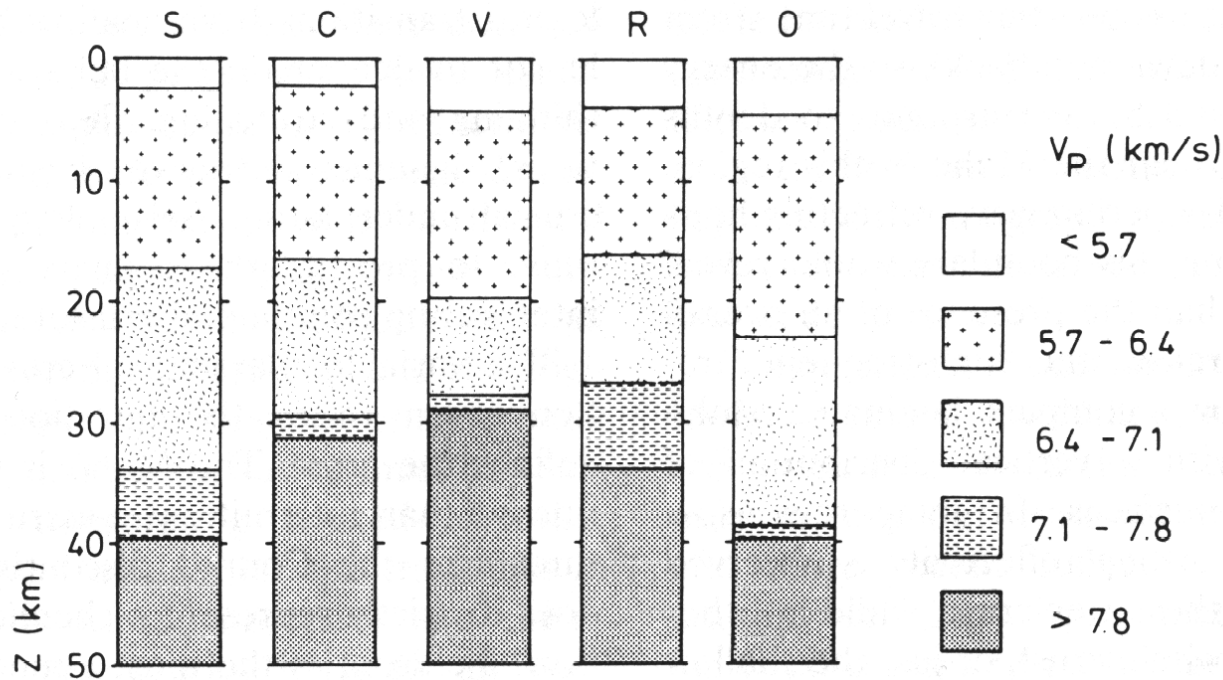
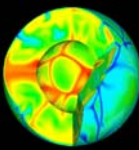
Mineral	Volume percent	ρ (g/cm ³)	V_p (km/s)	V_s (km/s)
Quartz	12	2.65	6.05	4.09
K-feldspar	12	2.57	5.88	3.05
Plagioclase	39	2.64	6.30	3.44
Micas	5	2.8	5.6	2.9
Amphiboles	5	3.2	7.0	3.8
Pyroxene	11	3.3	7.8	4.6
Olivine	3	3.3	8.4	4.9



Average crustal abundance, density and seismic velocities of major crustal minerals.



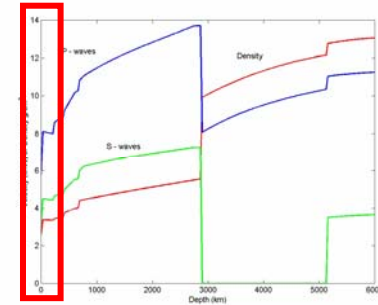
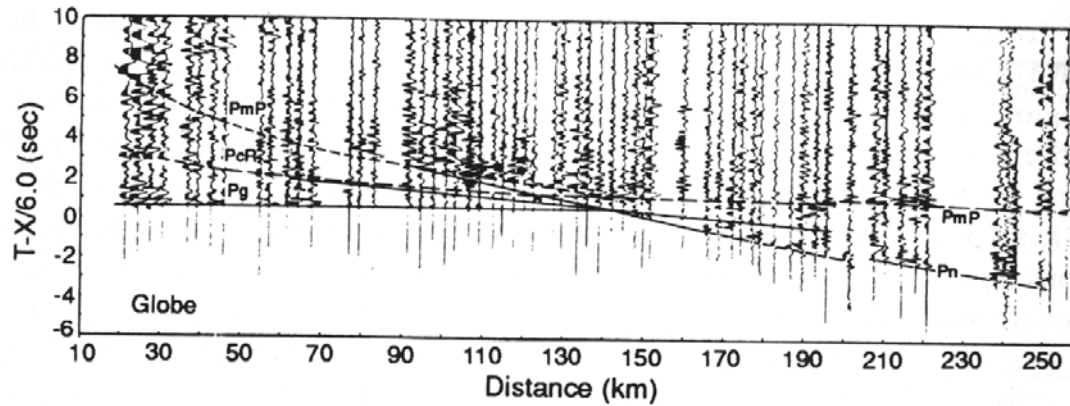
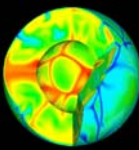
The Earth's Crust: Crustal Types



