# Dynamic Rupture Along Bimaterial Interfaces in 3D

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Abstract. We perform numerical simulations of dynamic rupture propagation on a 5 plane in a model consisting of two different elastic half spaces connected via a planar 6 frictional interface governed by regularized Coulomb friction. Therefore, ruptures in this 7 study are purely driven by the presence of a material contrast. Ruptures are nucleated 8 on the fault using a circular symmetric expanding increase of pore-pressure in a limited 9 source region. We show how a wrinkle-like rupture pulse can mature also in the 3D 10 case where we have a mixing of in-plane and anti-plane modes, the instability specific 11 of a bimaterial interface acting only for the in-plane mode. The pulse develops inside a 12 cone-shaped region around the in-plane direction of slip in the softer material. 13

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#### 1. Introduction

Geological faults with a long slip history are likely to bring into contact materials with 17 different elastic properties. Recent geologic mapping and laboratory experiments both 18 suggest that the principal zones of slip are localized along interfaces that separate con-19 siderably different rocks [Dor et al., 2006a, b]. Contrasts of elastic properties across large 20 faults have been imaged by seismic reflection and refraction studies Fuis et al., 2001, 2003; 21 Lutter et al., 2004, body and coda wave tomography [Eberhart-Phillips and Michael, 1998; 22 Magistrale and Sanders, 1995; Shapiro et al., 2005, modeling of geodetic data [Le Pichon 23 et al., 2005] and analysis of head waves that refract along material interfaces in the fault 24 zone structure [McGuire and Ben-Zion, 2005, and references therein]. The range of the 25 velocity contrast across the San Andreas and other large faults is estimated to be about 26 up to 30%, with values of 5-20% often reported. 27

A fault surrounded by identical materials on both sides cannot become unstable when 28 the governing friction law has a single, constant coefficient of friction. However, an inter-29 face separating materials of different elastic properties can become unstable even under 30 this condition [Weertman, 1980]. How much earthquake ruptures are influenced by such 31 material contrasts has been under debate recently [Andrews and Harris, 2005; Harris and 32 Day, 2005; Ben-Zion, 2006a, b]. The model of rupture propagation along a bimaterial 33 interface with a single, constant friction coefficient evidently excludes the weakening be-34 havior of friction during sliding and is unrealistic in this respect. Nevertheless, it is also 35 believed that simple weakening models of friction and their parameters do not have a 36

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<sup>37</sup> clean physical basis and additional physical knowledge has to be developed to come to
<sup>38</sup> physically consistent models [*Rice and Cocco*, 2006].

Destabilization of slip on a bimaterial interface is only present in the 2D in-plane case 39 and it is not present in the 2D anti-plane case. It has been mentioned by *Ben-Zion* 40 and Andrews [1998] that the results of bimaterial driven 2D in-plane rupture simulations 41 might be modified considerably in cases of 3D rupture propagation. Harris and Day [2005] 42 show results of dynamic rupture calculations in 3D with slip-weakening and Kelvin-Voigt 43 viscosity in the bulk. However, (1) it is not yet clear that the Kelvin-Voigt viscosity does 44 regularize ill-posedness, and (2) we wish to isolate the bimaterial instability from that 45 coming from the intrinsic frictional weakening. Therefore the problem of a rupture along 46 a bimaterial interface in 3D still needs examination. 47

#### 2. Ill-posedness, and numerical convergence

#### 2.1. Ill-posedness and Regularization

A frictional interface governed by Coulomb friction in a homogeneous medium with a 48 uniform initial stress along the fault less than the frictional strength never becomes un-49 stable no matter how forcefully an event is initiated in the nucleation zone. As mentioned 50 above, in order to study unstable slip on a bimaterial interface independently from other 51 sources of instability (e.g., slip- or rate-dependent friction) one would therefore ideally 52 wish to use Coulomb friction. However, sliding along a planar bimaterial interface under 53 Coulomb friction is often not well-posed [Adams, 1995; Ranjith and Rice, 2001] in the 54 sense that such problems do not possess any solution, even if not-too-refined numerical 55 simulations might appear to be stable. 56

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It has been shown by *Ranjith and Rice* [2001] that there is a connection between the ill-posedness of the problem and the existence of the generalized Rayleigh wave. A summarizing table is given in *Cochard and Rice* [2000]. The source of ill-posedness has been studied by *Adams* [1995] and is referred to as the Adams instability [*Cochard and Rice*, 2000].

