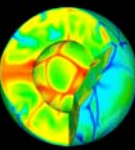




Seismotectonics



Seismic moment and magnitude

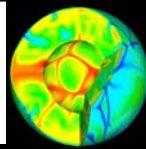
- Fault scarps
- Elastic rebound
- Richter scale
- Energy of earthquakes
- Seismic moment
- Fault area, horizontal slip

Fault plane solutions

- Fault displacement and double couple
- Source radiation pattern
- Beach balls
- Fault plane solutions



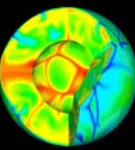
Fault scarps



California



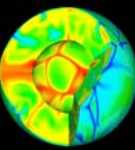
Fault scarps



Grand Canyon



Fault scarps

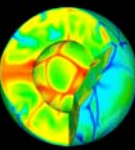


California





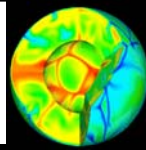
Fault scarps



California



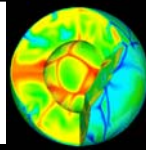
Fault scarps



Taiwan



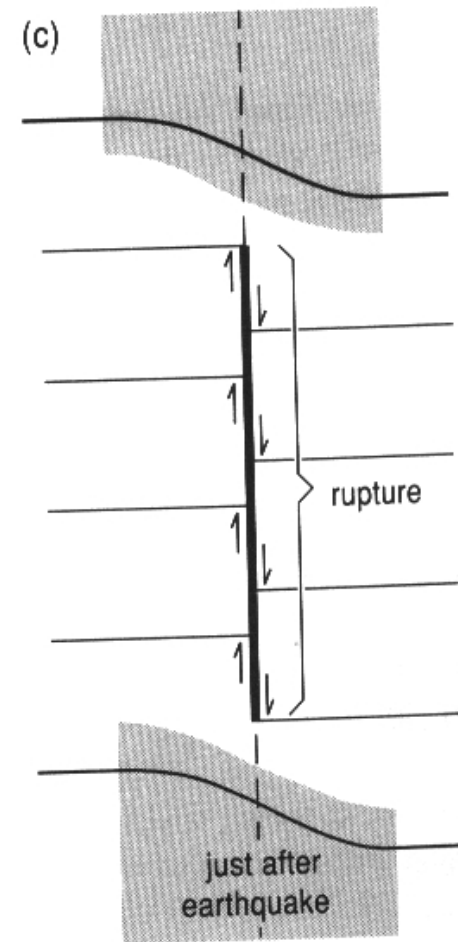
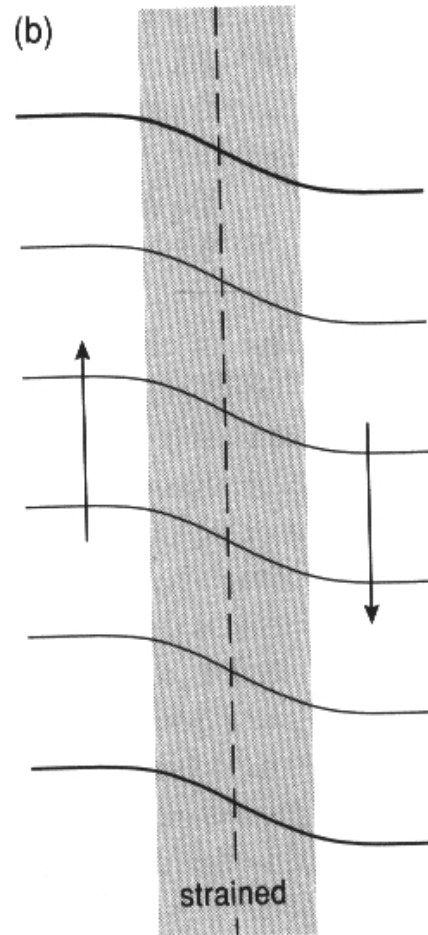
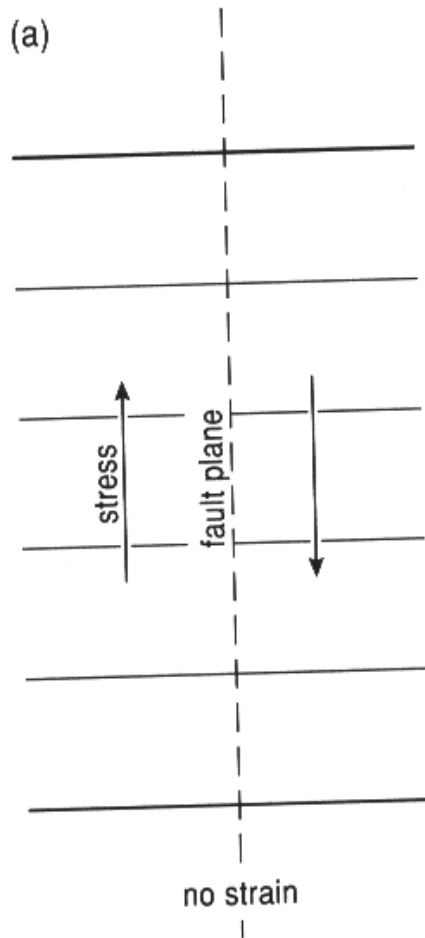
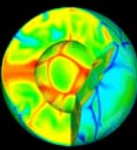
Fault scarps



Taiwan

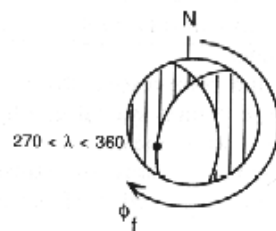
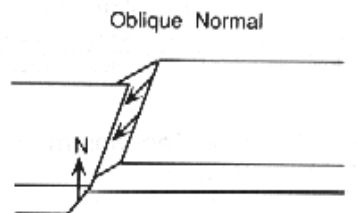
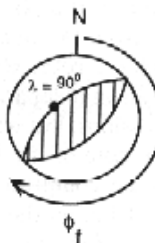
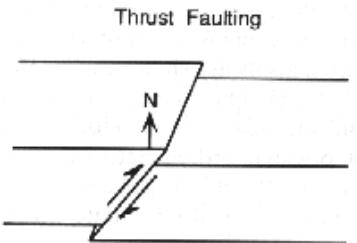
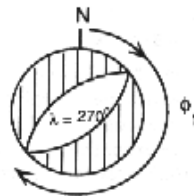
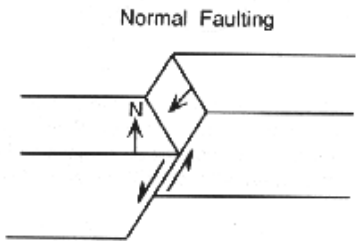
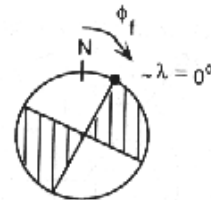
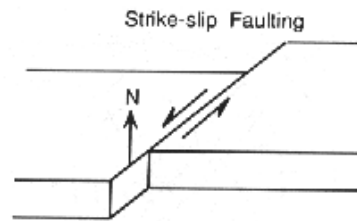
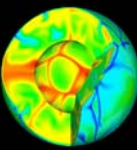


Elastic rebound





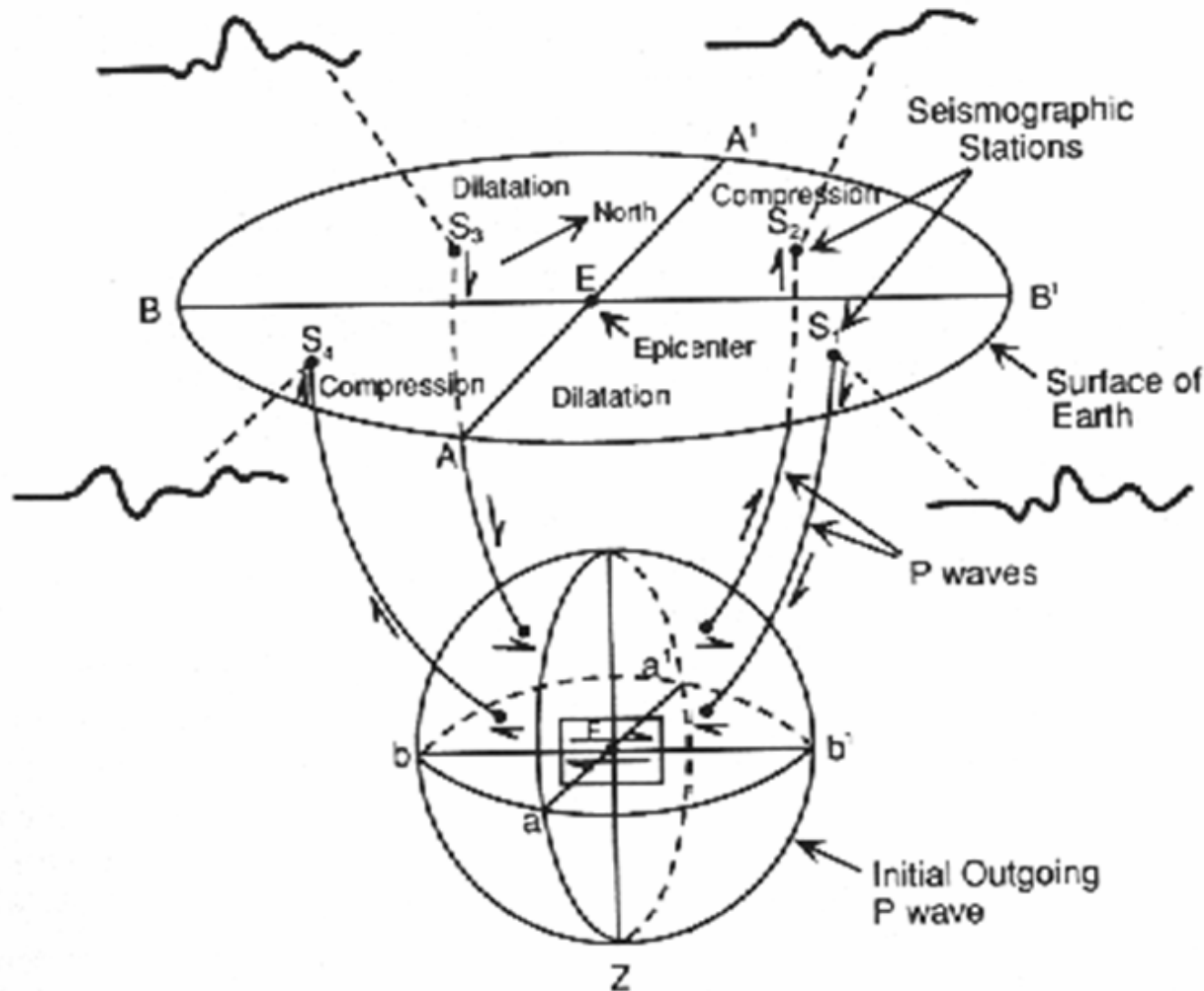
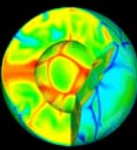
Fault types



Basis fault types and their appearance in the focal mechanisms. Dark regions indicate compressional P-wave motion.



Radiation from shear dislocation

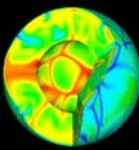


First motion of P waves at seismometers in various directions.

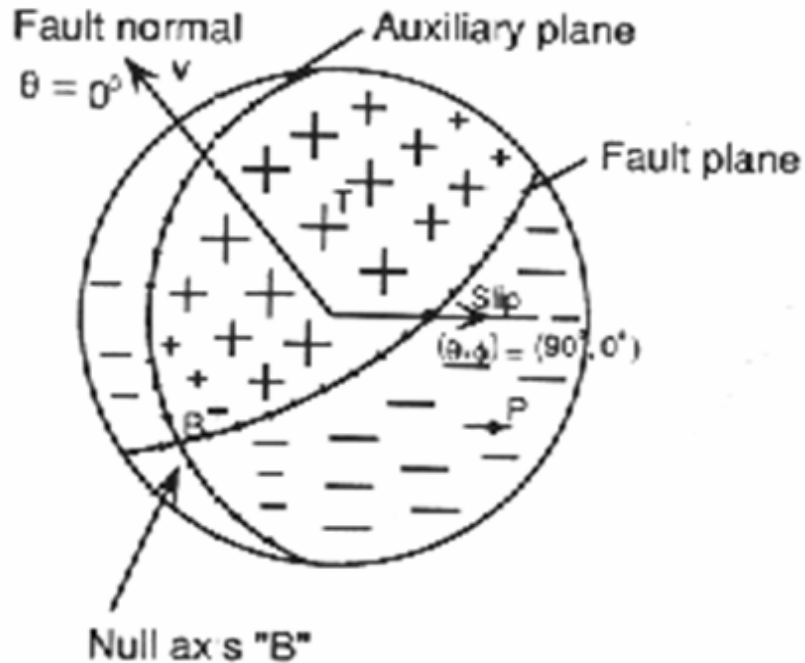
The polarities of the observed motion is used to determine the point source characteristics.