S shields, C Caledonian provinces, V Variscan provinces, R rifts, O orogens

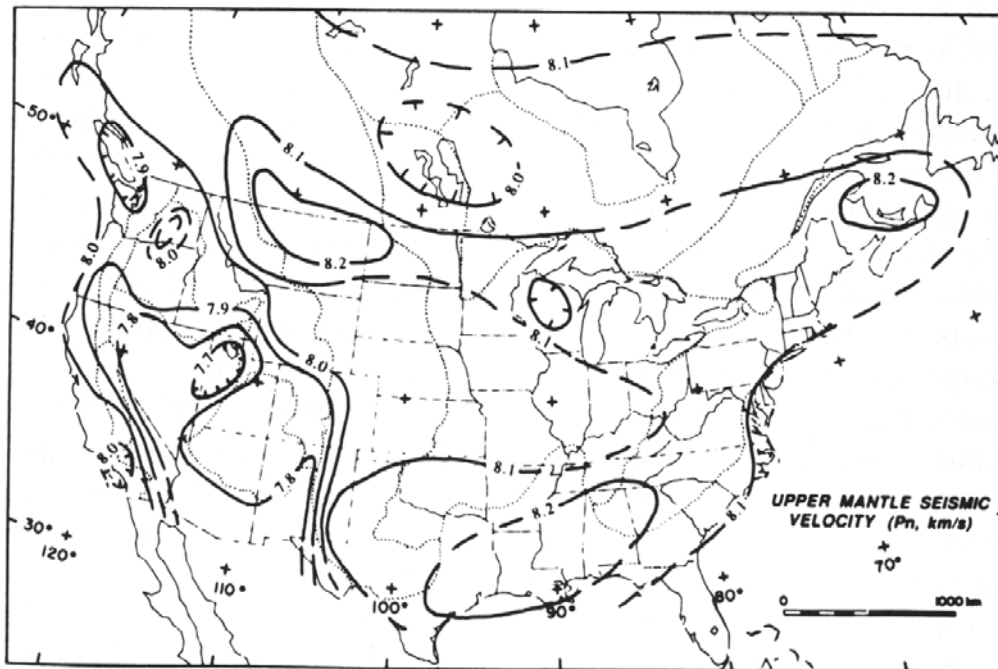


The Earth's Crust: Refraction Studies



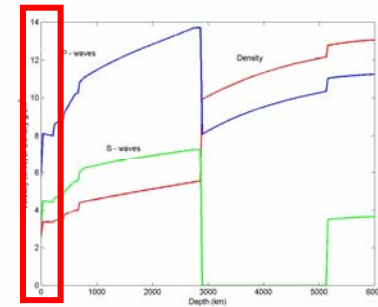
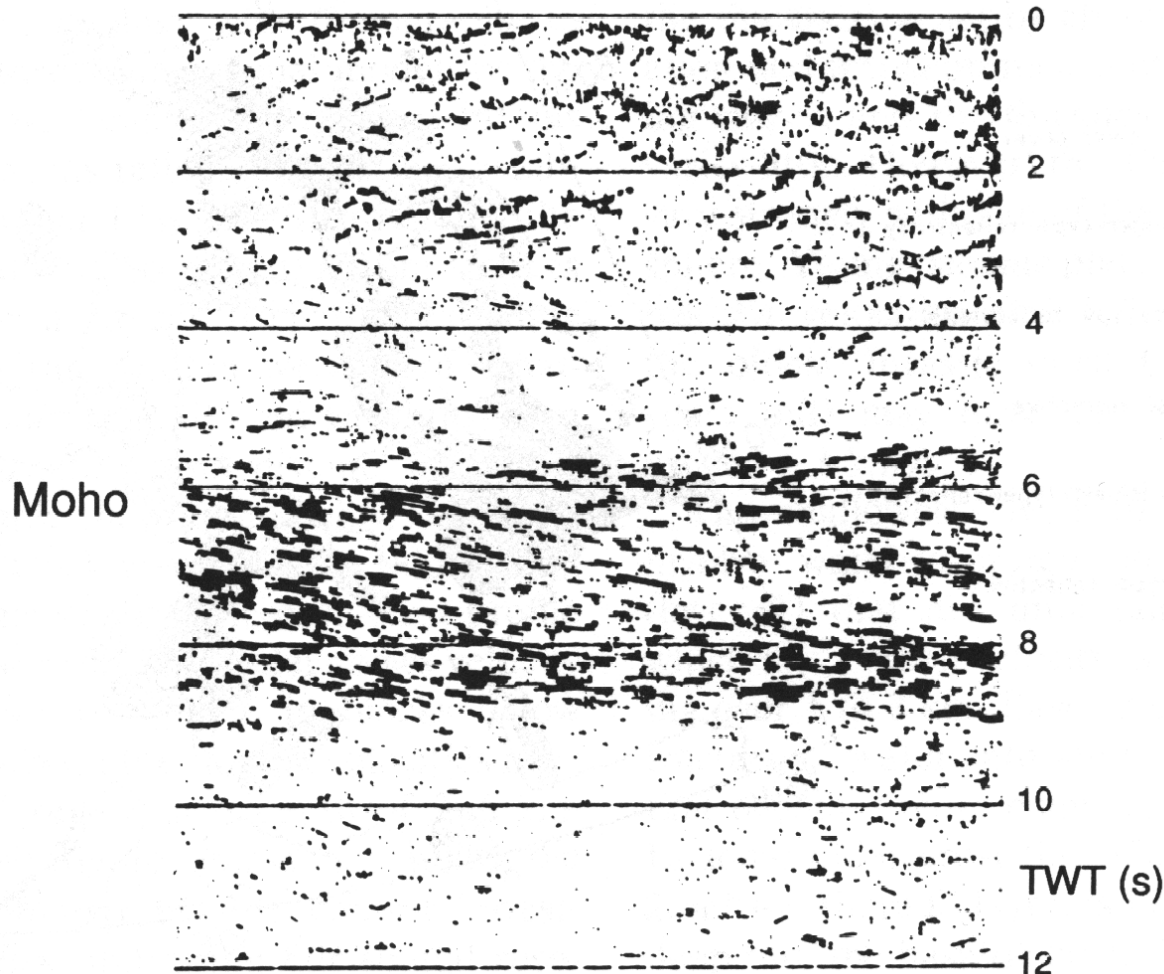
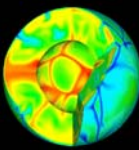
Refraction profiles across North America, (reduction velocity 6km/s) all the determination of lateral velocity variations:

PmP Moho reflection
Pn Moho refraction
Pg direct crustal wave





The Earth's crust: Crustal Types

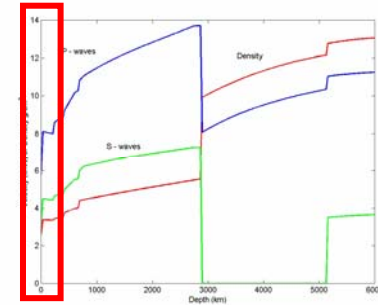
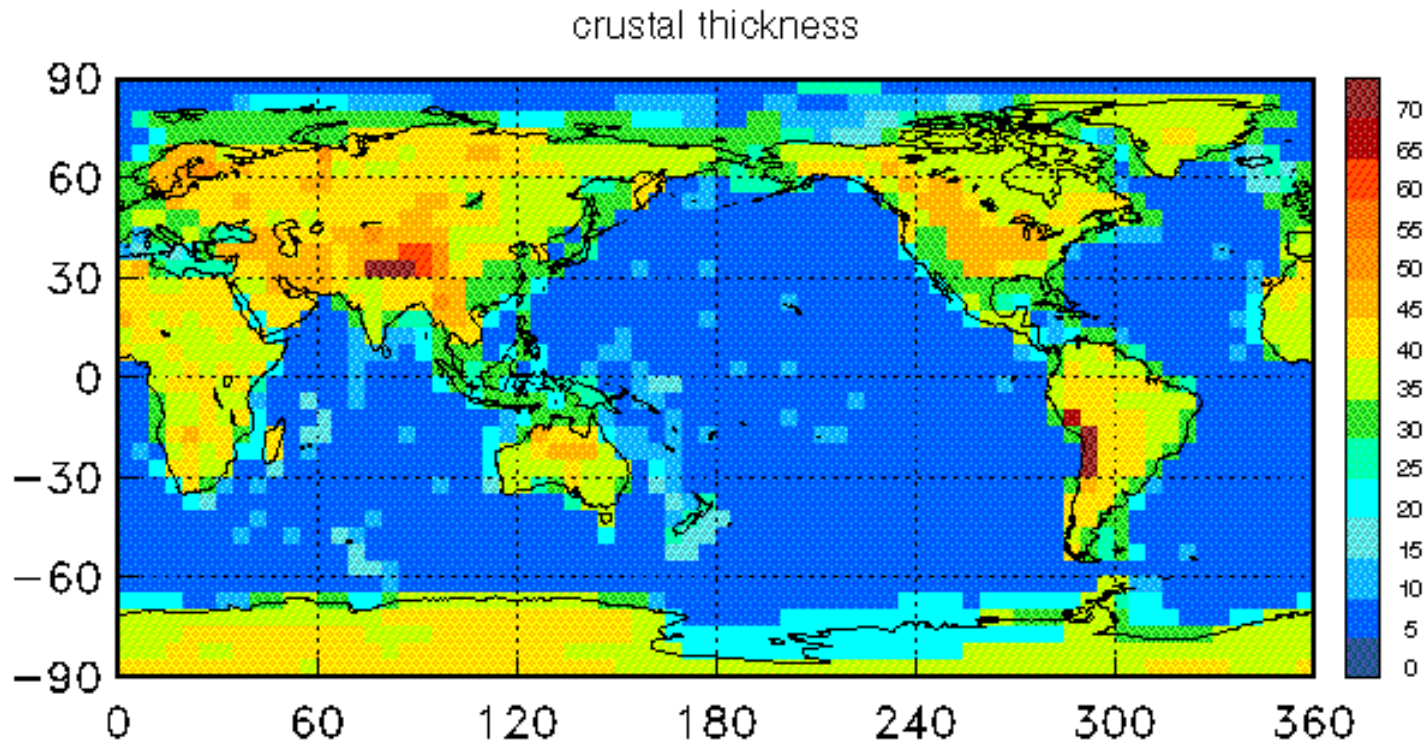
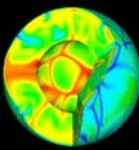


Reflection data often show a highly reflective lower crust. This may indicate fine layering or lamination, some transition from crust to upper mantle.

TWT two-way traveltimes



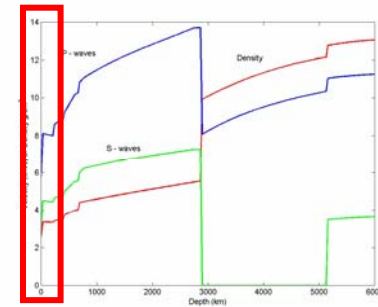
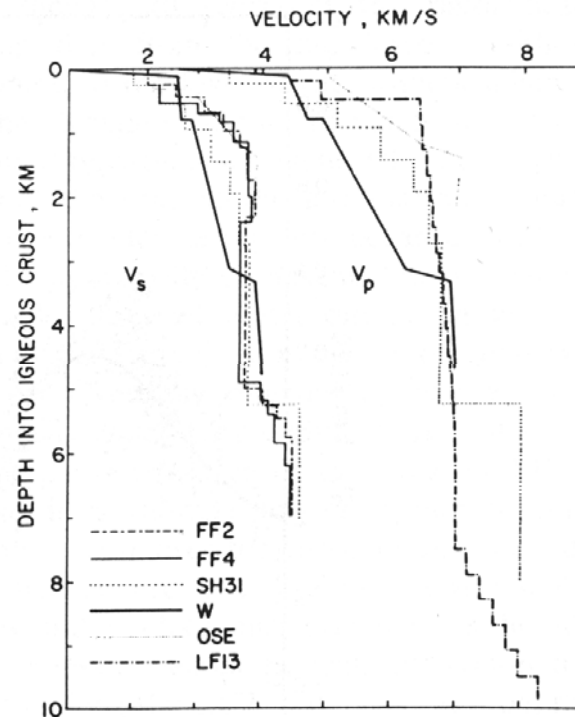
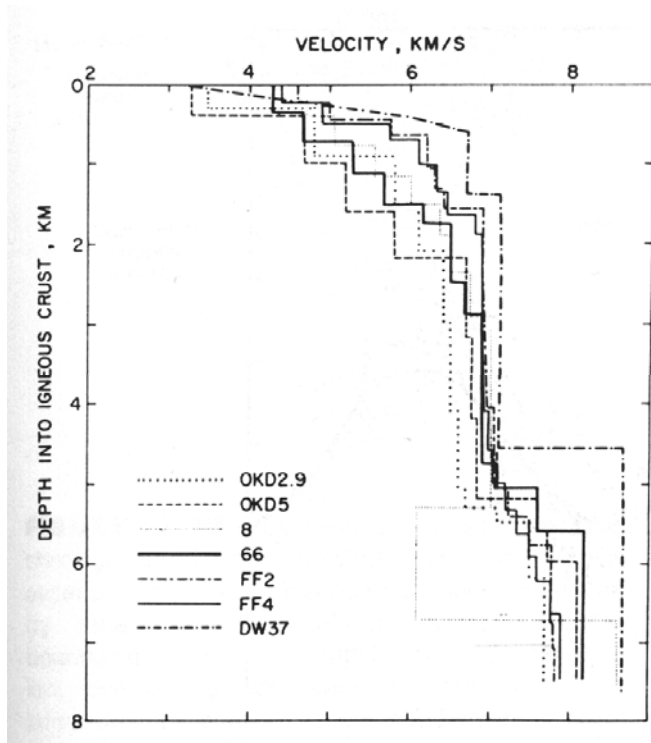
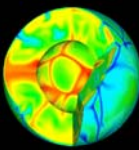
The Earth's crust: Crustal Types



Recently compiled world-wide crustal thickness (km) indicates cratonic areas and mountain ranges with active tectonics. These data are important to correct travel times regionally, i.e. calculate the contribution of crustal thickness to a teleseismic travel-time perturbation.