In this study we use an experimentally based constitutive law derived by *Prakash and Clifton* [1993] and *Prakash* [1998] in its simplified form [*Cochard and Rice*, 2000] with its characteristic differential equation for the shear strength  $\tau_1^s$  given by:

$$\dot{\tau}_1^s = -\frac{|V| + V^*}{L} \left[\tau_1^s - f \max\left(0, -\tau_2\right)\right] \tag{1}$$

with slip velocity V, characteristic slip velocity  $V^*$ , characteristic length L, and friction coefficient f letting the shear stress  $\tau_1$  to respond gradually rather than instantaneously to an abrupt change of normal stress  $\tau_2$ . It has been shown by *Ranjith and Rice* [2001] to regularize the previously discussed ill-posedness. Classical slip-weakening or rate- and state-dependent constitutive laws with strength proportional to local normal stress do not provide a regularization [*Cochard and Rice*, 2000].

#### 2.2. Numerical convergence

Even in the well-posed regime it is numerically challenging to resolve a wrinkle-like rupture pulse travelling along a bimaterial interface, because of the above mentioned intrinsic instability of such an interface [*Cochard and Rice*, 2000; *Ben-Zion and Huang*, 2002]. To test the numerical results for their physical plausibility we strive for converging results of grid-refined simulations. In Figure 1 we show converging results achieved with the finite-difference method used in this study (details in section 3) for three levels of

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grid refinement ( $\Delta x = 0.5, 0.25, 0.125$  m). Because of the huge computational expense 74 we cannot accomplish further refinements or larger propagation distances at the same 75 time. The top panel shows slip velocity as a function of time at three different points 76  $(x_{\rm pd} = 100m, 150m, 200m)$  along the in-plane direction on the fault. At about 0.03 s we 77 see a slip pulse that corresponds to what is labeled 'supershear' in figure 2 - it has already 78 totally disappeared at 150 m away from the nucleation. Thus, as in 2D, the main feature 79 is the so called Weertman pulse [e.g., Andrews and Ben-Zion, 1997], appearing in all three 80 stations. For the least refined grid (green curve) we can see numerical noise, especially for 81 the station at 200 m propagation distance. The corresponding features can also be seen 82 in the bottom panel which shows normal stress. We can see that suitable convergence is 83 achieved for  $\Delta x = 0.25$  m. 84

#### 3. Numerical Model and Parameter Setup

#### 3.1. Physical Model and Numerical Implementation

The model consists of two elastic halfspaces of different elastic properties which are cou-85 pled by a frictional interface. We use a standard 4th-order staggered-grid finite-difference 86 technique for solving the elastic wave equations inside the elastic medium which has been 87 described by various authors [e.g., Igel et al., 1995; Graves, 1996]. Frictional sliding is 88 numerically included via the stress-glut method which has been introduced by Andrews 80 [1999]. To keep the simulations clean from artificial reflections caused by the finiteness of 90 the numerical implementation of the model we use perfectly matched layers [Marcinkovich 91 and Olsen, 2003] of appropriate width on all sides of the modelspace. Although Ampuero 92 and Dahlen [2005] show that the elastic constants on a material discontinuity are am-93

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<sup>94</sup> biguous, we found that we could achieve the best duplication of results of *Cochard and* <sup>95</sup> *Rice* [2000] using the harmonic mean for the elastic  $c_{av} = 2c_1c_2/(c_1 + c_2)$  constants and <sup>96</sup> the arithmetic mean for the density  $\rho_{av} = (\rho_1 + \rho_2)/2$ , as has been suggested elsewhere <sup>97</sup> [*Moczo et al.*, 2002; *Wu and Chen*, 2003], for points on the material boundary where the <sup>98</sup> elastic properties are not defined in the staggered scheme.

#### 3.2. Nucleation

We nucleate each event in a circular symmetric way by increasing the fluid pressure in 99 a limited space-time region. We thus prevent a propagation direction that is privileged 100 by the nucleation procedure. Our procedure is similar to the one used by Brietzke and 101 Ben-Zion [2006]. In 3D it is a spatially ring-shaped nucleation pulse with the spatial 102 width a, expanding at the nucleation velocity  $v_{nuc}$  until the maximum radius b, which can 103 be expressed using the geometric variables  $\xi = (|r| - v_{\text{nuc}}t)/a$ , and  $\eta = (|r| + v_{\text{nuc}}t)/b - b$ 104  $(\sqrt{a^2+b^2}/b)$ , with radius  $|r| = \sqrt{y^2+z^2}$ . Within the source, the fluid pressure is given 105 as  $P_{\rm f} = P_0 \left(1 - \xi^2 - \eta^2\right)^2$ , while outside this region it is zero. 106

#### 3.3. Model Parameters

To make our study comparable to other numerical studies [*Cochard and Rice*, 2000; *Ben-Zion and Huang*, 2002; *Ben-Zion and Shi*, 2005] we use parameters that are similar to the ones used in those studies. The range of the investigated parameters is summarized in Table 1.