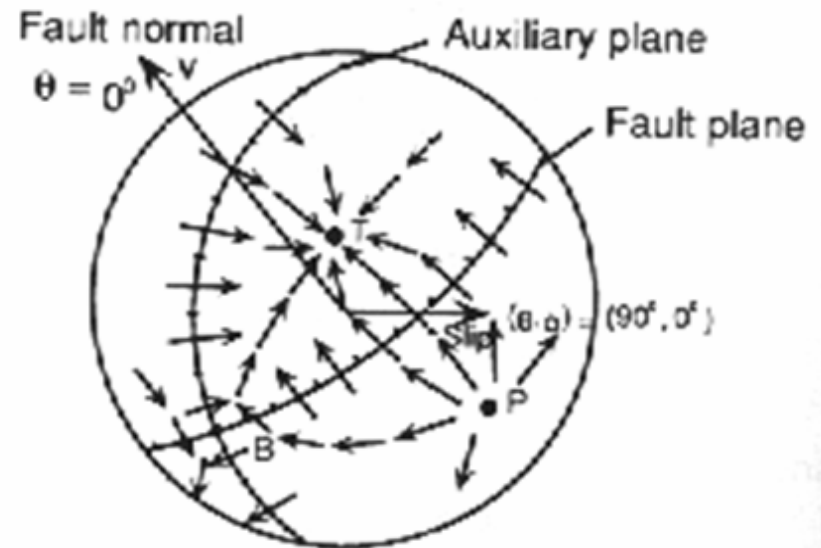
Focal Mechanisms



Focal mechanism for an oblique-slip event.



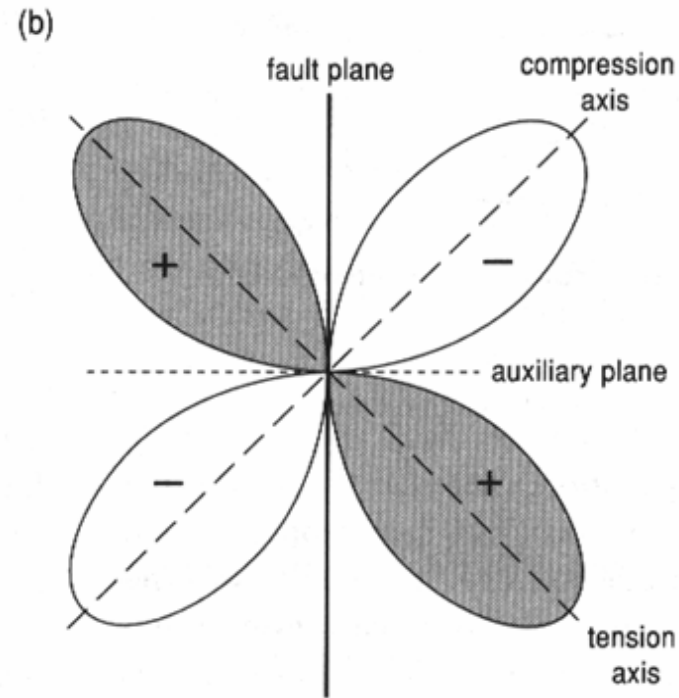
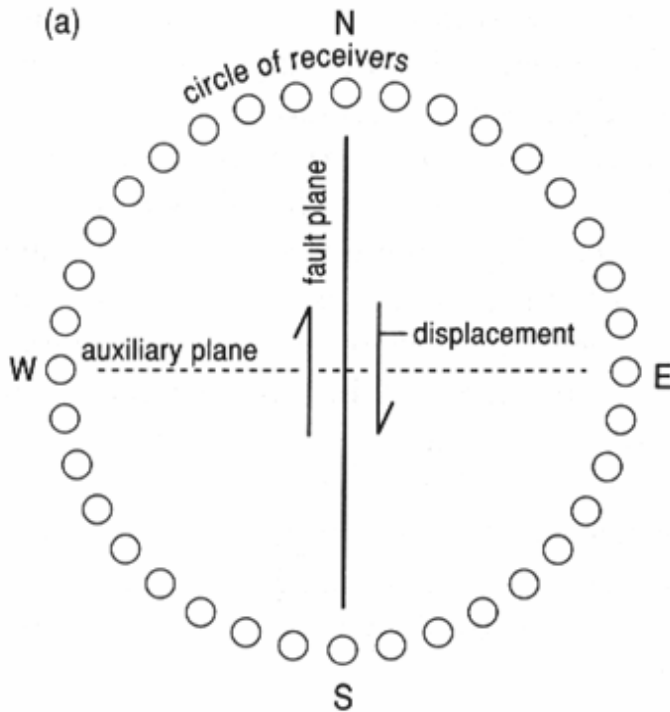
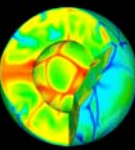
P-wave polarities and relative amplitudes



S-wave polarizations and amplitudes



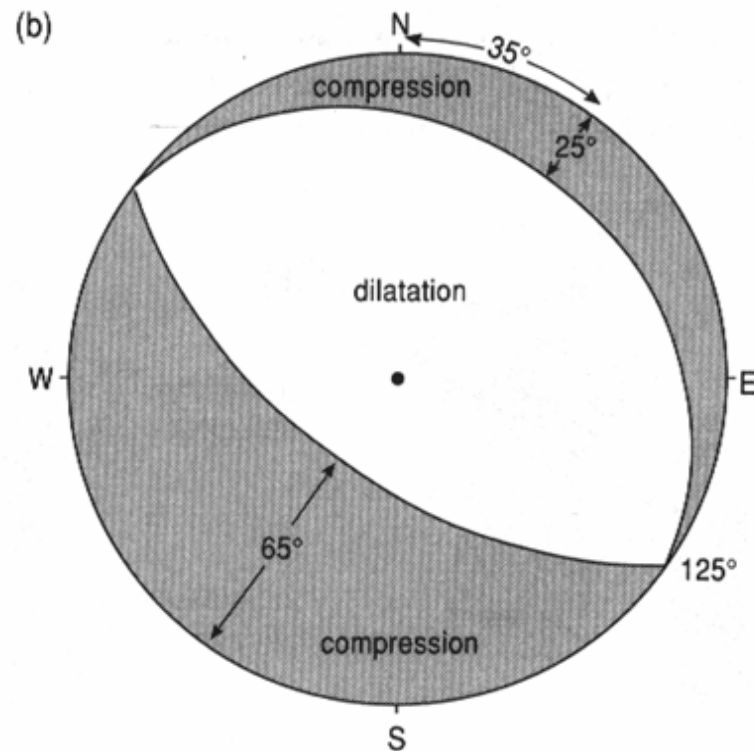
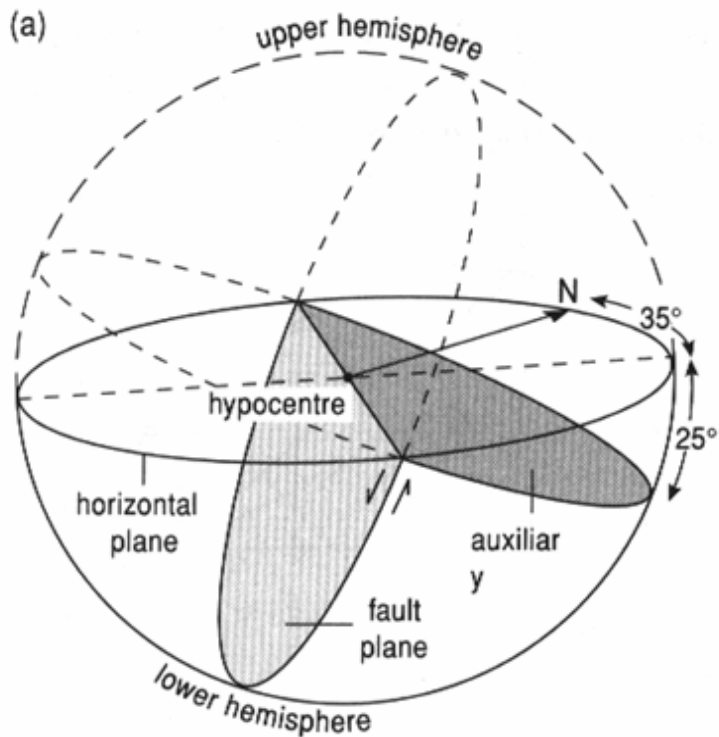
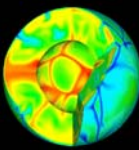
Seismic sources



The basic physical model for a source is two fault planes slipping in opposite directions



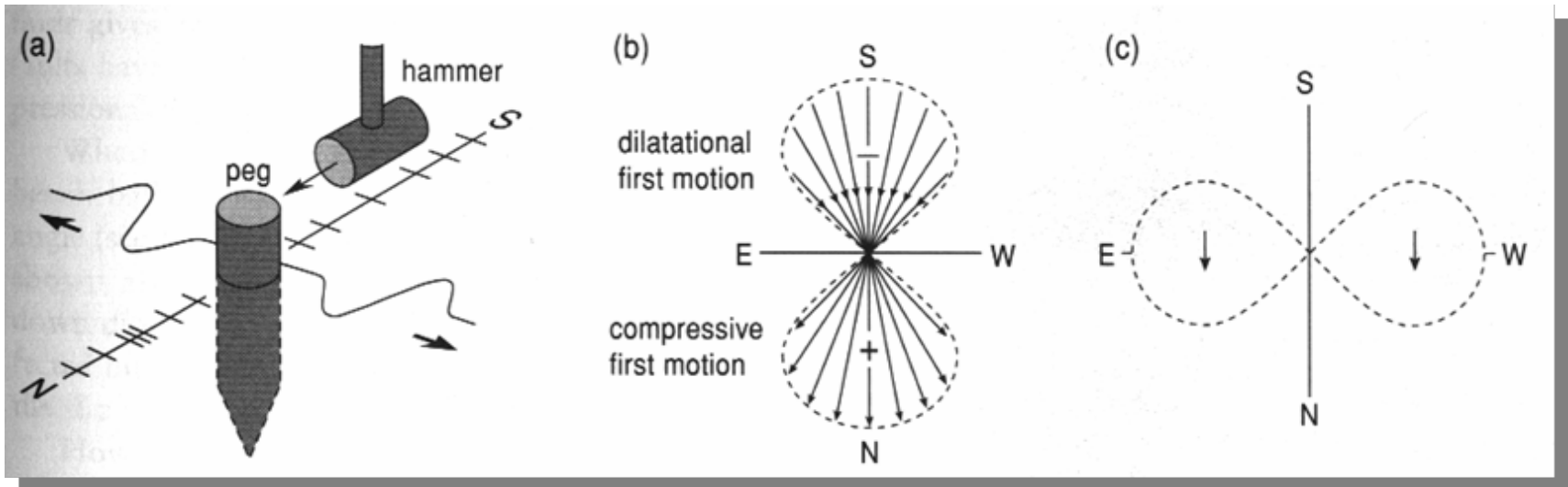
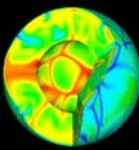
Seismic sources



Our goal: find the **fault plane** and the **slip direction**



Seismic sources



The radiation from seismic sources is in general strongly direction-dependent



Radiation from a point source

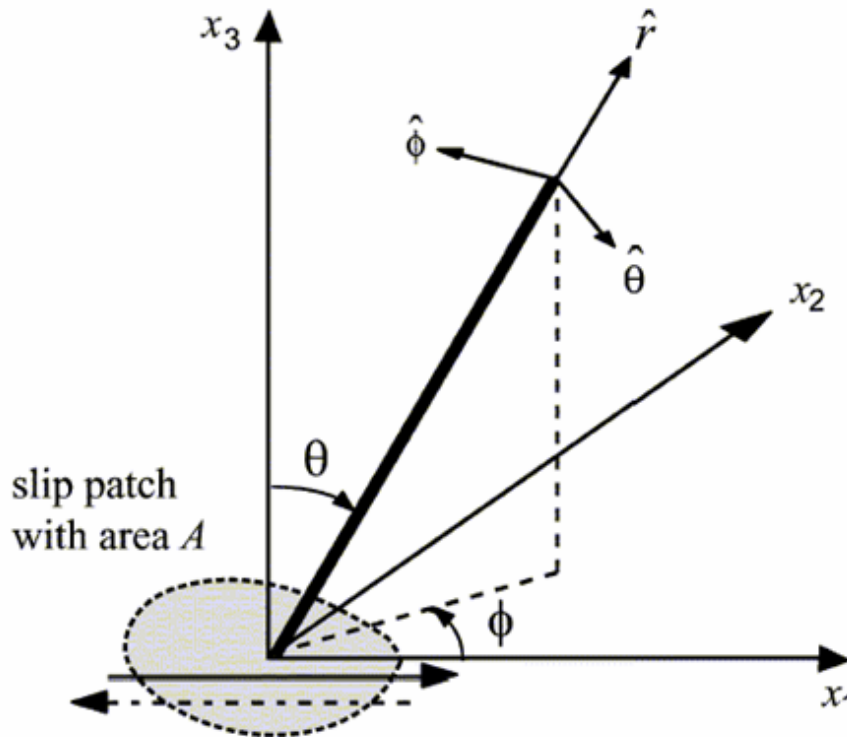
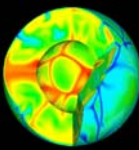
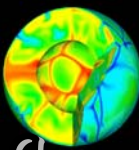


FIGURE 5 Cartesian and polar coordinate systems for analysis of radiation by a slip patch with area A and average slip $\langle \Delta u(t) \rangle$.