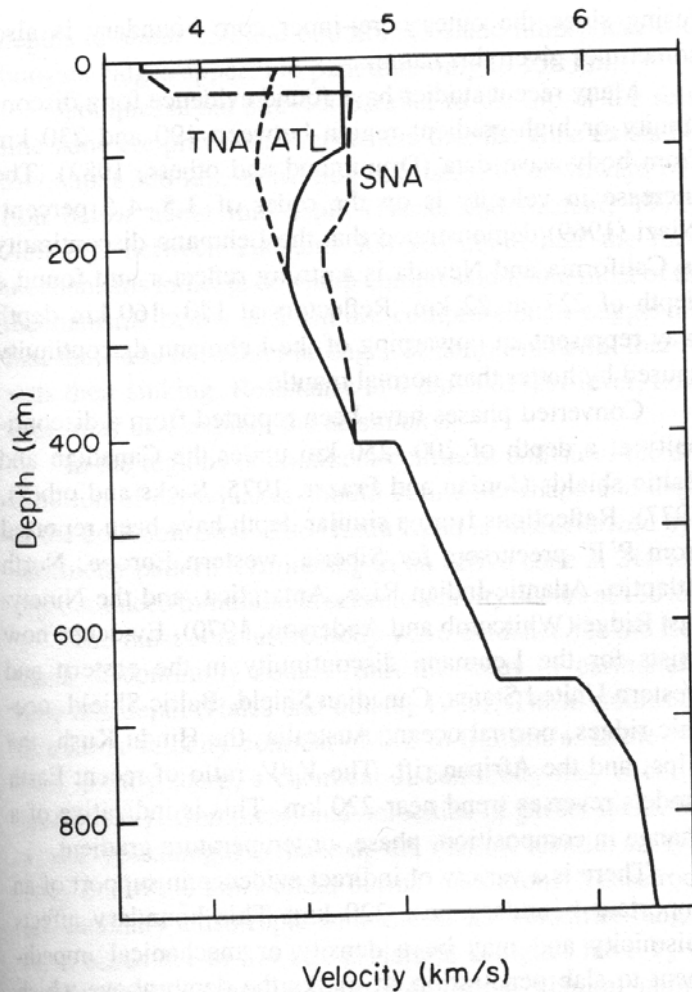
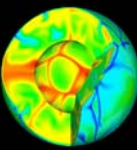
The Earth's crust: Crustal Types



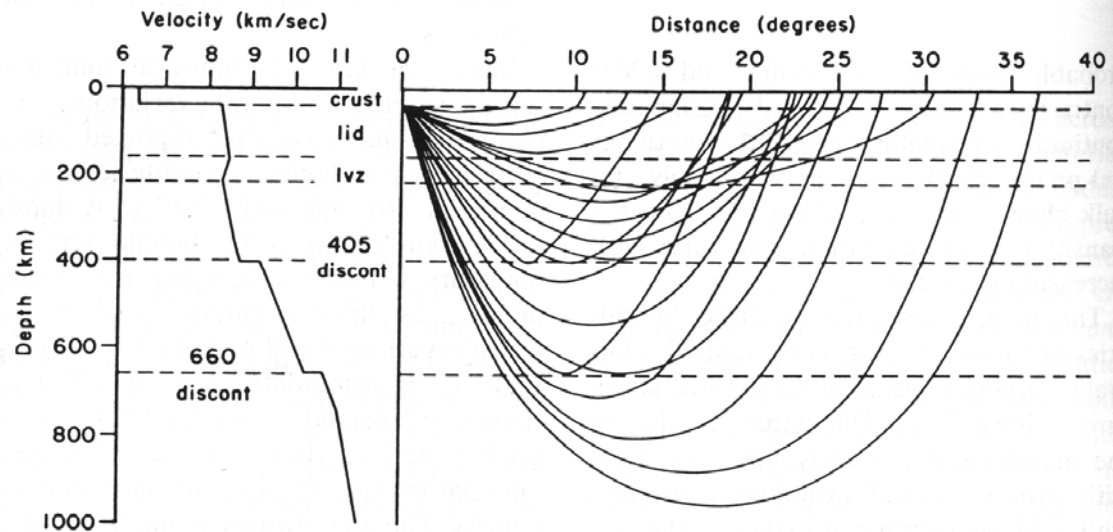
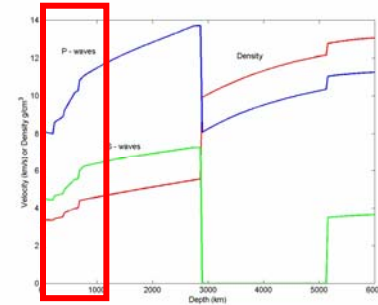
Left: Crust P-velocity profiles for young (<20 million year) oceanic basin structures. Right: Crustal P and S velocities for oceanic regions older than 20 million years.



The Earth's Upper Mantle: Athenosphere



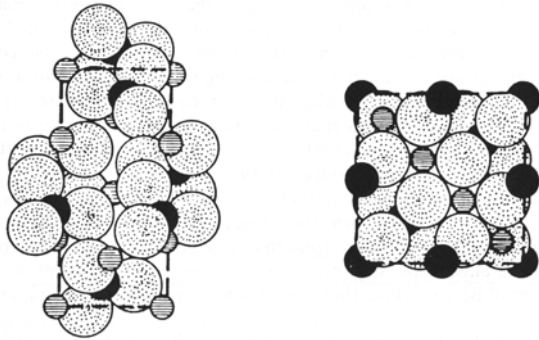
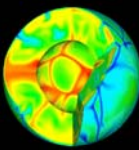
The high-velocity lid above the low velocity zone (**asthenosphere**) is called the **lithosphere**.



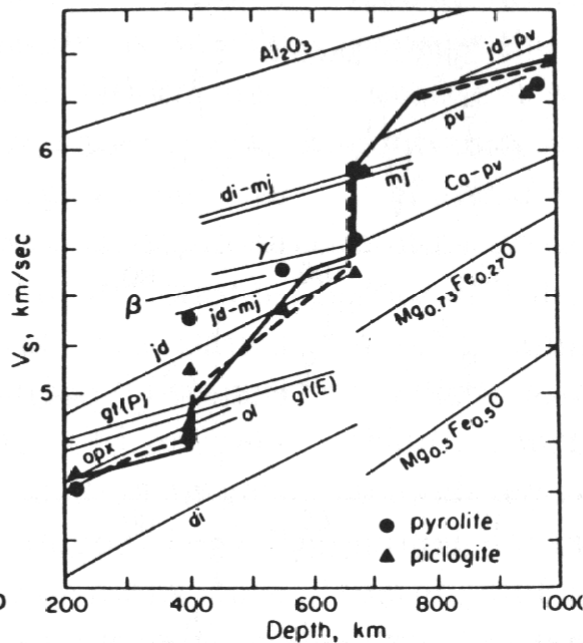
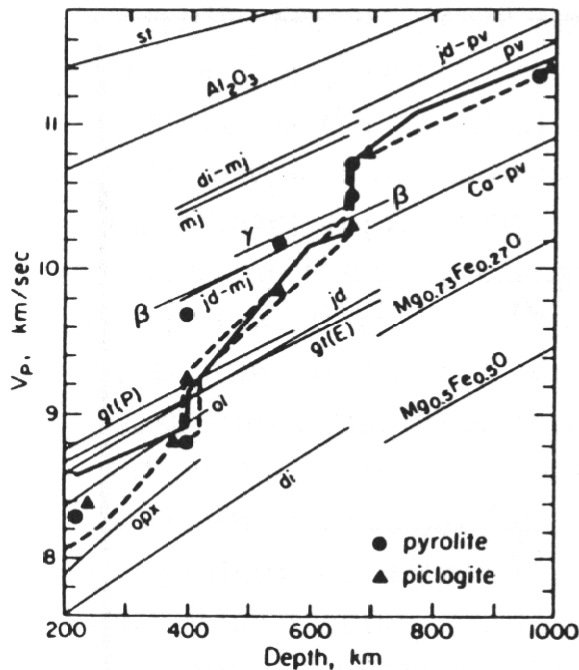
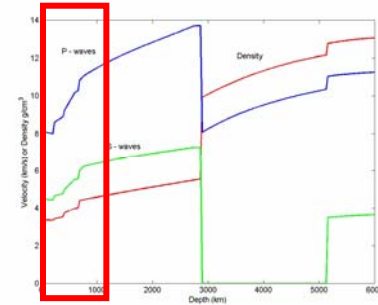
The upper-mantle velocity structure leads to complex ray paths.



Upper Mantle: Phase transitions



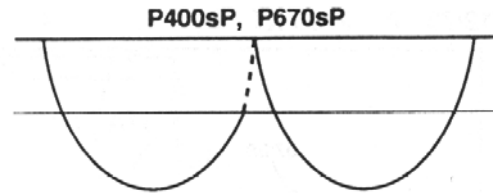
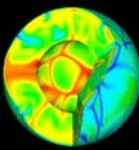
Upper mantle discontinuities (e.g. 410km) are caused by phase transitions (left: low pressure olivine, right: high pressure β -spinel)



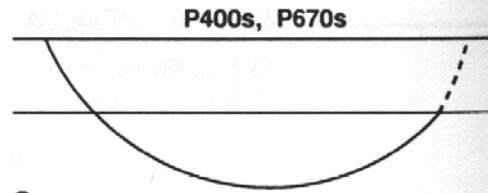
Various upper mantle seismic models and experimental results for minerals and mineral assemblages.