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#### 4. Results

Snapshots of slip velocity on the fault at three instances in time are shown in Figure 2 111 for  $\Delta x = 0.25$  m. The nucleation initiates rupture phases in the subshear as well as the 112 supershear range. The supershear part of the rupture dies out for the shown parameter 113 set in all directions. In the in-plane direction and some time-dependent angle the subshear 114 rupture phase develops towards a self-amplifying wrinkle-like pulse in the positive direction 115 (direction of slip in the softer material), and it dies out in the opposite (negative) direction. 116 This self-amplifying wrinkle-like pulse travels at about the generalized Rayleigh velocity 117 within a limited region of slip that has the shape of a fraction of a circle. Therefore 118 the propagation velocity of this pulse is constant in all directions where it exists. In 119 the anti-plane direction the pulse symmetrically dies out for all propagation velocities. 120 Although this might look trivial knowing the corresponding 2D cases, where we have 121 the development of a unilateral rupture-pulse in the positive direction of the in-plane 122 case, and dying rupture pulses in the anti-plane case, we want to emphasize that this 123 result cannot be deduced trivially from the separate 2D cases. The snapshots show how 124 a self-amplifying, purely material-contrast driven wrinkle-like rupture pulse can appear 125 also in the 3D case. We found that the allowance of rake rotation in the 3D case does not 126 change the results qualitatively (in a few simulations, not shown here, we found that the 127 self-amplification is stronger for the cases for which we prohibited rake-rotation). 128

In order to investigate how the pulse evolves after some larger propagation distance we performed a 3D simulation with a coarse numerical grid ( $\Delta x = 0.5$  m) allowing us to enlarge the propagation distance to  $x_{pd} = 1$  km. The resulting slip-distribution after

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<sup>132</sup> 475 ms is shown in figure 3. The buldge of slip on the rightmost part shows a very <sup>133</sup> small disturbance (< 4% of the total slip) propagating ahead of the main pulse at the <sup>134</sup> same velocity, and which originates at a propagation distances of about 400 m. The <sup>135</sup> non-circular rupture front in figure 3 also indicates a transition at the edge, the rupture <sup>136</sup> propagation velocity being slightly smaller there.

<sup>137</sup> Our results show that for adequate parameters a wrinkle-mode of rupture that is purely <sup>138</sup> driven by the decrease in normal stress during slip and its positive feedback (amplification) <sup>139</sup> can exist also in the 3D case. However, in 3D it appears that the wrinkle-like rupture <sup>140</sup> pulse degenerates possibly faster than in 2D towards a potentially unrealistic state with <sup>141</sup> a huge slip-acceleration and very small width of the pulse (sharpening).

#### 5. Discussion

The problem of rupture on a bimaterial interface is numerically extremely challenging because of the highly unstable physical mechanism associated with it (compared to, e.g., the classical slip-weakening instability). Using finite-difference calculations we show that a self-sustained pulse can exist under the simplified Prakash-Clifton law also in 3D. Such a pulse travels inside a cone with a time-dependent angle on the fault plane around the positive in-plane direction.

How robust are the results of this study? Based on our experience in 2D and the very limited experience in 3D, we found that there exists ranges of parameters, e.g., regarding the nucleation procedure or a too low level of initial stress, for which self-sustained pulses have not been observed. Nevertheless we are confident that self-sustained propagation

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<sup>152</sup> occurs in at least some significant neighborhood of the present parameters. However, a <sup>153</sup> full parameter space study remains to be done, even in 2D.