Geometry we use to express the seismic wavefield radiated by point double-couple source with area A and slip Δu

Here the fault plane is the x_1x_2 -plane and the slip is in x_1 -direction.

Which stress components are affected?



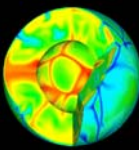
Radiation from a point source

$$\begin{aligned}
 u(x, t) = & \frac{1}{4\pi\rho} A^N \frac{1}{r^4} \int_{r/v_P}^{r/v_S} \tau M_0(t - \tau) d\tau \\
 & + \frac{1}{4\pi\rho v_P^2} A^{IP} \frac{1}{r^2} M_0(t - r/v_P) \\
 & + \frac{1}{4\pi\rho v_S^2} A^{IS} \frac{1}{r^2} M_0(t - r/v_S) \\
 & + \frac{1}{4\pi\rho v_P^3} A^{FP} \frac{1}{r} \dot{M}_0(t - r/v_P) \\
 & + \frac{1}{4\pi\rho v_S^3} A^{FS} \frac{1}{r} \dot{M}_0(t - r/v_S).
 \end{aligned}$$

... one of the most important results of seismology!
... Let's have a closer look ...

u ground displacement as a function of space and time
 ρ density
 r distance from source
 V_s shear velocity
 V_P P-velocity
 N near field
 IP/S intermediate field
 FP/S far field
 M_0 seismic moment

$$\begin{aligned}
 A^N &= 9 \sin 2\theta \cos \phi \hat{r} - 6(\cos 2\theta \cos \phi \hat{\theta} - \cos \theta \sin \phi \hat{\phi}), \\
 A^{IP} &= 4 \sin 2\theta \cos \phi \hat{r} - 2(\cos 2\theta \cos \phi \hat{\theta} - \cos \theta \sin \phi \hat{\phi}), \\
 A^{IS} &= -3 \sin 2\theta \cos \phi \hat{r} + 3(\cos 2\theta \cos \phi \hat{\theta} - \cos \theta \sin \phi \hat{\phi}), \\
 A^{FP} &= \sin 2\theta \cos \phi \hat{r}, \\
 A^{FS} &= \cos 2\theta \cos \phi \hat{\theta} - \cos \theta \sin \phi \hat{\phi},
 \end{aligned}$$



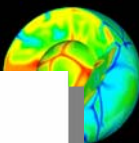
Radiation from a point source

$$\begin{aligned} u(x, t) = & \frac{1}{4\pi\rho} A^N \frac{1}{r^4} \int_{r/v_P}^{r/v_S} \tau M_0(t - \tau) d\tau \\ & + \frac{1}{4\pi\rho v_P^2} A^{IP} \frac{1}{r^2} M_0(t - r/v_P) \\ & + \frac{1}{4\pi\rho v_S^2} A^{IS} \frac{1}{r^2} M_0(t - r/v_S) \\ & + \frac{1}{4\pi\rho v_P^3} A^{FP} \frac{1}{r} \dot{M}_0(t - r/v_P) \\ & + \frac{1}{4\pi\rho v_S^3} A^{FS} \frac{1}{r} \dot{M}_0(t - r/v_S). \end{aligned}$$

Near field
term contains
the static
deformation

Intermediate
terms

Far field
terms: the
main ingredient
for source
inversion, ray
theory, etc.



Radiation pattern

$$A^N = 9 \sin 2\theta \cos \phi \hat{r} - 6(\cos 2\theta \cos \phi \hat{\theta} - \cos \theta \sin \phi \hat{\phi}),$$

$$A^{IP} = 4 \sin 2\theta \cos \phi \hat{r} - 2(\cos 2\theta \cos \phi \hat{\theta} - \cos \theta \sin \phi \hat{\phi}),$$

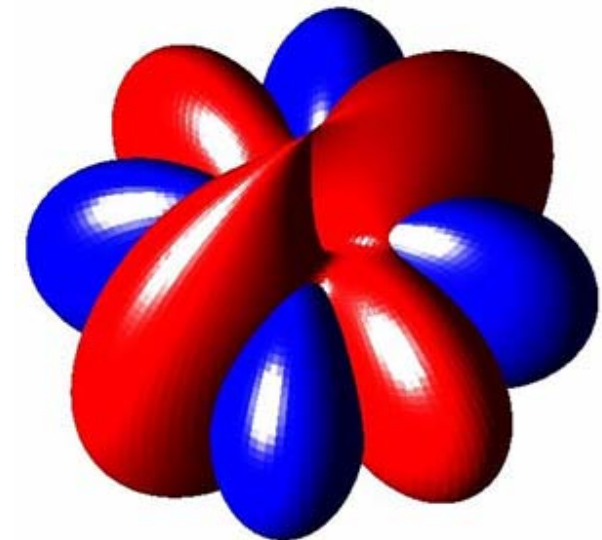
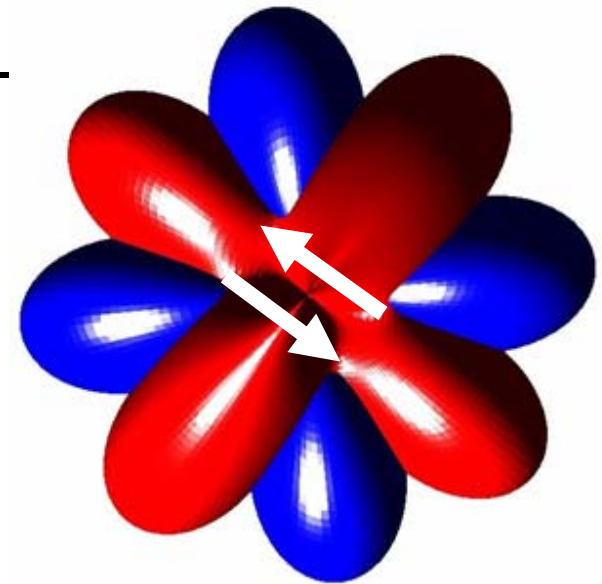
$$A^{IS} = -3 \sin 2\theta \cos \phi \hat{r} + 3(\cos 2\theta \cos \phi \hat{\theta} - \cos \theta \sin \phi \hat{\phi}),$$

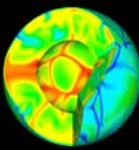
$$A^{FP} = \sin 2\theta \cos \phi \hat{r},$$

$$A^{FS} = \cos 2\theta \cos \phi \hat{\theta} - \cos \theta \sin \phi \hat{\phi},$$

Far field P - blue

Far field S - red





Seismic moment

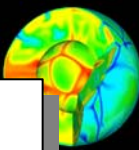
$$M_0$$

$$M_0 = \mu \langle \Delta u(t) \rangle A$$

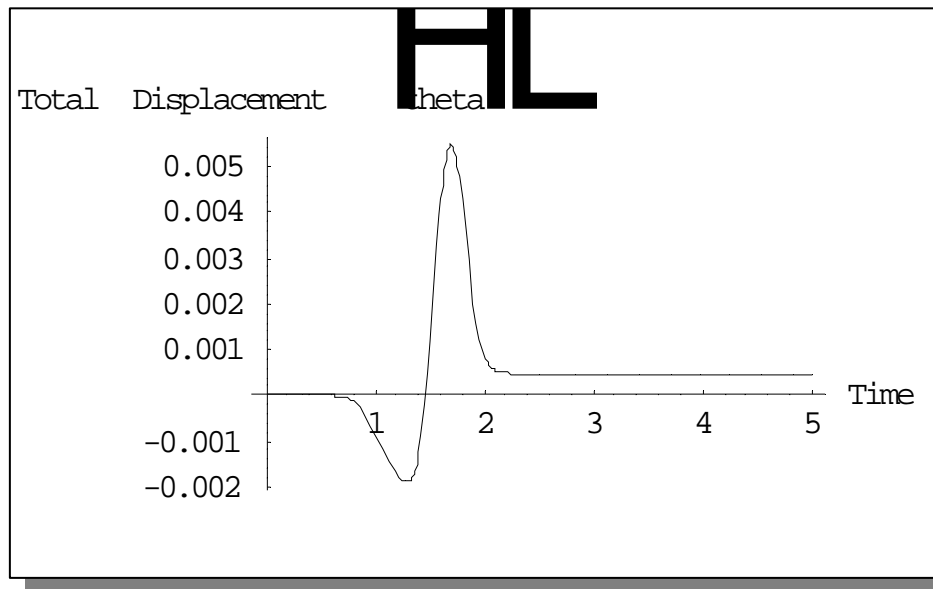
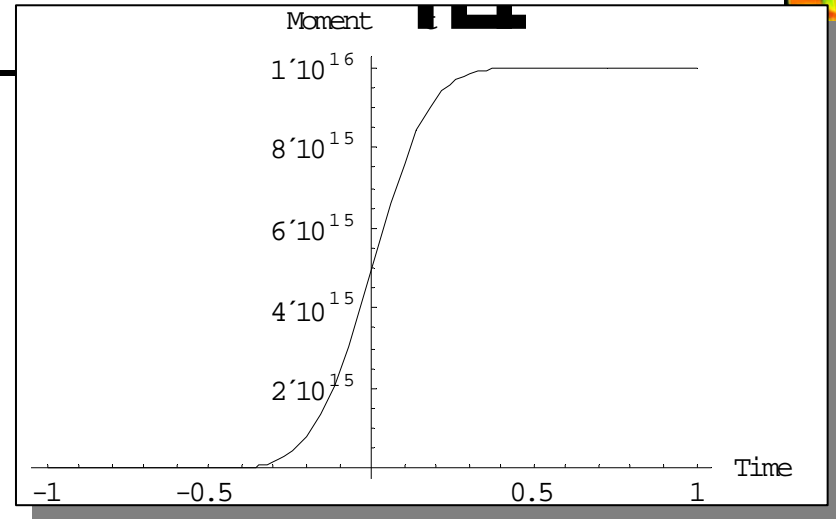
M_0	seismic moment
μ	rigidity
$\langle \Delta u(t) \rangle$	average slip
A	fault area

$$\begin{aligned} u(x, t) = & \frac{1}{4\pi\rho} A^N \frac{1}{r^4} \int_{r/v_P}^{r/v_S} \tau M_0(t - \tau) d\tau \\ & + \frac{1}{4\pi\rho v_P^2} A^{IP} \frac{1}{r^2} M_0(t - r/v_P) \\ & + \frac{1}{4\pi\rho v_S^2} A^{IS} \frac{1}{r^2} M_0(t - r/v_S) \\ & + \frac{1}{4\pi\rho v_P^3} A^{FP} \frac{1}{r} \dot{M}_0(t - r/v_P) \\ & + \frac{1}{4\pi\rho v_S^3} A^{FS} \frac{1}{r} \dot{M}_0(t - r/v_S). \end{aligned}$$

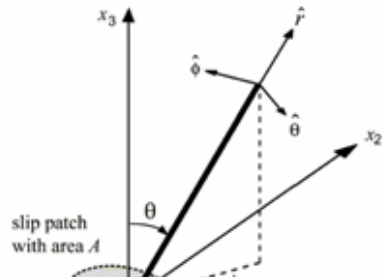
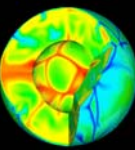
Note that the far-field displacement is proportional to the **moment rate**!



Seismograms

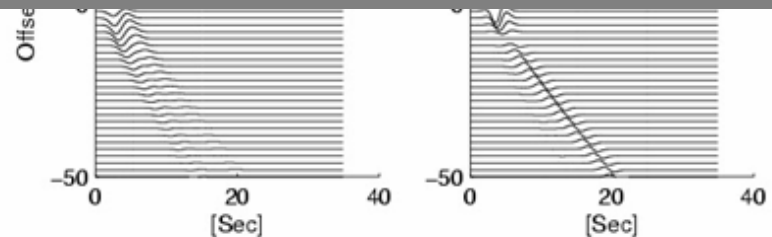
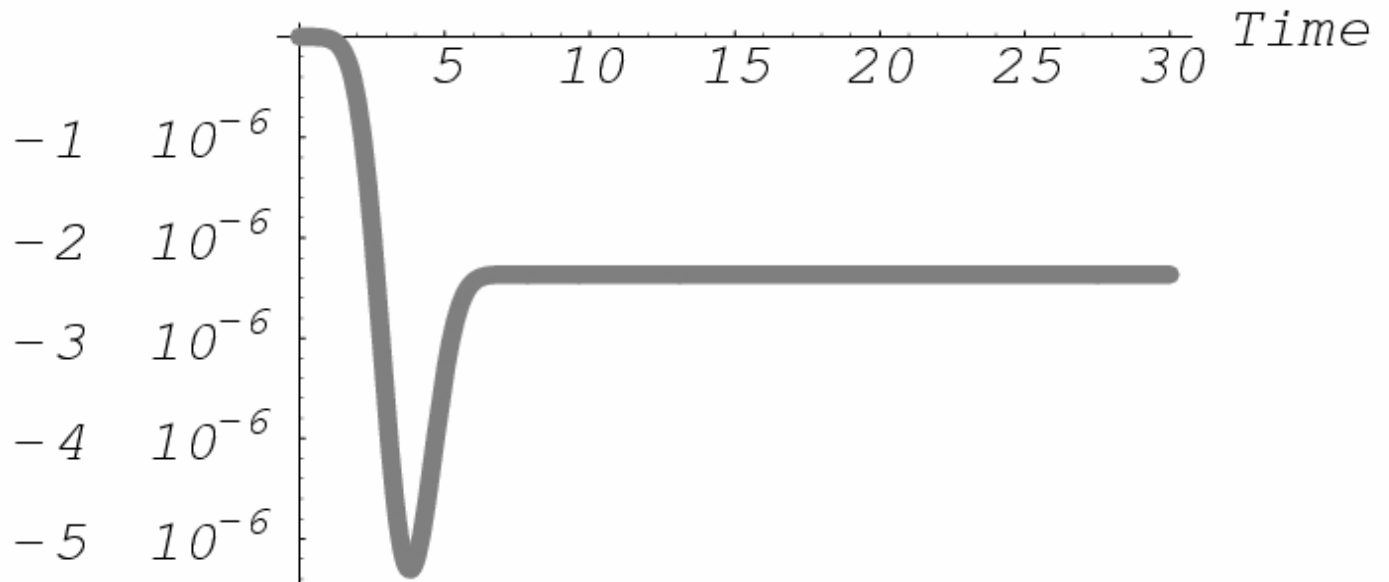