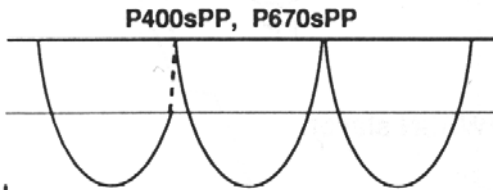
Upper Mantle: Discontinuities



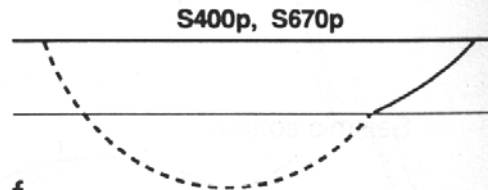
a



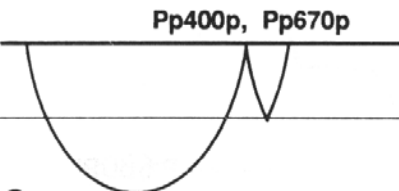
e



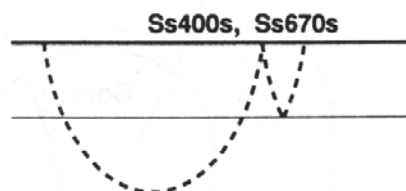
b



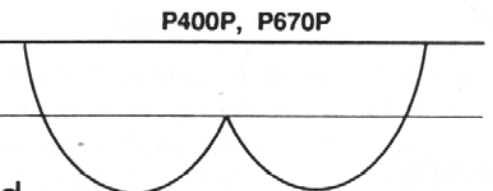
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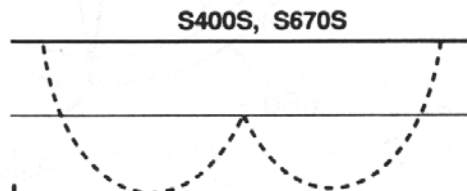
c



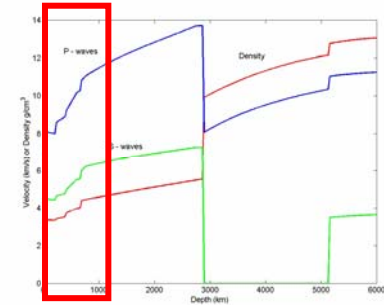
g



d



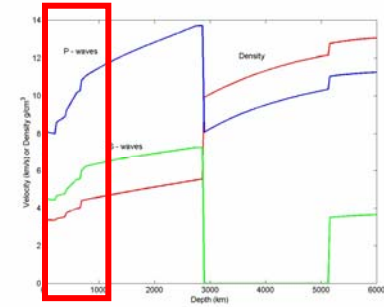
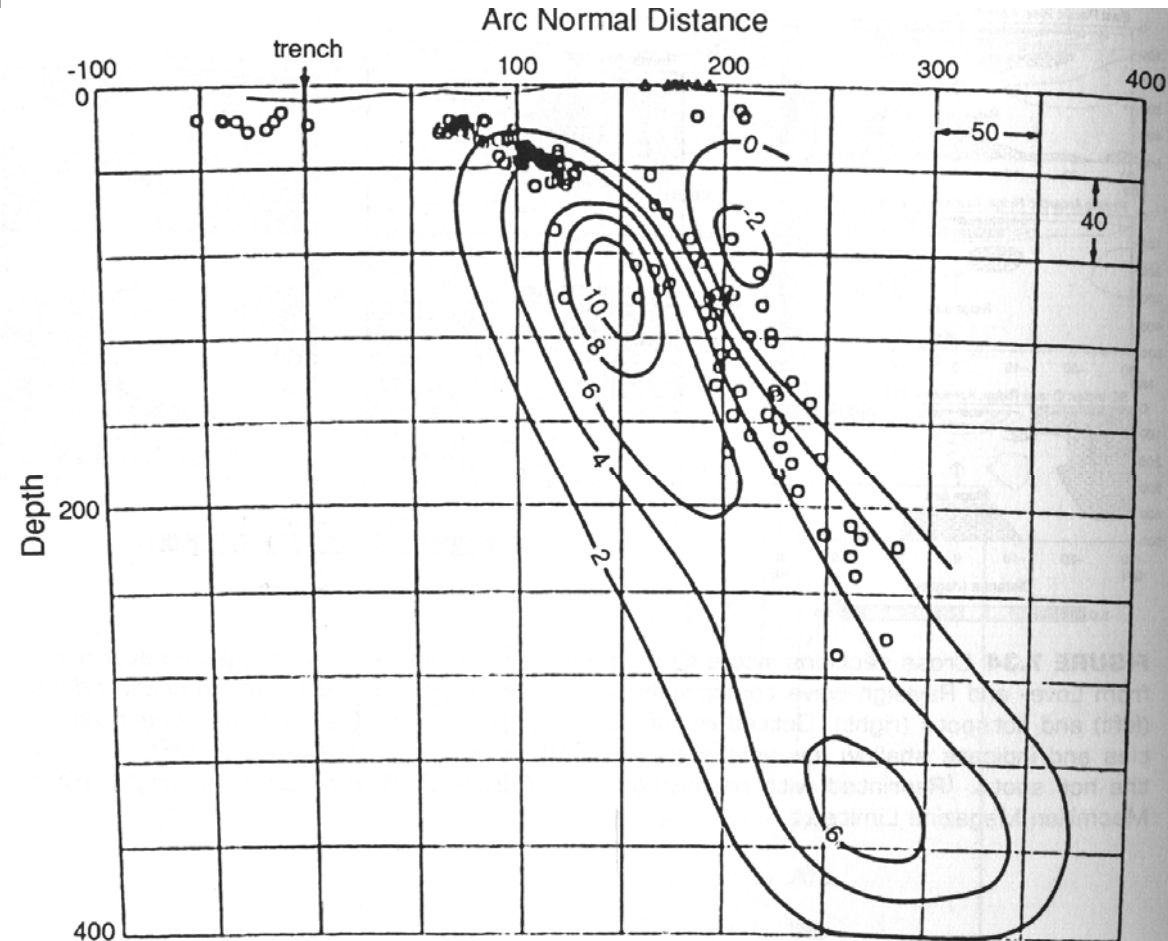
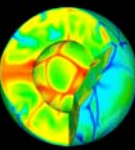
h



Various reflections from upper mantle discontinuities are being used to investigate the structural details of the transition zones (e.g. vertical gradients, thickness of transition zone, topography of discontinuities, etc.)



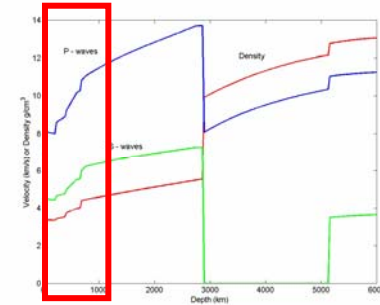
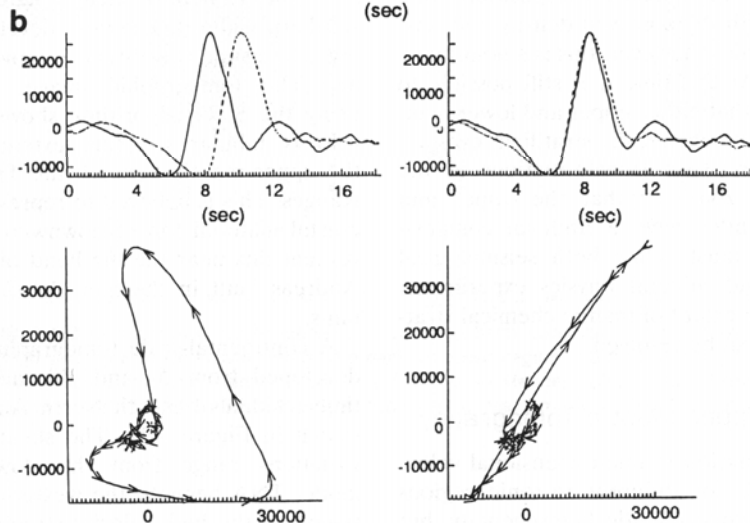
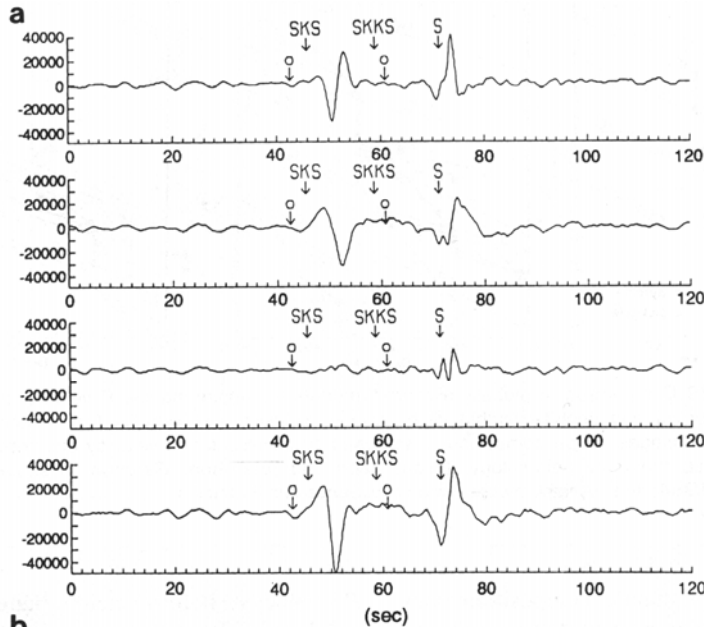
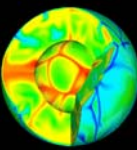
Upper Mantle: Phase transitions