Based on very few numerical simulations of rupture propagation along a bimaterial 154 interface governed by slip-weakening friction Andrews and Harris [2005] state that this 155 phenomenon of the wrinkle-like pulse is not important for realistic earthquake rupture. 156 In contrast, it has been under debate recently what is a realistic and physically consistent 157 earthquake model. Additional physical knowledge has to be developed *Rice and Cocco*, 158 2006] to narrow the involved uncertainties and to evaluate physically more consistent, 159 more realistic models of earthquake rupture; off-fault energy dissipation due to plastic 160 strain, visco-elasticity, melt lubrication, thermal pressurization are just examples of what 161 could be taken into account. 162

Our study has the clearly defined conceptual limit of a purely material driven effect. 163 The results of recent studies on bimaterial interfaces bring in complementary insights (e.g., 164 Rudnicki and Rice [2006]; Shi and Ben-Zion [2006]; Brietzke and Ben-Zion [2006]; Rubin 165 and Ampuero [2006]). There are indeed good examples for which the bimaterial mech-166 anism seems to be necessary to properly interpret the observations: Rubin and Gillard 167 [2000] observed asymmetric alongstrike distribution of aftershocks on the San Andreas 168 fault and Dor et al. [2006b] observed asymmetric rock damage across faults of the San 169 Andreas system. 170

In order to address the question of whether or not the wrinkle-like slip pulse (or, more generally, the bimaterial mechanism) is relevant in existing earthquake rupture mechanisms, a much wider range of models and parameter combinations would need to be tested,

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<sup>174</sup> owing to the aforementioned uncertainties in our knowledge of source physics. The strong <sup>175</sup> self-sharpening behavior of the wrinkle-like rupture pulse suggests that, with increasing <sup>176</sup> propagation distance, it degenerates towards unrealistically large slip velocities (as already <sup>177</sup> noted by *Ben-Zion and Huang* [2002] in 2D), and perhaps vanishes as a consequence of <sup>178</sup> pulse thinning, this latter aspect obviously deserving a deeper investigation.

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Table 1.	Range	of simu	lation	parameters
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parameter	value or value range	
numerical method	finite-differences	
grid type	staggered cartesian grid	
grid-spacing $\Delta x$	0.5  m, 0.25  m, 0.125  m	
maximum grid size $n_x \times n_y \times n_z$	$500 \times 3200 \times 3200$	
maximum size of the physical model $x \times y \times z$	$230\times1580\times1580~{\rm m}^3$	
density material 1 $\rho_1$	$3333.3~\rm kg/m^3$	
wave velocities material 1 $v_{s1}$ , $v_{p1}$	3000.0, 5196.2 m/s	
density material 2 $\rho_2$	$2777.7~\rm kg/m^3$	
wave velocities material 2 $v_{\rm s2},v_{\rm p2}$	2500.0, 4330.1 m/s	
size of the fault model $y \times z$	$1580 \times 1580 = 2496400 \text{ m}$	
type of friction	simplified Prakash-Clifton	
friction coefficient $f$	0.75	
characteristic slip velocity $V^*$	$1 \mathrm{m/s}$	
characteristic length $L$	4 mm	
initial shear stress $\tau^{\infty}$	70 MPa	
initial normal stress $\sigma^{\infty}$	100 MPa	
size of nucleation zone $\oslash_{nuc} = 2b$	120 m	
nucleation velocity $v_{\rm nuc}$	$2475~\mathrm{m/s}$	
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Figure 1. Slip velocity, shear-stress, friction, and normal stress at three stations at 100 m, 150 m and 200 m propagation distance along the in-plane direction for the parameter case shown in table 1. For the shown propagation distances, a resolution of  $\Delta x = 0.25$  m provides already good convergence, since the results of the next refinement with  $\Delta x = 0.125$  m basically shows the same solution.

Figure 2. Snapshots of the slip velocity on the fault-plane at three instances in time for the same parameters as in figure 1. The nucleation takes place inside the pink circle  $(\oslash_{nuc} = 2b = 120 \text{ m})$ . The left panel shows slip velocity shortly after the nucleation. Two supershear half-moon shaped rupture phases, and a subshear rupture phase traveling at about the generalized Rayleigh velocity in all directions (see left panel) are initiated. The subshear rupture phase develops towards a self-amplifying wrinkle-like pulse within a cone around the positive in-plane direction, it dies out in the negative in-plane as well as in the anti-plane directions (see middle/right panel).

Figure 3. Distribution of slip after 475 ms simulation time. After passing the nucleation region at  $r_{\rm pd} < 60$  m the pulse travels through a transitional area of relatively stable pulse propagation (compare with figure 2) until it becomes clearly self-sustained.

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Figure 1.

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### Figure 2.

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## Figure 3.

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