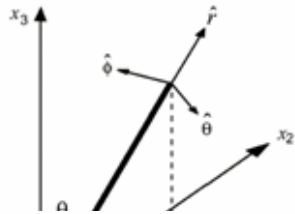
Horizontal displacement 5km away from the source



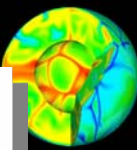
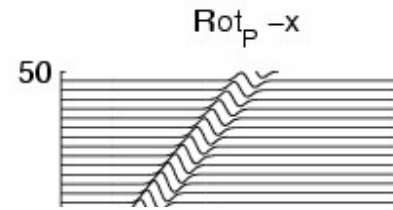
Velocity seismograms M6.5 point source

Total Displacement

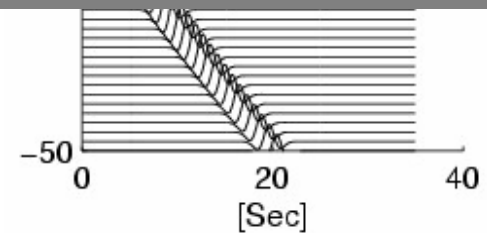
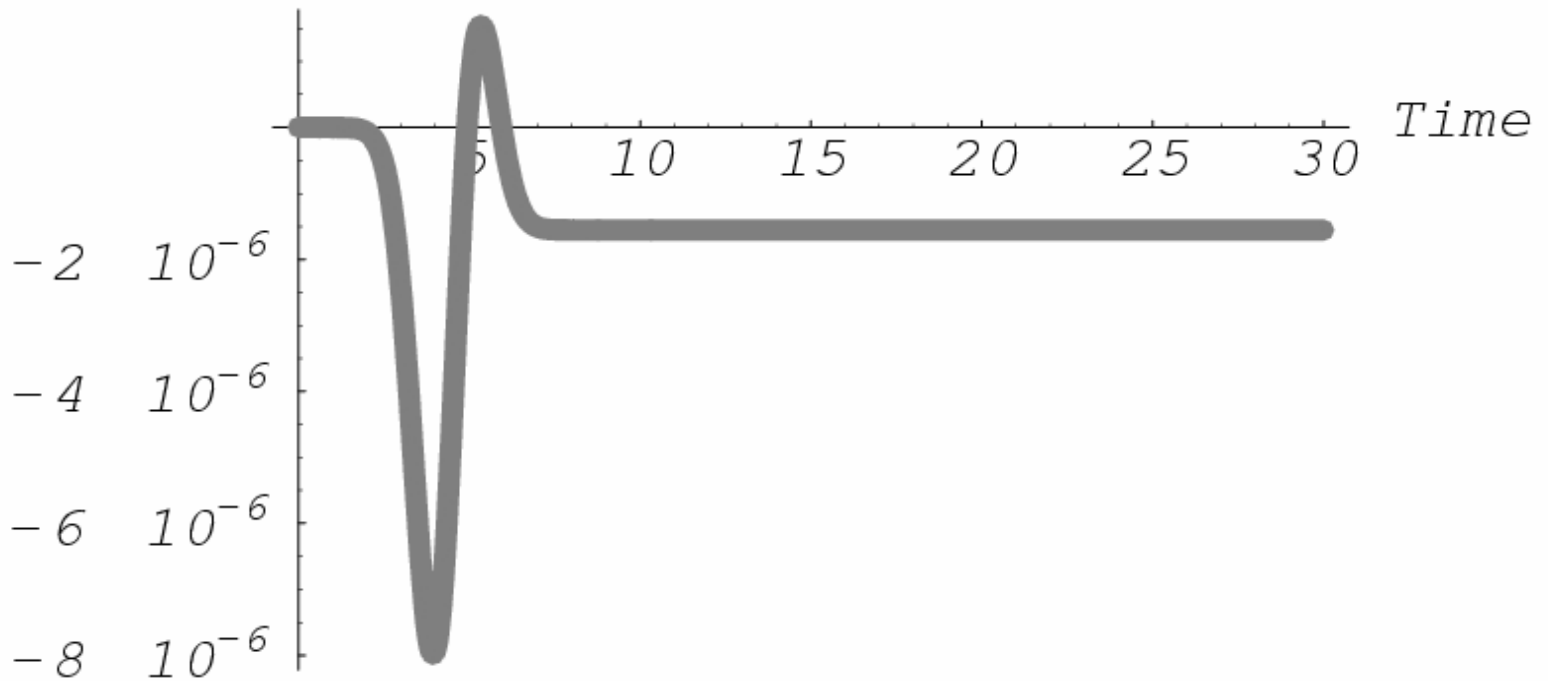




Rotational seismograms M6.5 point

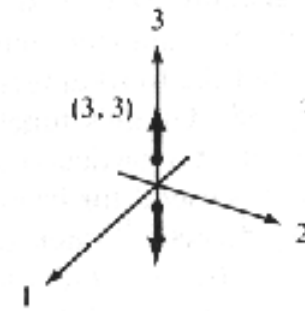
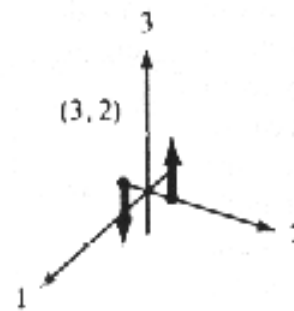
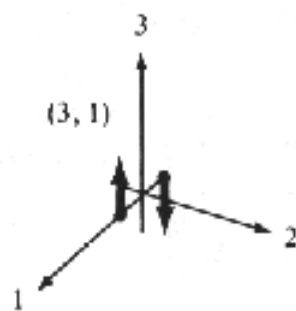
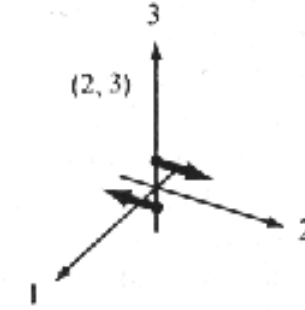
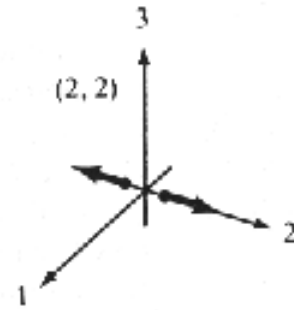
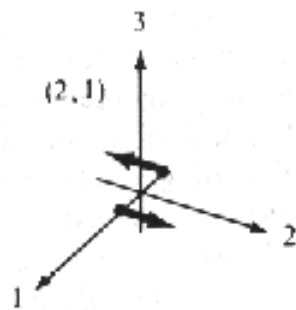
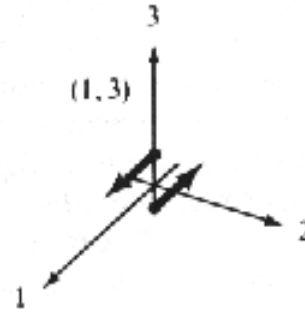
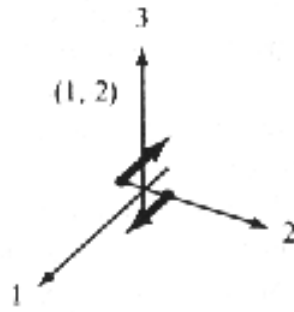
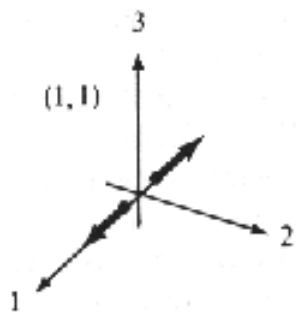
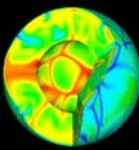


Total Curl HThetaL





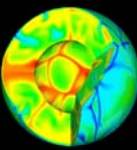
Moment tensor components



Point sources can be described by the seismic moment tensor M . The elements of M have clear physical meaning as forces acting on particular planes.



Beachballs and moment tensor



Moment Tensor	Beachball	Moment Tensor	Beachball
$\frac{1}{\sqrt{3}} \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix}$		$-\frac{1}{\sqrt{3}} \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix}$	
$-\frac{1}{\sqrt{2}} \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}$		$\frac{1}{\sqrt{2}} \begin{pmatrix} 0 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & -1 \end{pmatrix}$	
$\frac{1}{\sqrt{2}} \begin{pmatrix} 0 & 1 & 0 \\ 1 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}$		$\frac{1}{\sqrt{2}} \begin{pmatrix} 0 & 0 & 1 \\ 0 & 0 & 0 \\ 1 & 0 & 0 \end{pmatrix}$	
$\frac{1}{\sqrt{2}} \begin{pmatrix} 1 & 0 & 0 \\ 0 & -1 & 0 \\ 0 & 0 & 0 \end{pmatrix}$		$\frac{1}{\sqrt{2}} \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & -1 \end{pmatrix}$	
$\frac{1}{\sqrt{6}} \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & -2 \end{pmatrix}$		$\frac{1}{\sqrt{6}} \begin{pmatrix} 1 & 0 & 0 \\ 0 & -2 & 0 \\ 0 & 0 & 1 \end{pmatrix}$	
$\frac{1}{\sqrt{6}} \begin{pmatrix} -2 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix}$		$-\frac{1}{\sqrt{6}} \begin{pmatrix} -2 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix}$	

explosion - implosion

vertical strike slip fault

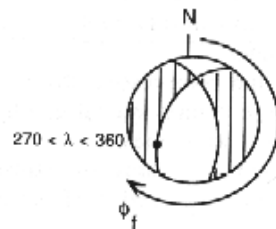
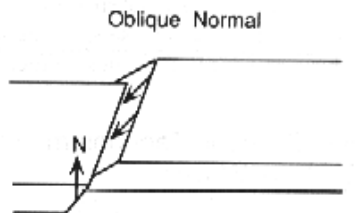
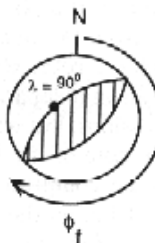
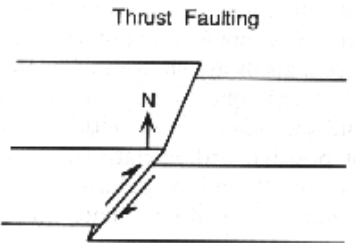
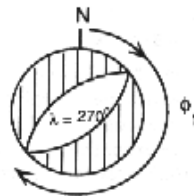
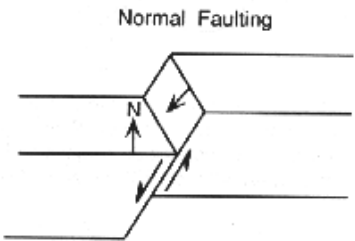
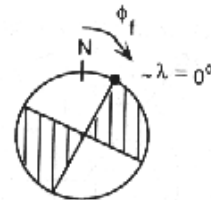
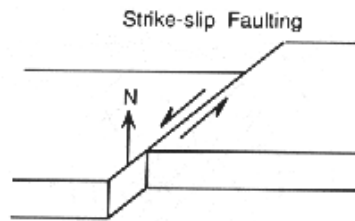
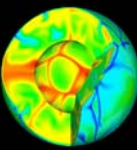
vertical dip slip fault