The location of seismic source within high velocity anomalies indicates downgoing slab structures. **Where do earthquakes seem to happen preferentially?**



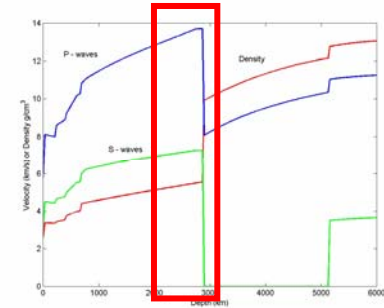
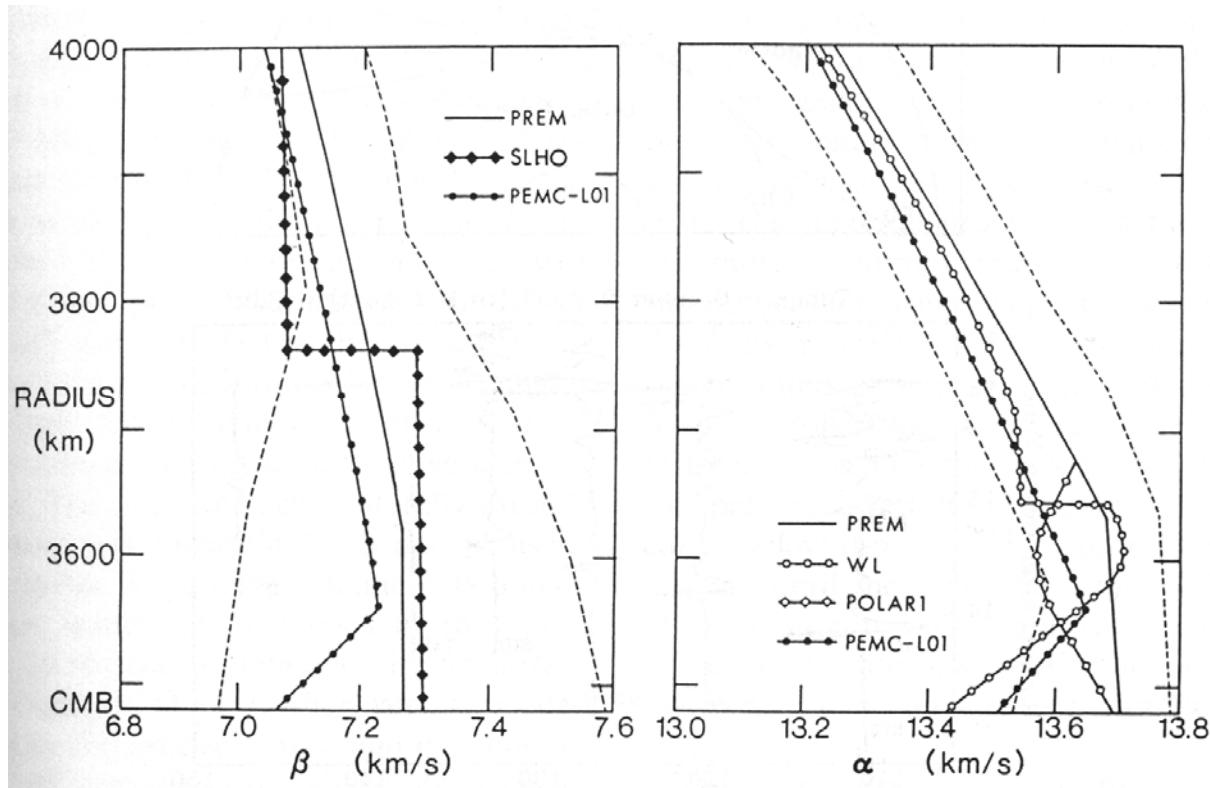
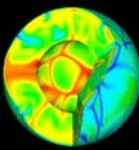
Upper Mantle: Anisotropy



Shear wave splitting of the SKS phase indicates seismic anisotropy in the upper mantle. The alignment of the anisotropic symmetry system is thought to be correlated with tectonic plate motion.



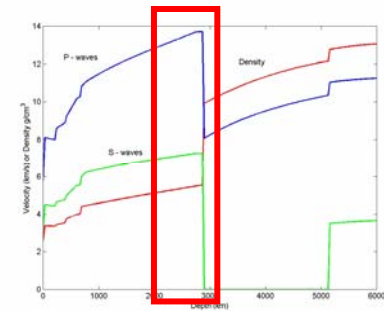
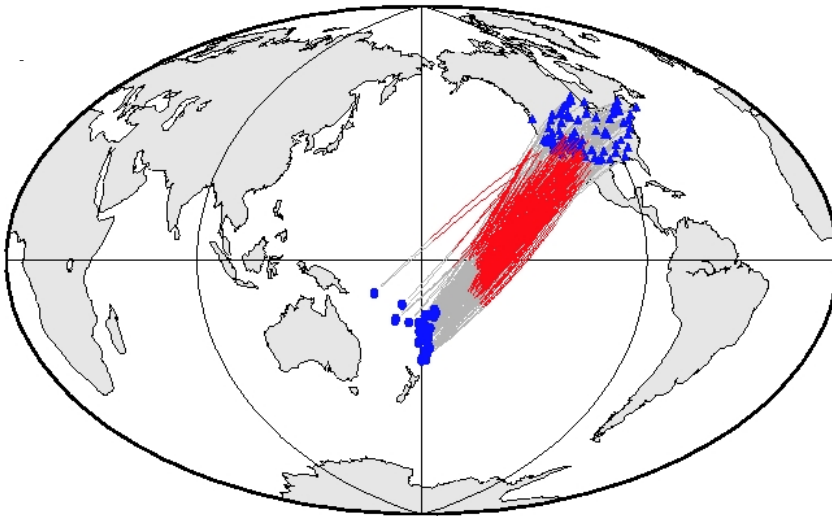
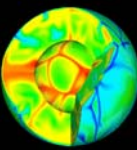
Lower Mantle: D''



The mid-mantle shows little lateral heterogeneity. The lowermost mantle (D'') has strong (possibly >10%) lateral velocity perturbations. This may originate in a thermal boundary layer or from subducted lithosphere.



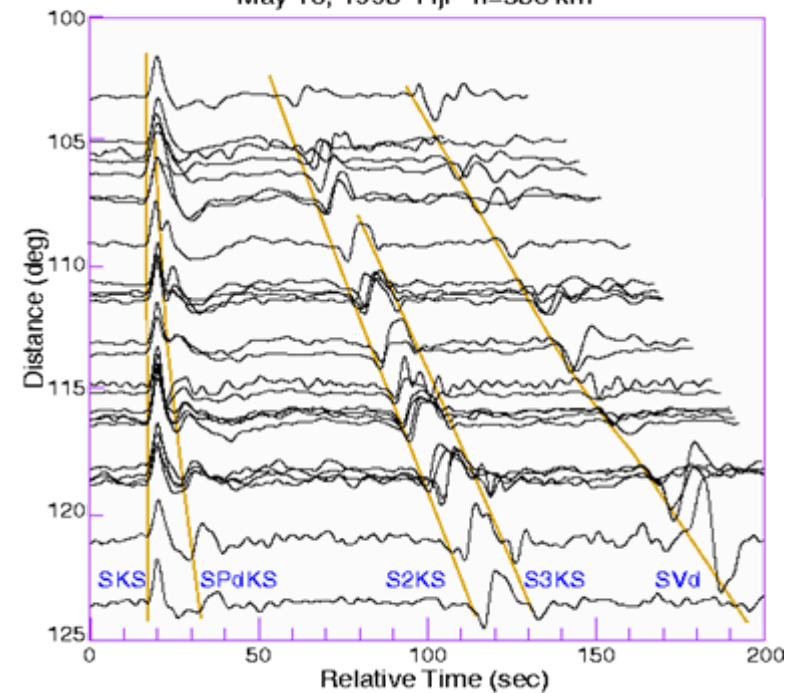
Lower Mantle: Diffracted Waves



The lowermost mantle structure can be studied using waves diffracted at the core-mantle boundary.

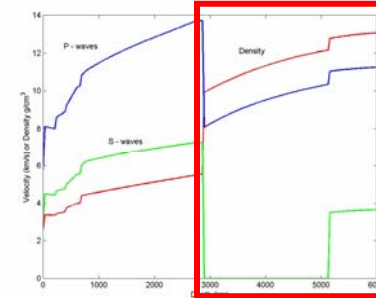
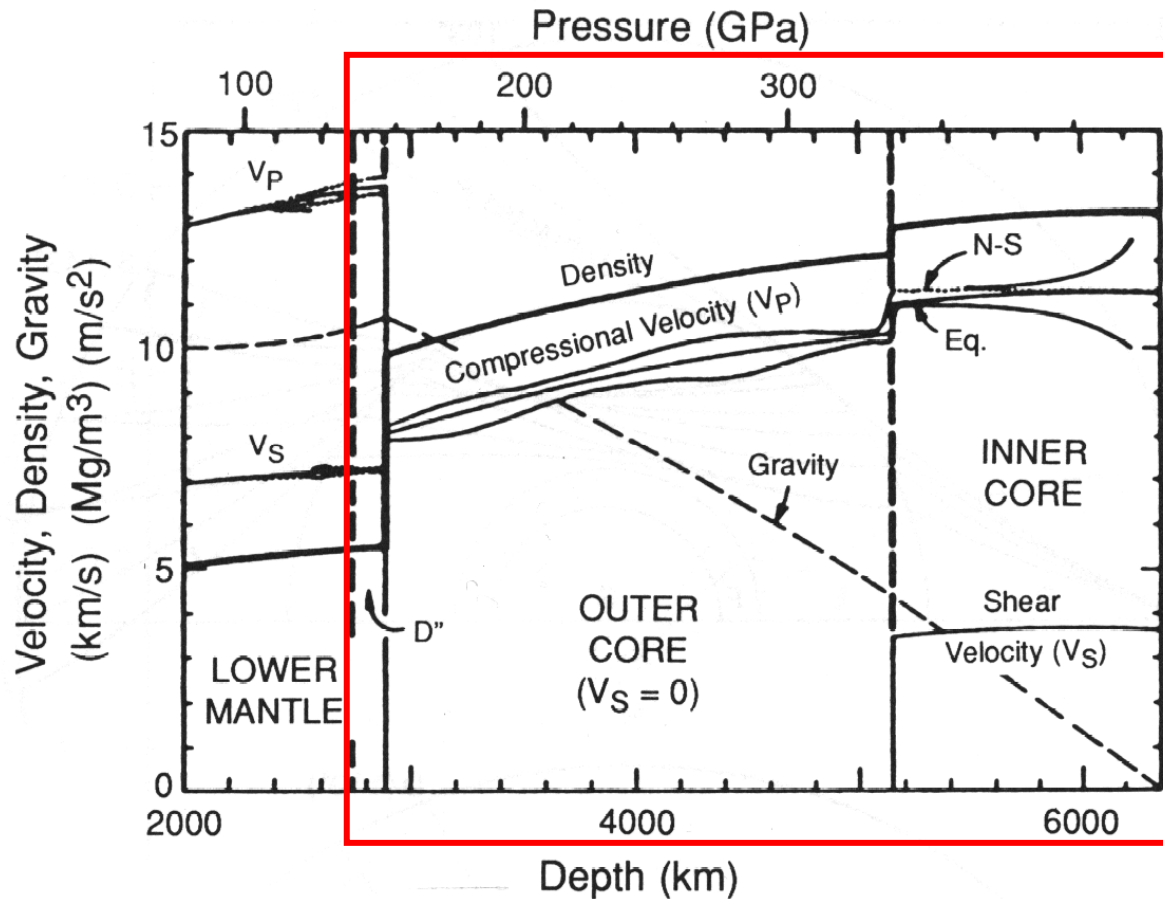
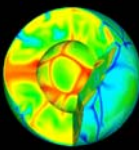
Fiji-Tonga Broadband Data

May 16, 1998 Fiji $h=586$ km





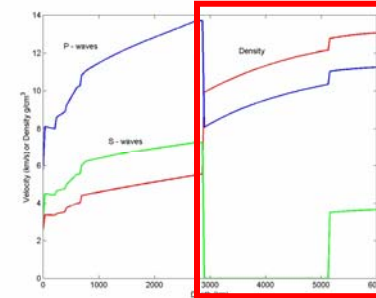
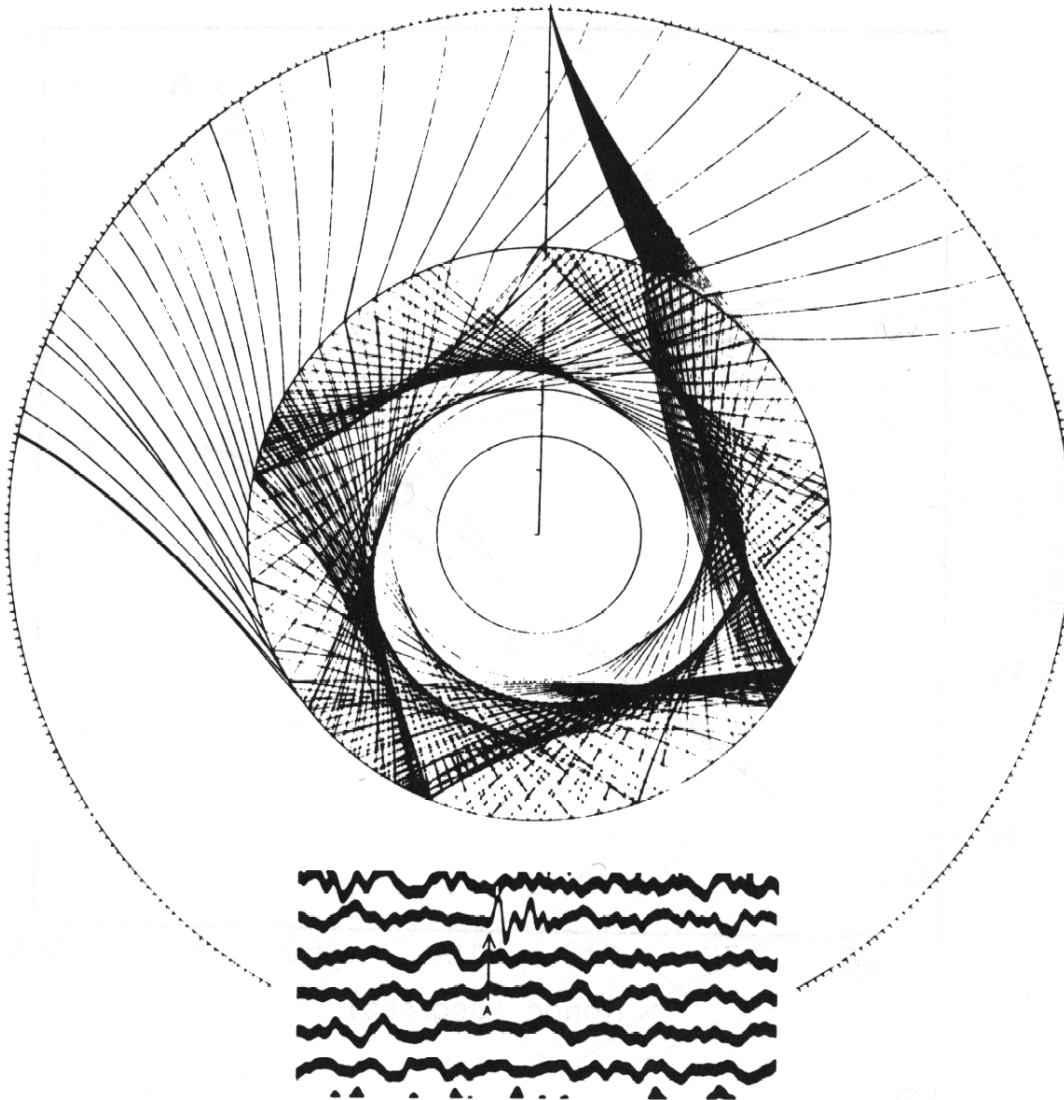
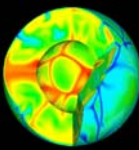
The Earth's Core



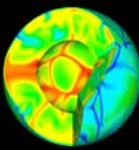
The Earth's inner core shows considerable anisotropy. Time-dependent differential travel times have led to the speculation that the Earth's inner core is rotating faster than the mantle.



The Earth's Core: Multiples

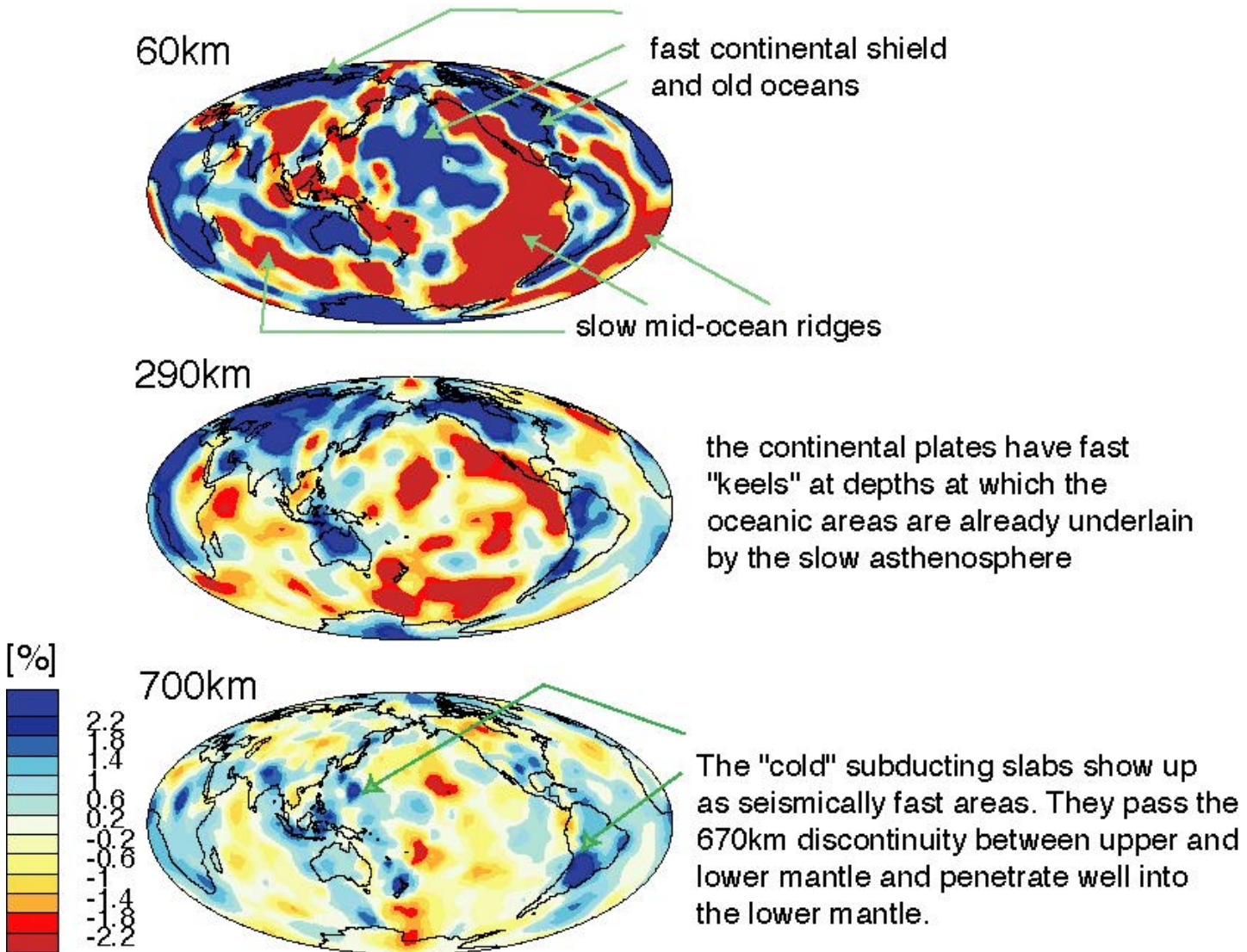


Multiple reflection ray paths PK_nP in the outer core and recording of PK_4P from an underground nuclear explosion.