45° dip thrust fault

compensated linear vector
dipoles



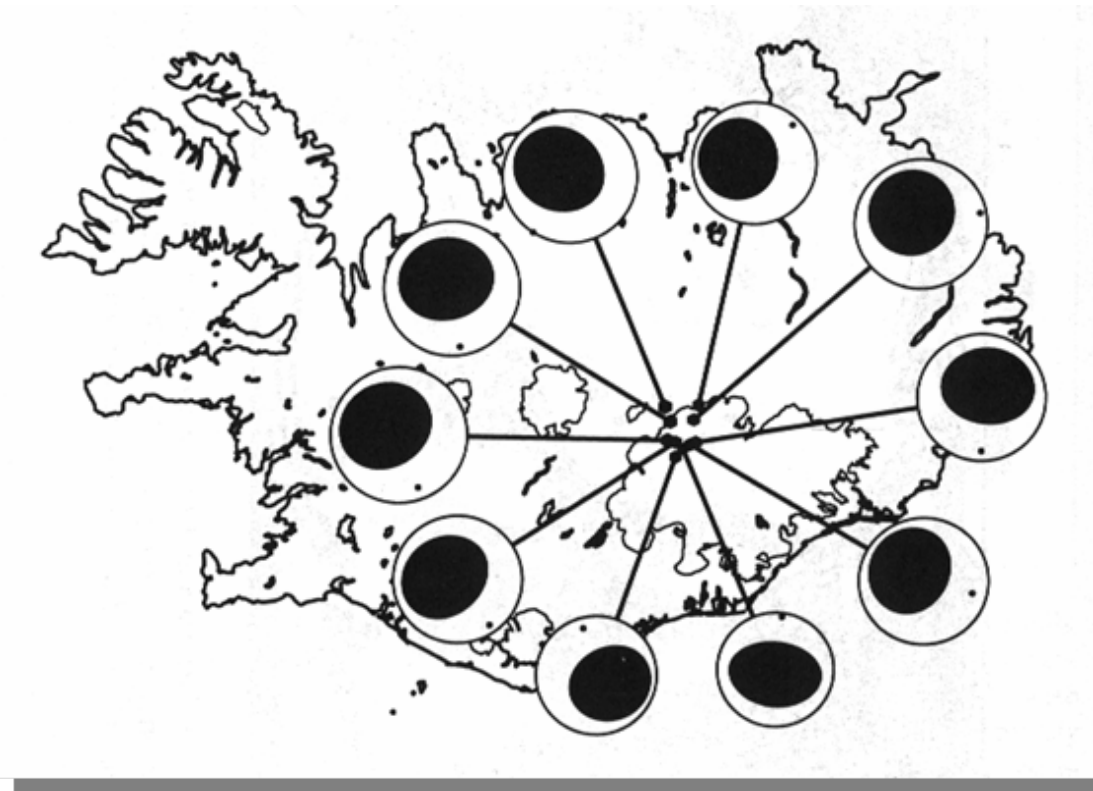
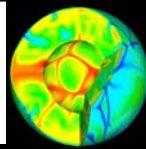
Fault types



Basis fault types and their appearance in the focal mechanisms. Dark regions indicate compressional P-wave motion.



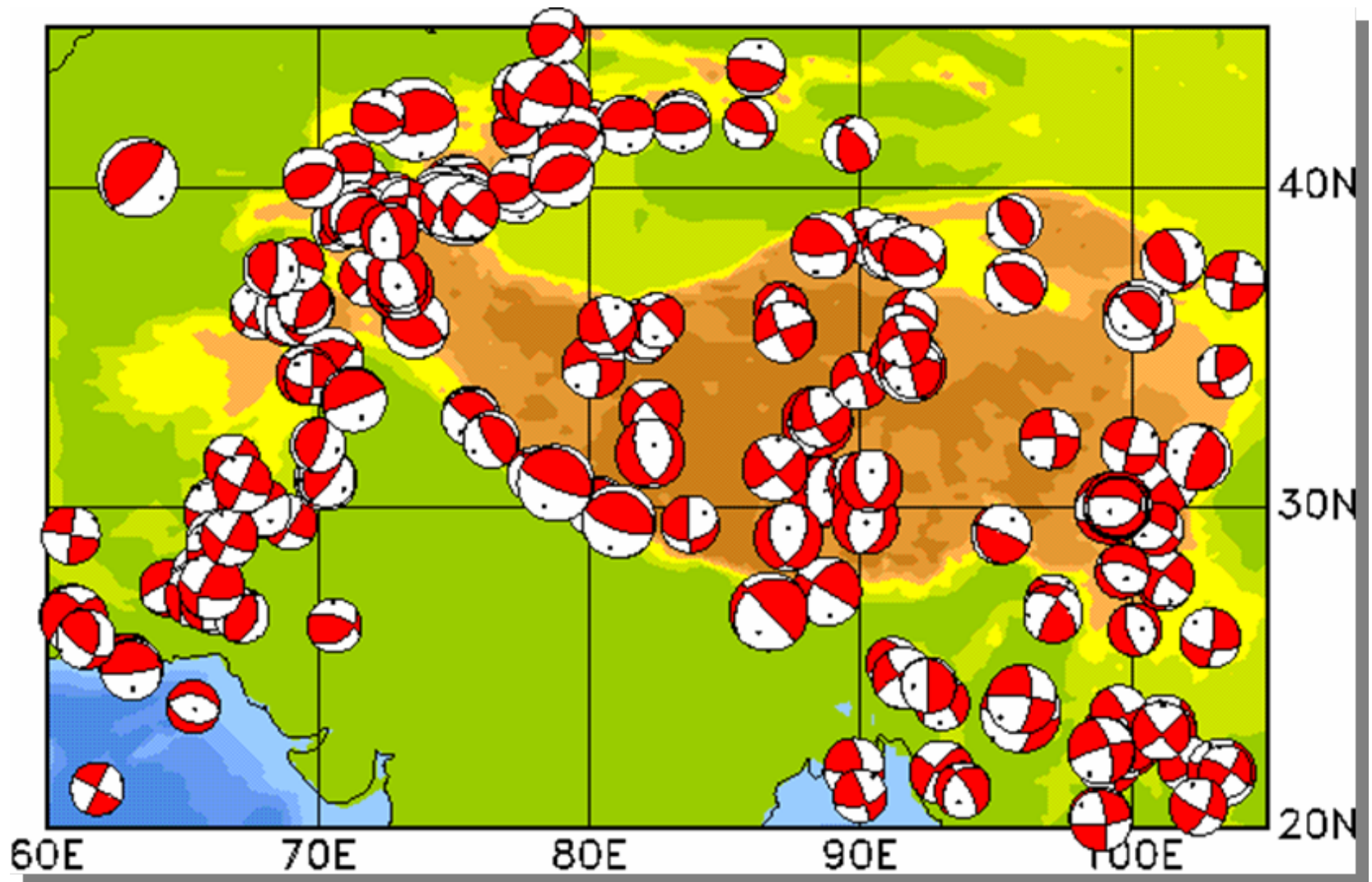
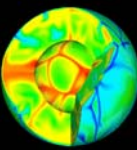
Beachballs - Iceland



Fried eggs: simultaneous vertical extension and horizontal compression

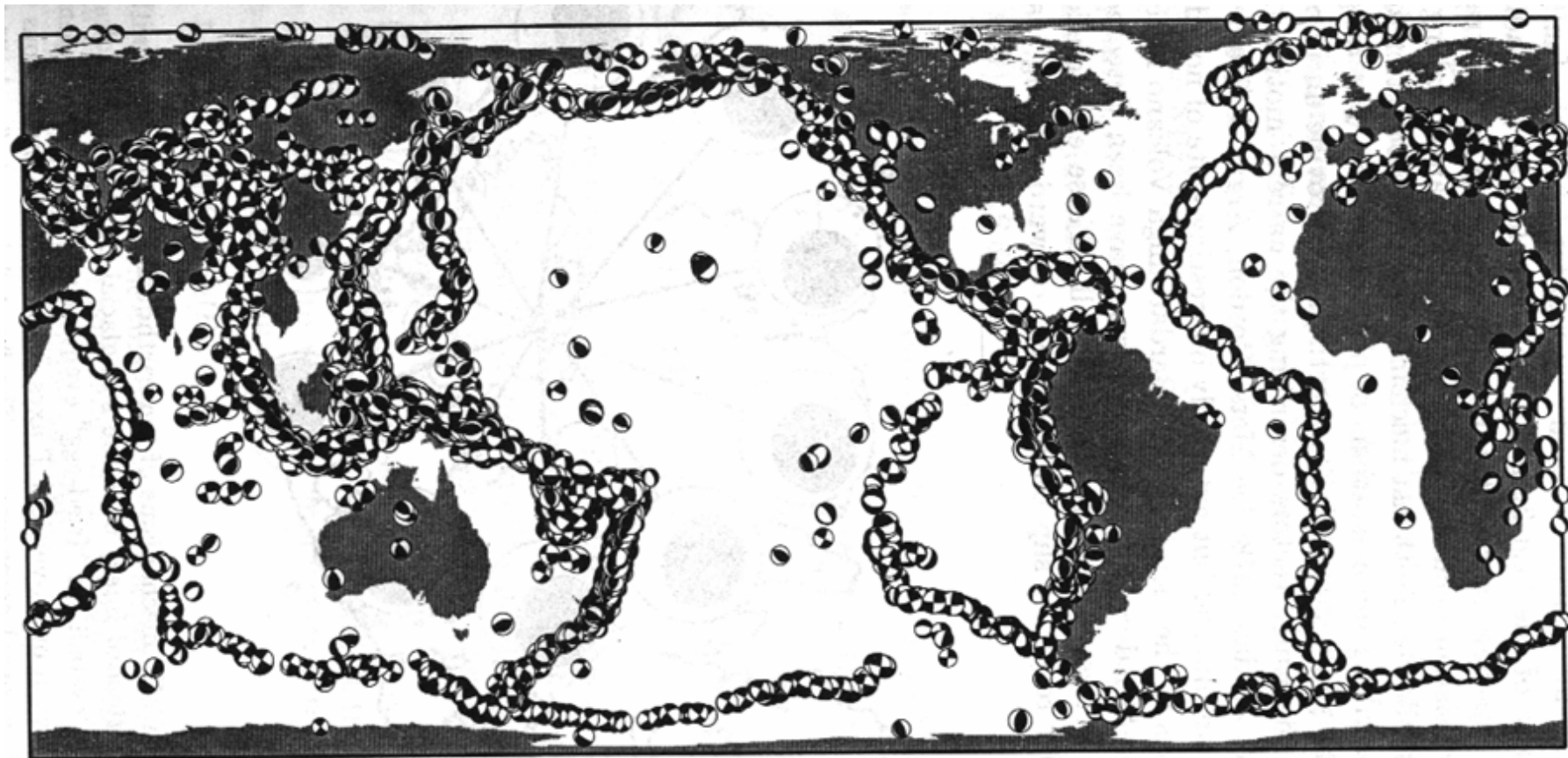
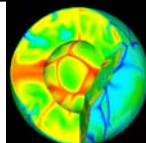


Beachballs - Himalaya



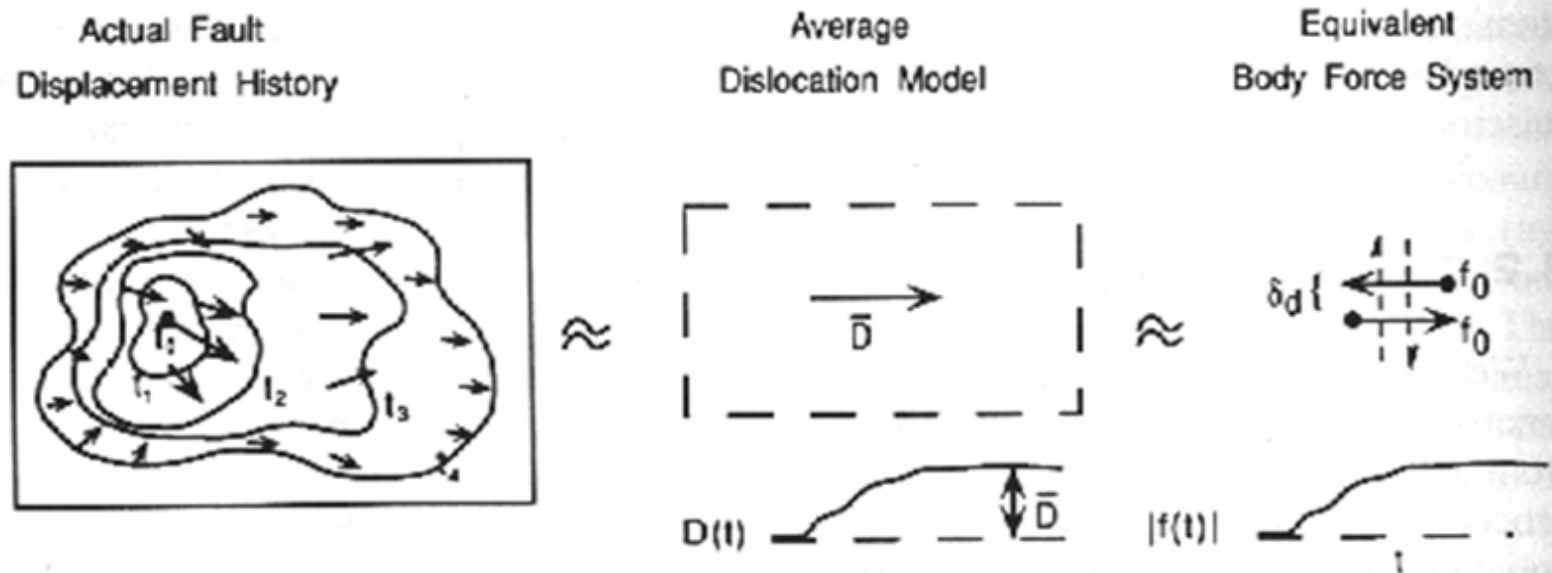
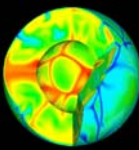


Beachballs - global





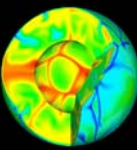
Equivalent Forces: concepts



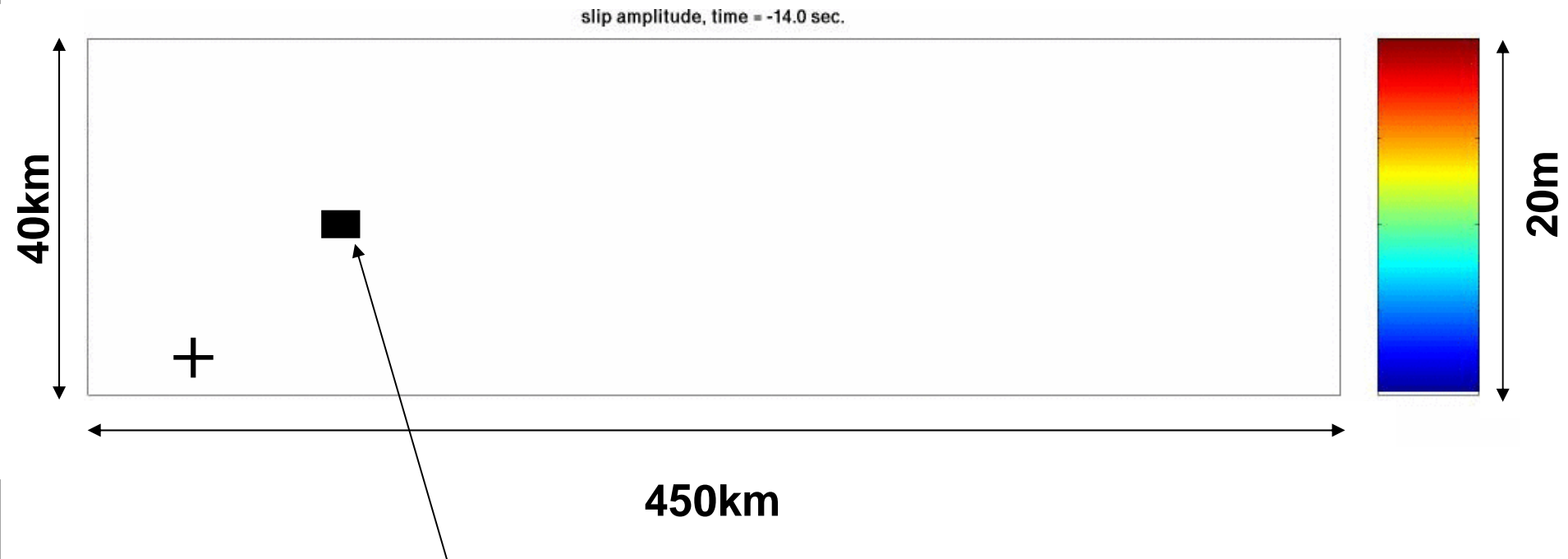
The actual slip process is described by superposition of equivalent forces acting in space and time.



26 Dec 2004 01:58:53MET



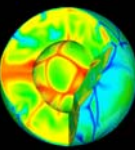
Der Bruchvorgang



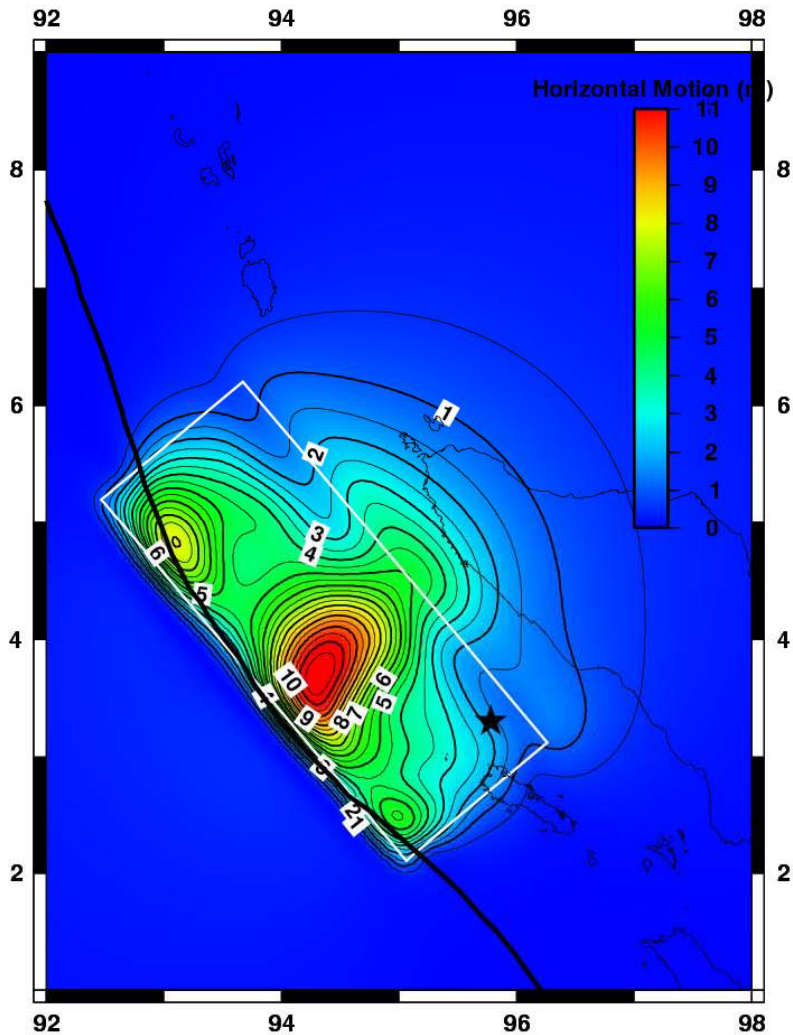
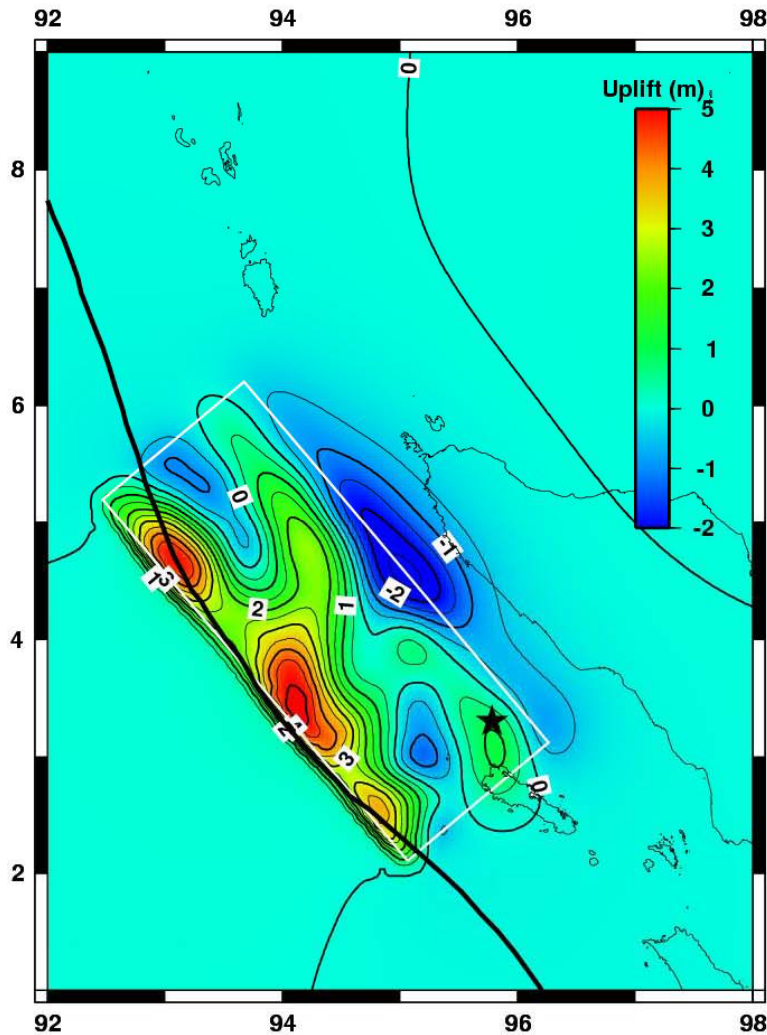
Größe zu erwartende Bruchfläche in Deutschland



26 Dec 2004 02:02:00MET

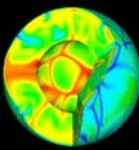


Verschiebung am Meeresboden





Simulation of rotational motions in 3D



(heterogeneous) media: finite faults

Mw: 6.5

L: 23 km

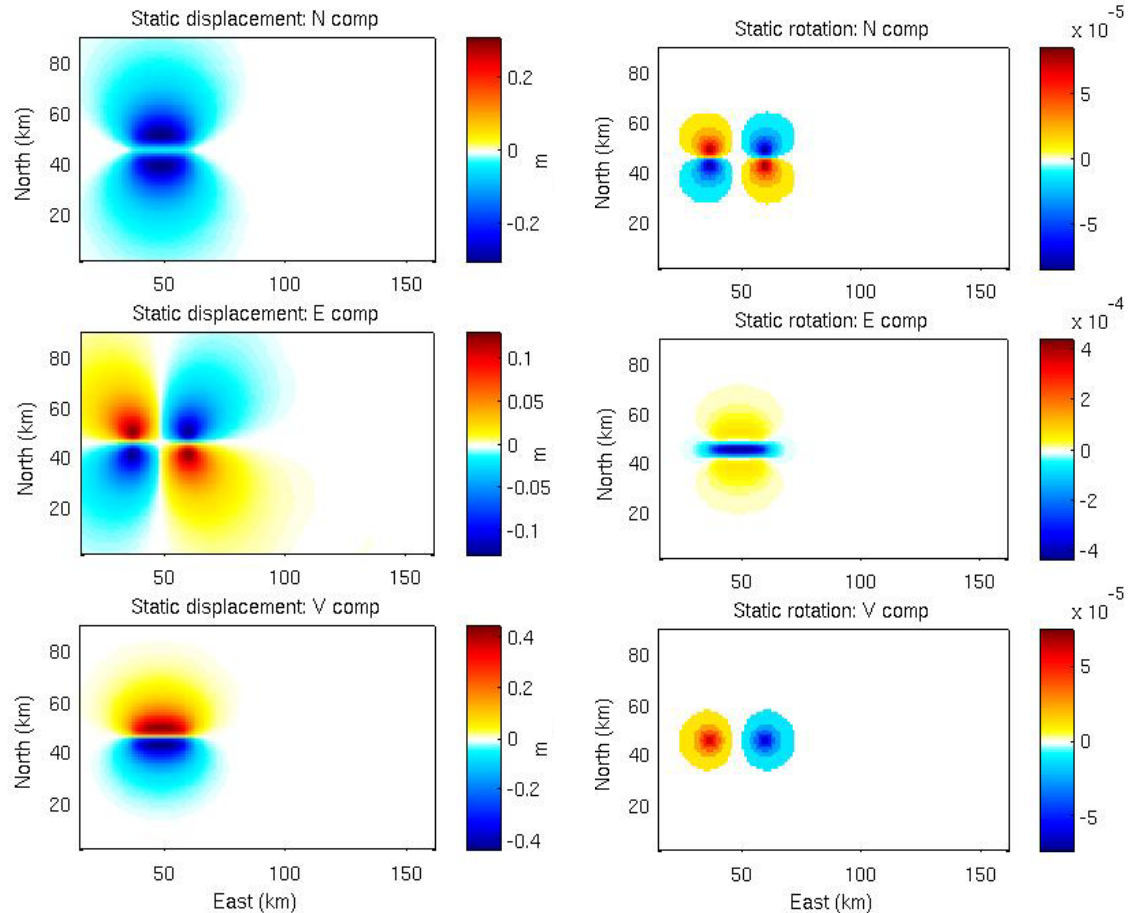
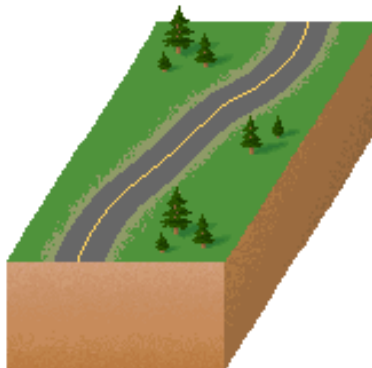
W: 14 km