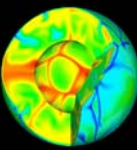
Upper mantle: 3-D structure

SB4L18-Upper Mantle



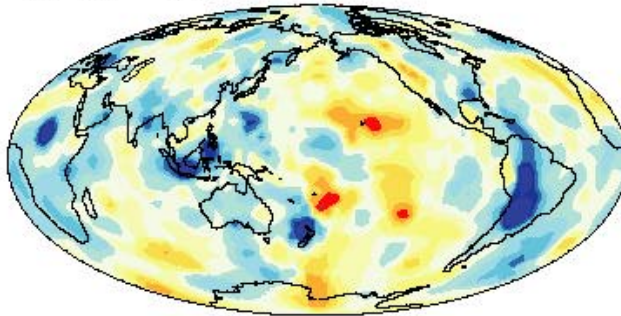


Mid-mantle: 3-D structure



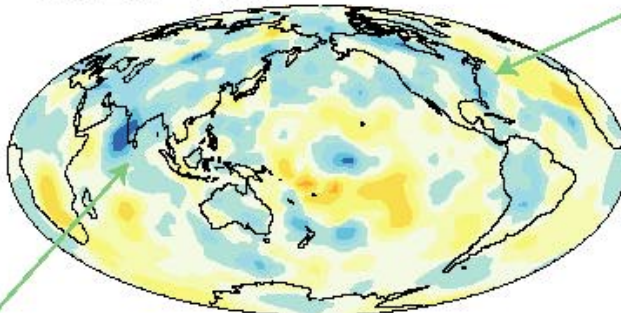
SB4L18-Mid-Mantle

925 km

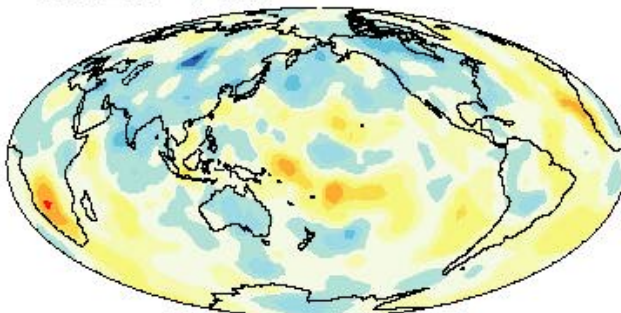


Some of the "cold" subducting slabs can be traced well into the lower mantle. E.g. old Farallon and Tethian subducting slabs.

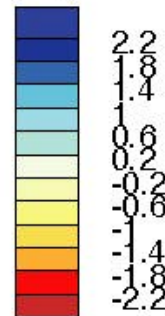
1525 km



1825 km

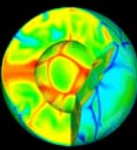


[%]



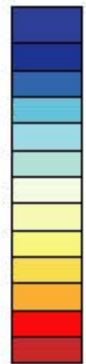


Lower Mantle: 3-D structure



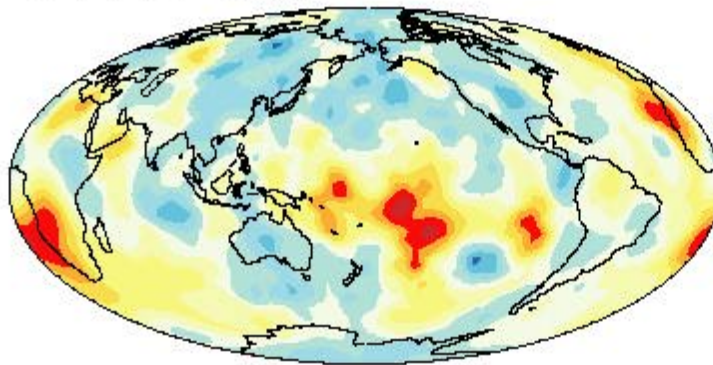
SB4L18-Lowermost Mantle

[%]



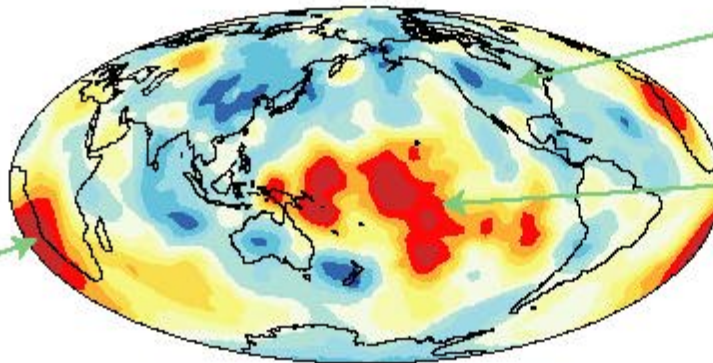
2.2
1.8
1.4
1.0
0.6
0.2
-0.2
-0.6
-1.0
-1.4
-1.8
-2.2

2425 km



The lowermost mantle is dominated by a ring of "fast" material around the Pacific and "slow" material in the Central Pacific and beneath Africa. The fast regions are thought to be the "graveyard of subducting slabs".