L_e : 1x1 km

N: 22x14

Haskell rupture model

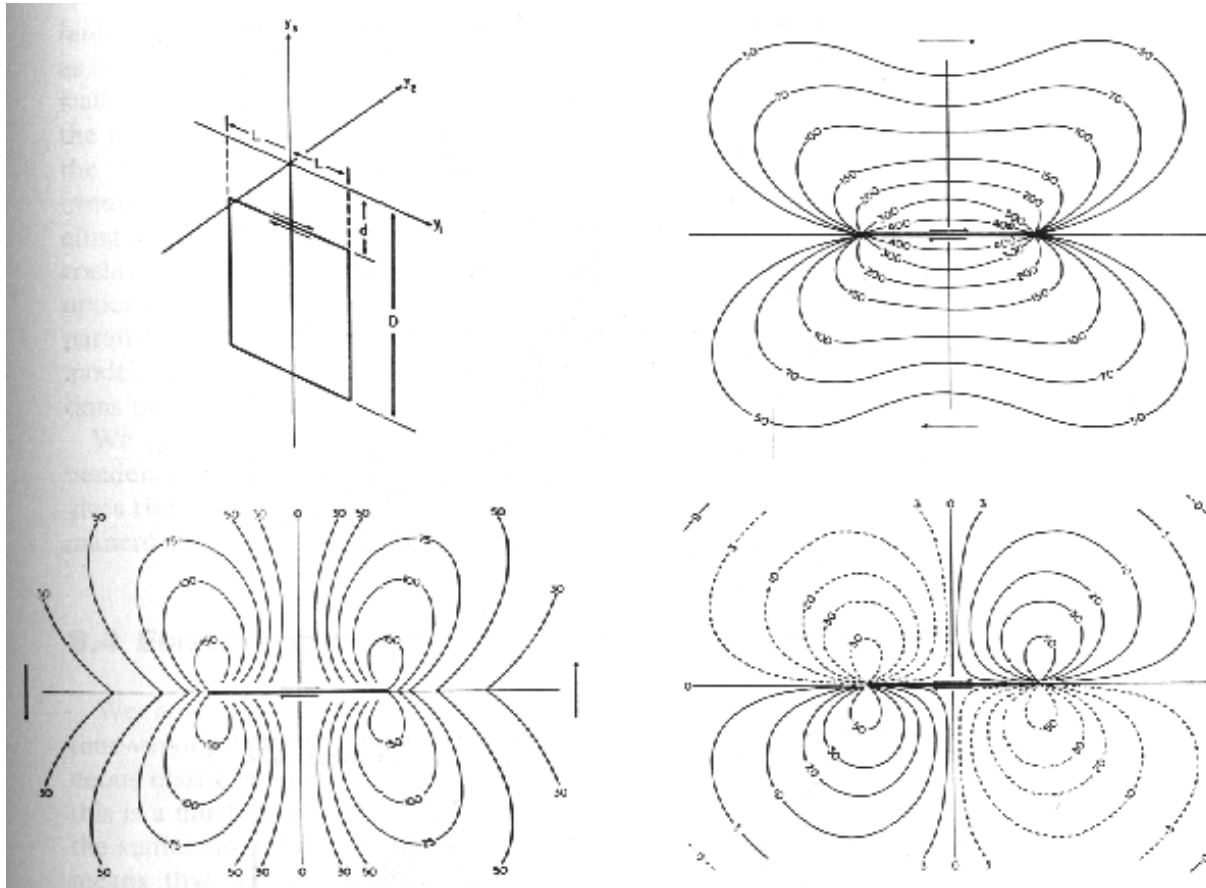
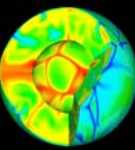
Strike slip



(Results shown for homogeneous model)



Static Displacements



Ground displacement at the surface of a vertical strike slip.

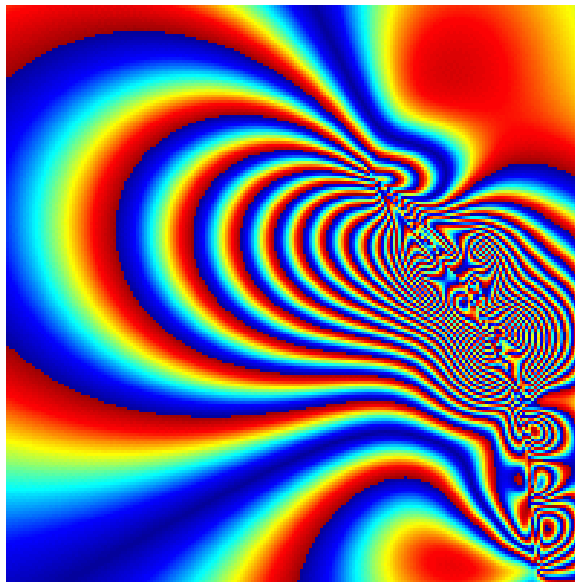
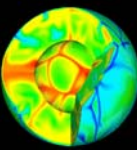
Top right: fault parallel motion

Lower left: fault perpendicular motion

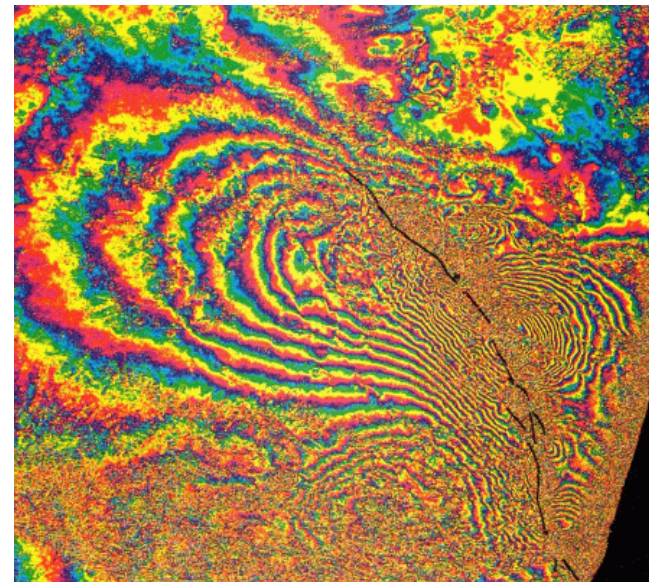
Lower right: vertical motion



Co-seismic deformation



Simulated deformation

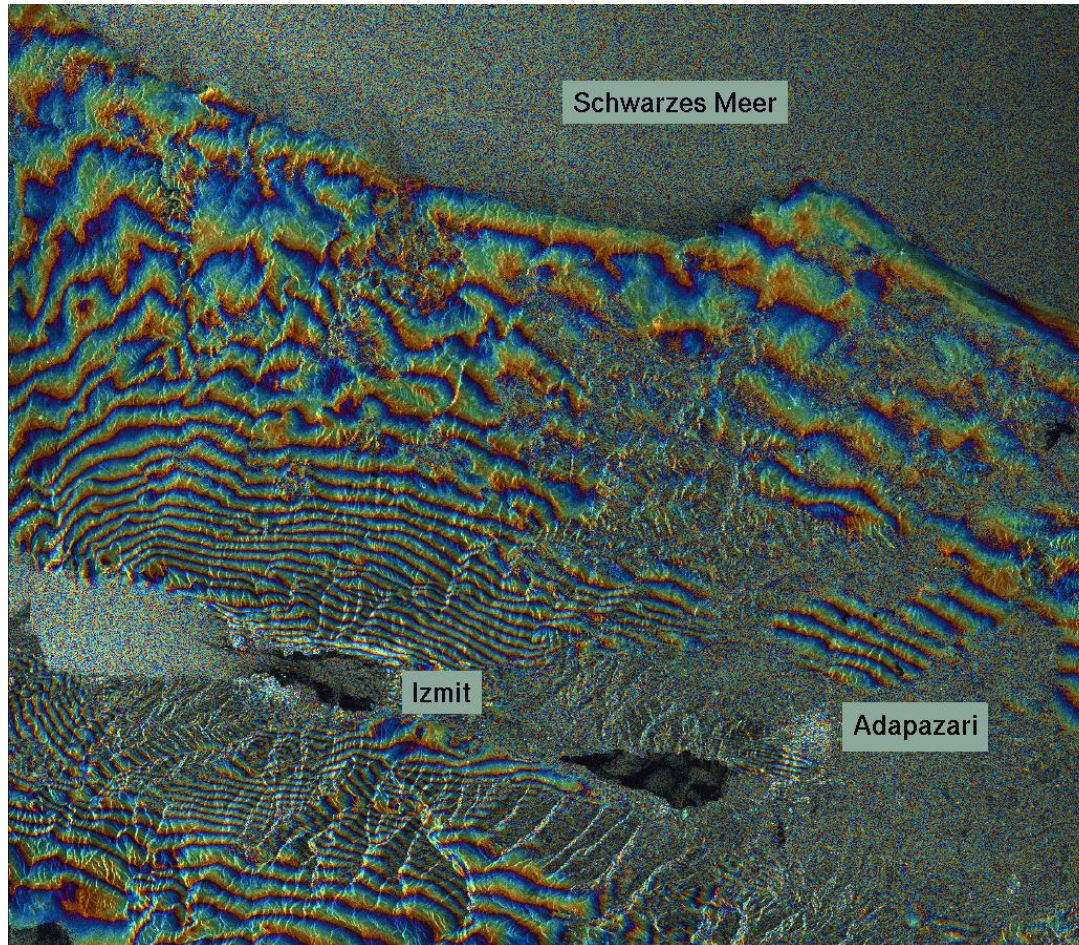
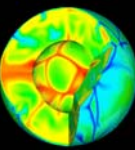


Observed deformation

Source Kim Olsen, UCSB



Static Displacements



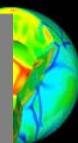
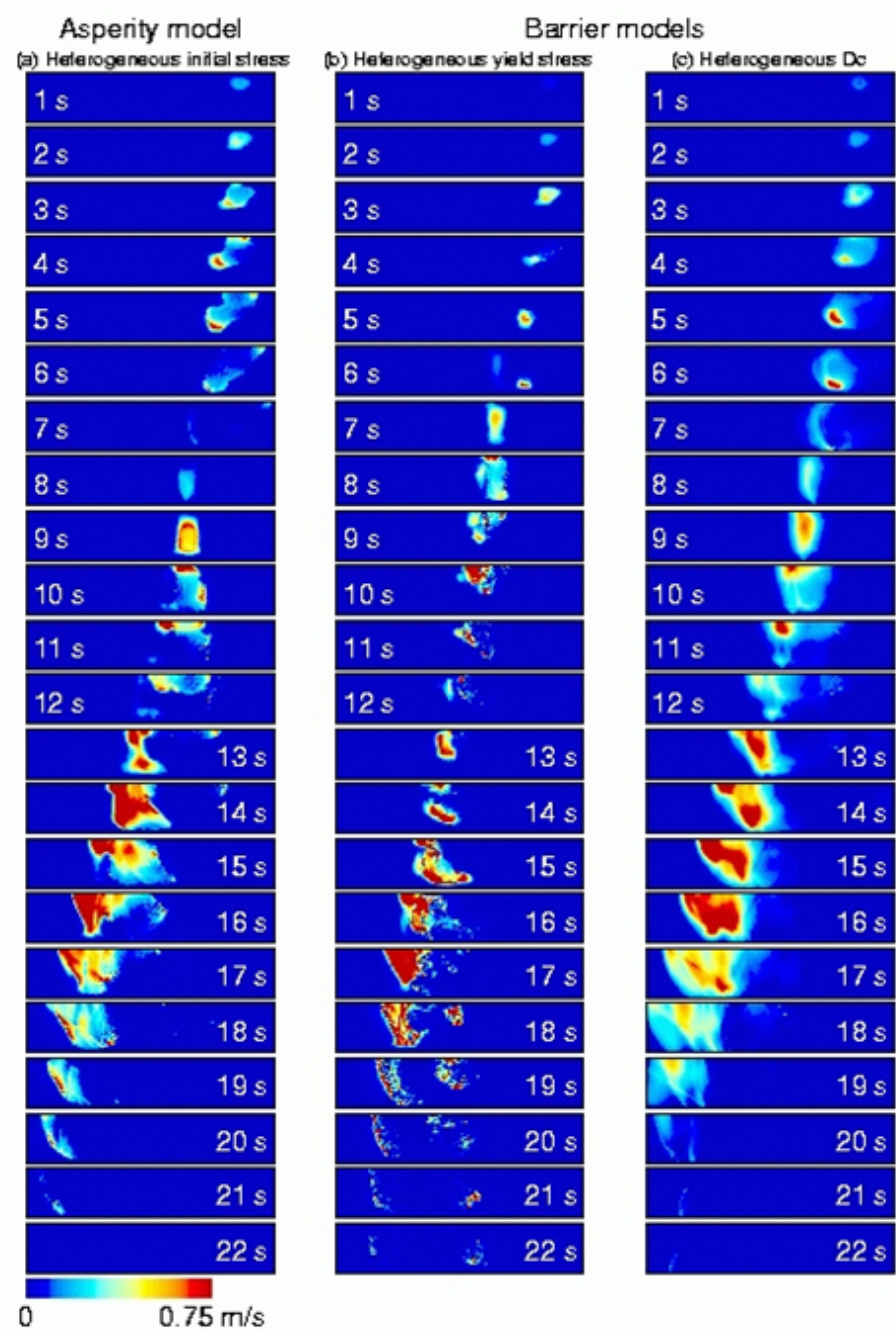
Displacements after Turkey earthquake 1999.