2770 km



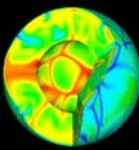
graveyard of slabs

Central Pacific Super Plume

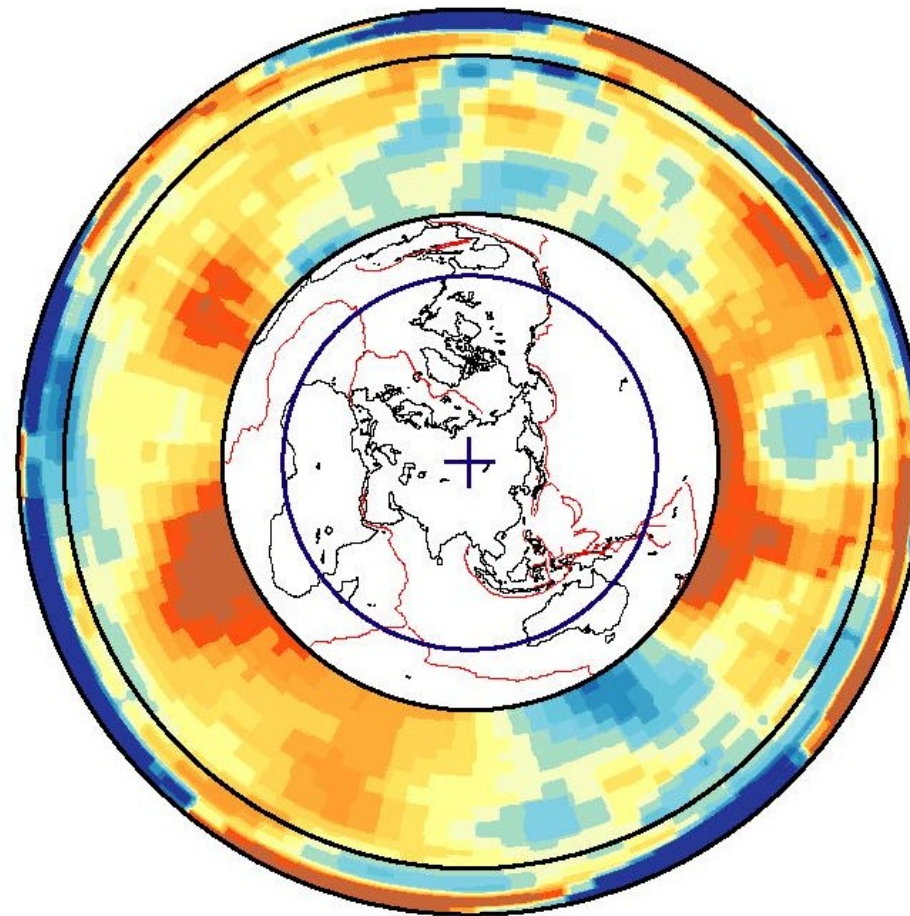
Great African Plume



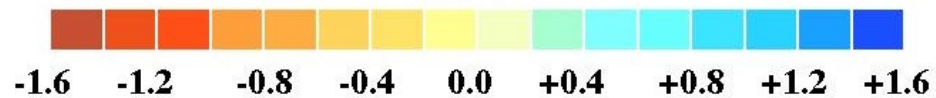
Global Cut: 3-D structure



SB4L18

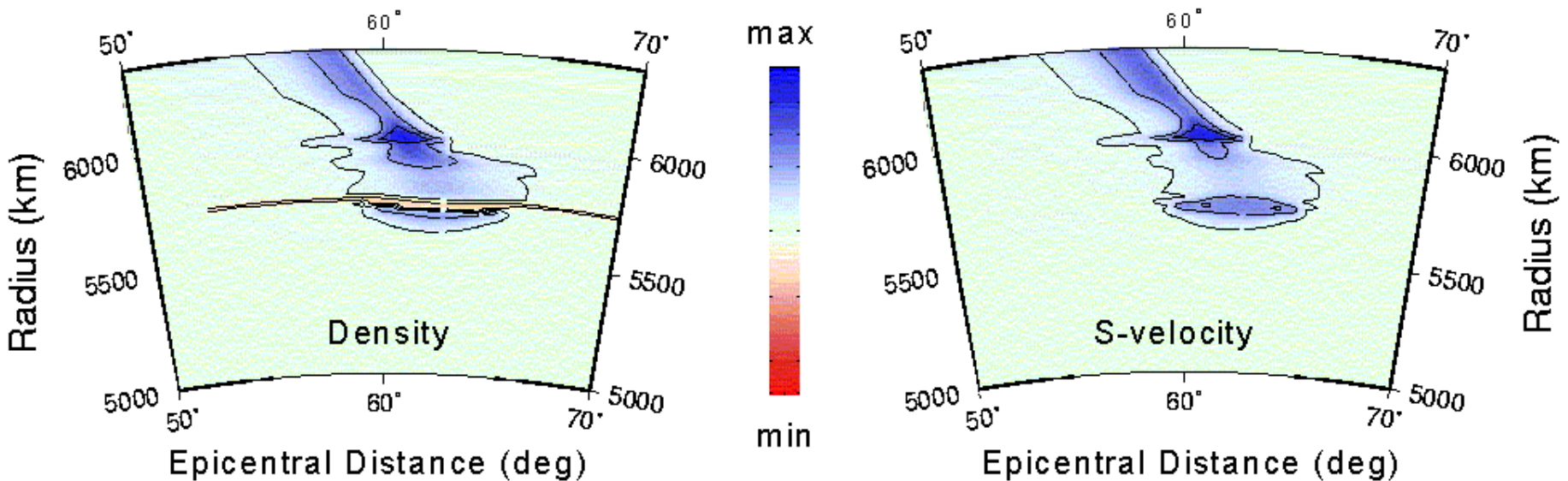
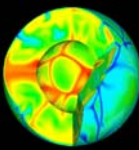


% dV_s/V_s





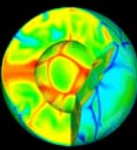
Geodynamic Modelling: Subduction Zones



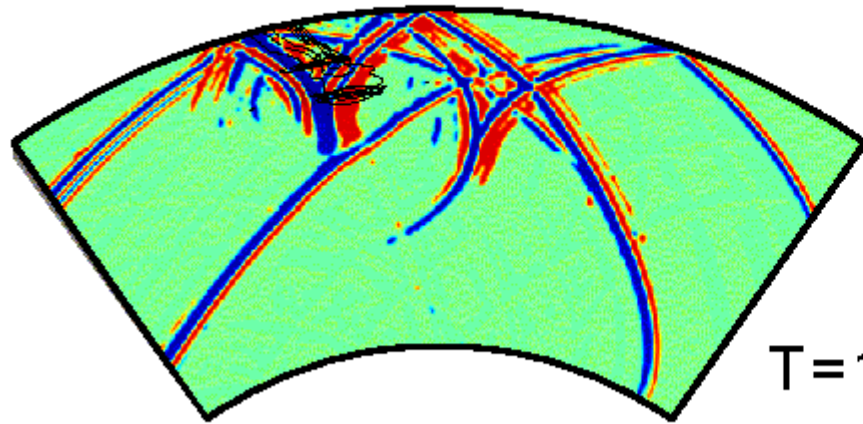
Perturbation of seismic velocity and density for a subducting plate obtained from numerical convection modelling including phase transitions.



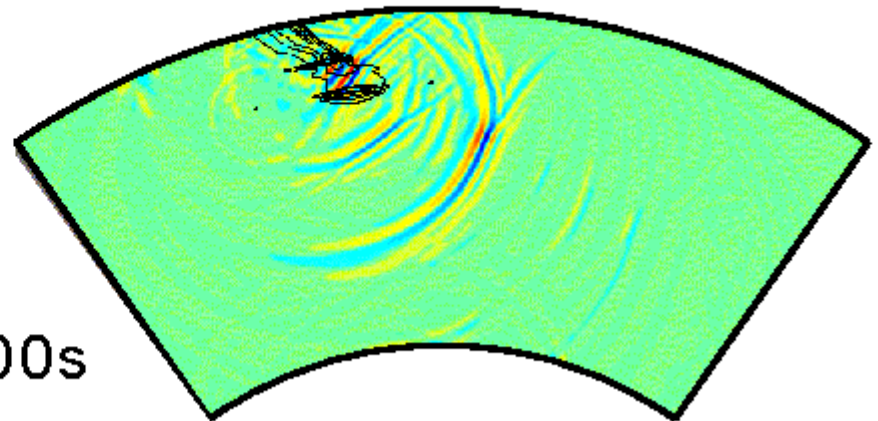
Geodynamic Modelling: Subduction Zones



SH-wavefield



Perturbation

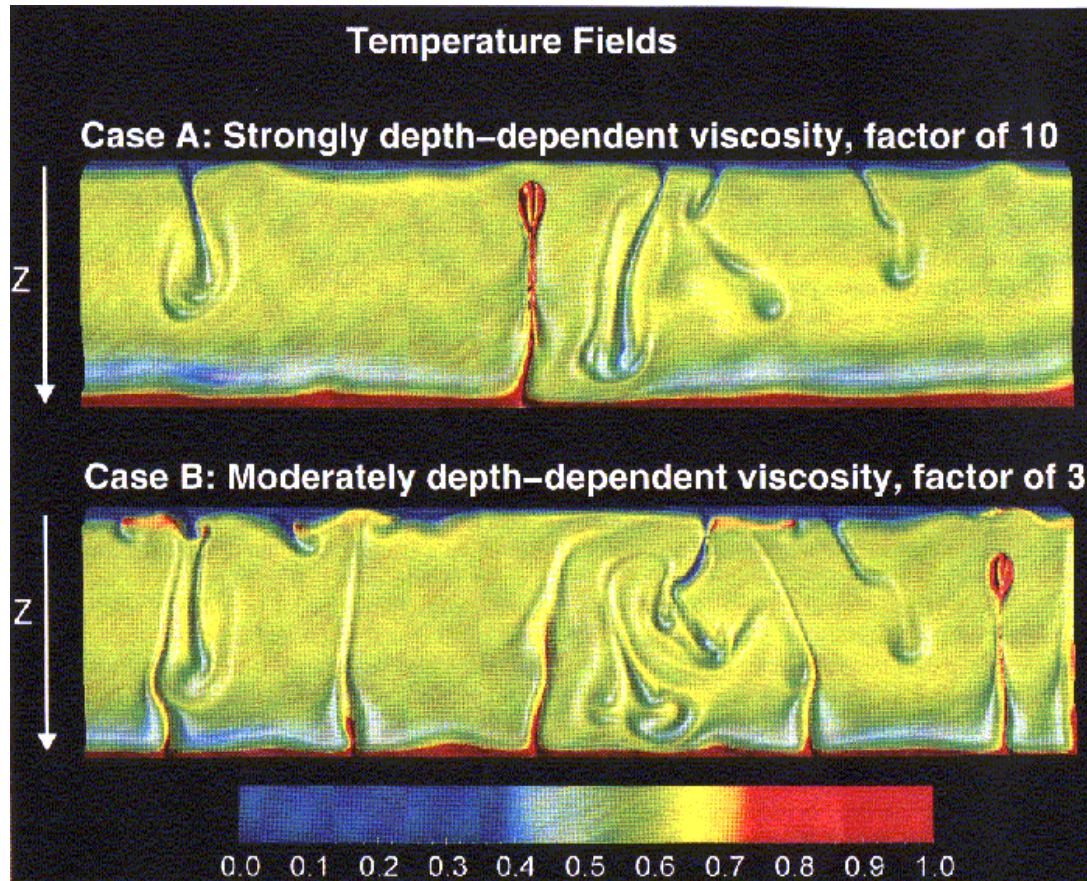
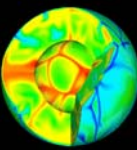


$T=1300s$

Snapshots through subducting slab model and the wavefield perturbation due to the slab. The background model is PREM.



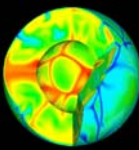
Geodynamic Modelling: Plumes



High-resolution numerical study of plumes and the effects of the mantle viscosity structure.



The Earth's Structure: Summary



The Earth's seismic velocity structure can be determined from inverting seismic travel times (e.g. using the **Wiechert-Herglotz** technique for spherically symmetric media).

The Earth's radial structure is dominated by the **core-mantle boundary**, the **inner-core boundary**, the upper-mantle discontinuities (**410km** and **670km**) and the crust-mantle transition (**Moho**).

The 3-D structure of the Earth's interior can be determined by inverting the travel-time perturbations with respect to a spherically symmetric velocity model (e.g. PREM). The **positive** and **negative** velocity perturbations are thought to represent **cold** (**dense**) or **hot** (**buoyant**) regions, respectively.

There is remarkable correlation between fast regions and subductin zones as well as slow regions with hot-spot (plume) activity.