Source kinematics

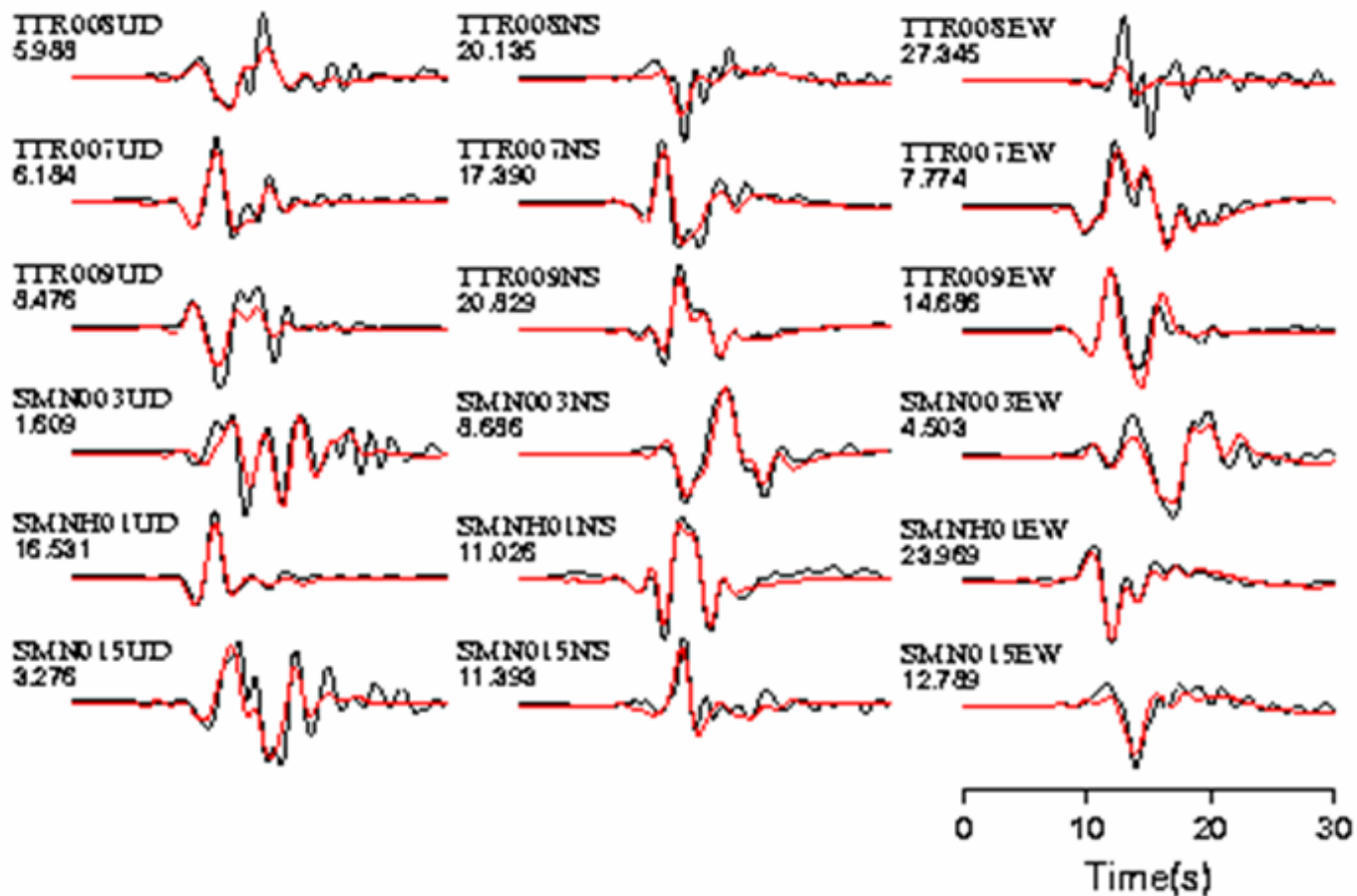
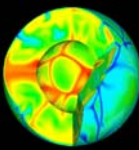
Slip rate as a function of various fault conditions (Landers earthquake)

Source: K Olsen, UCSB



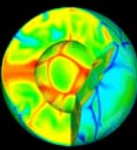


Source kinematics



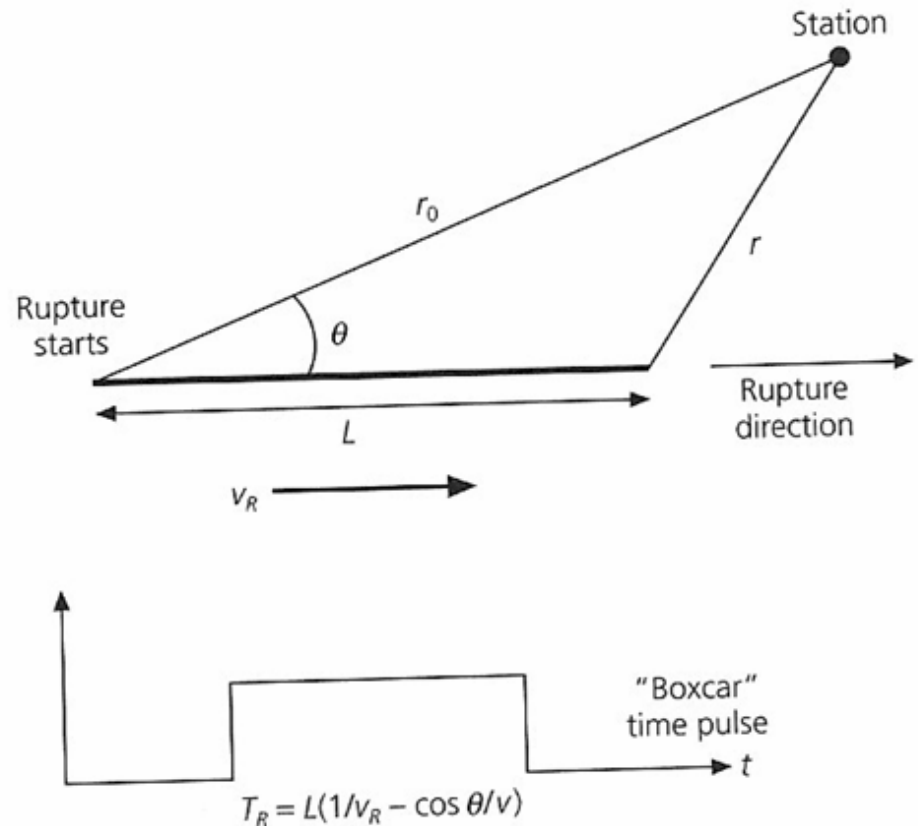


Source directivity



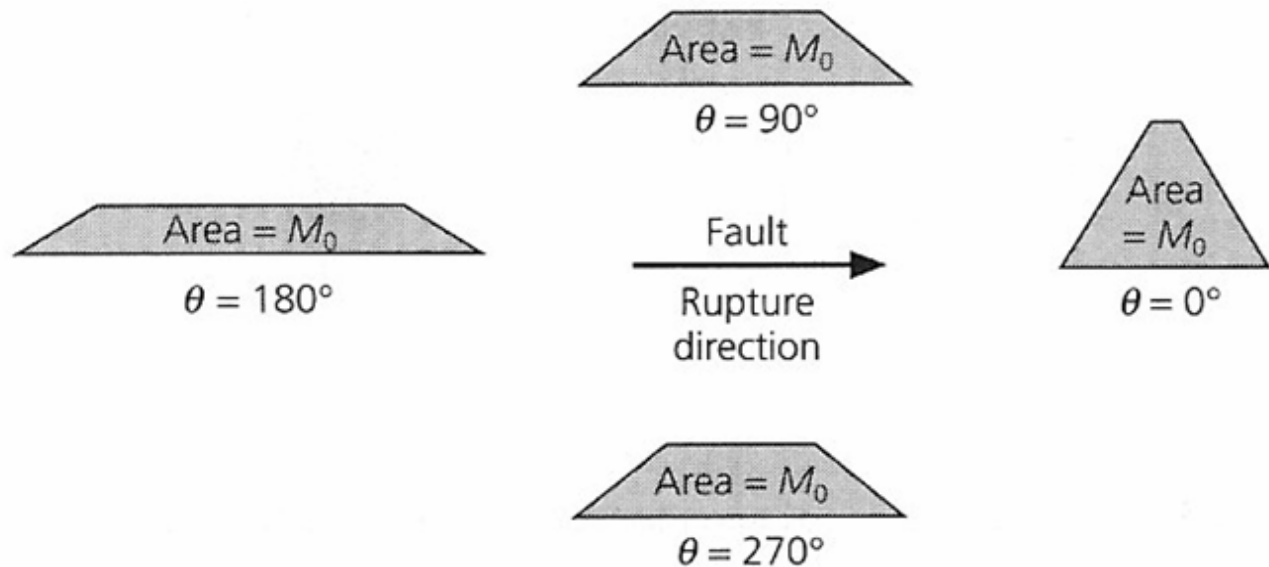
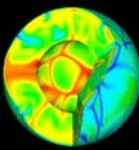
When a finite fault ruptures with velocity v_r , the time pulse is a boxcar with duration

$$T_R = L(1/v_r - \cos(\theta/v))$$





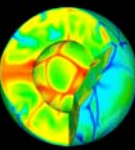
Source directivity



The energy radiation becomes strongly anisotropy (Dopple effect). In the direction of rupture propagation the energy arrives within a short time window.



Source kinematics



Point source characteristics (source moment tensor, rise time, source moment, rupture dimensions) give us some estimate on what happened at the fault. However we need to take a closer look. We are interested in the space-time evolution of the rupture.

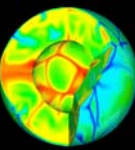
Here is the fundamental concept:

The recorded seismic waves are a superpositions of many individual double-couple point sources.

This leads to the problem of estimating this space-time behavior from observed (near fault) seismograms. The result is a **kinematic** description of the source.



Seismic moment



Seismologists measure the size of an earthquake using the concept of seismic moment. It is defined as the force times the distance from the center of rotation (torque). The moment can be expressed surprisingly simple as:

$$M_0 = \mu A d$$

M_0 seismic moment

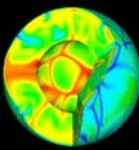
μ Rigidity

A fault area

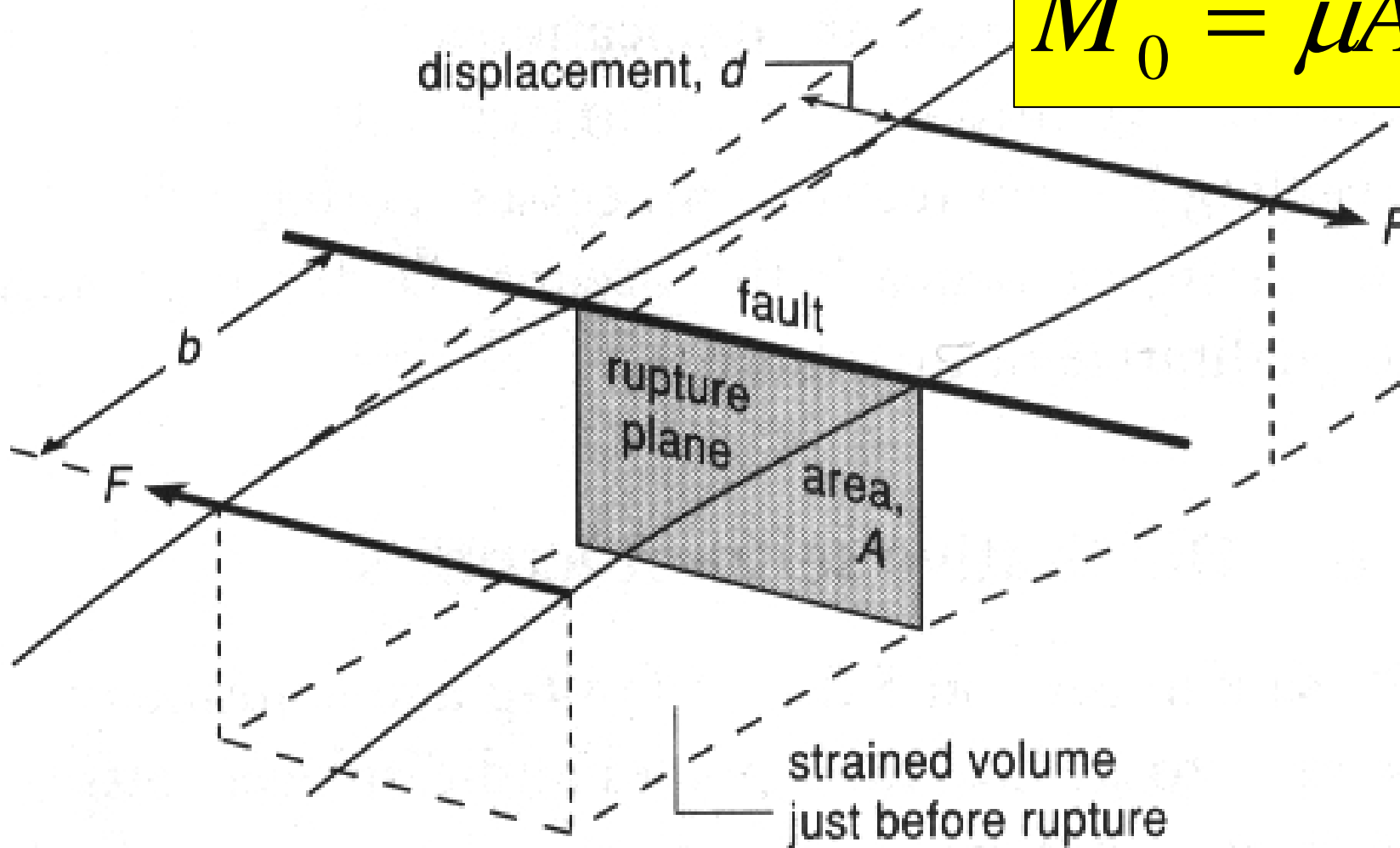
d slip/displacement



Seismic moment

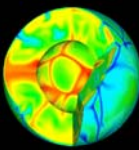


$$M_0 = \mu A d$$

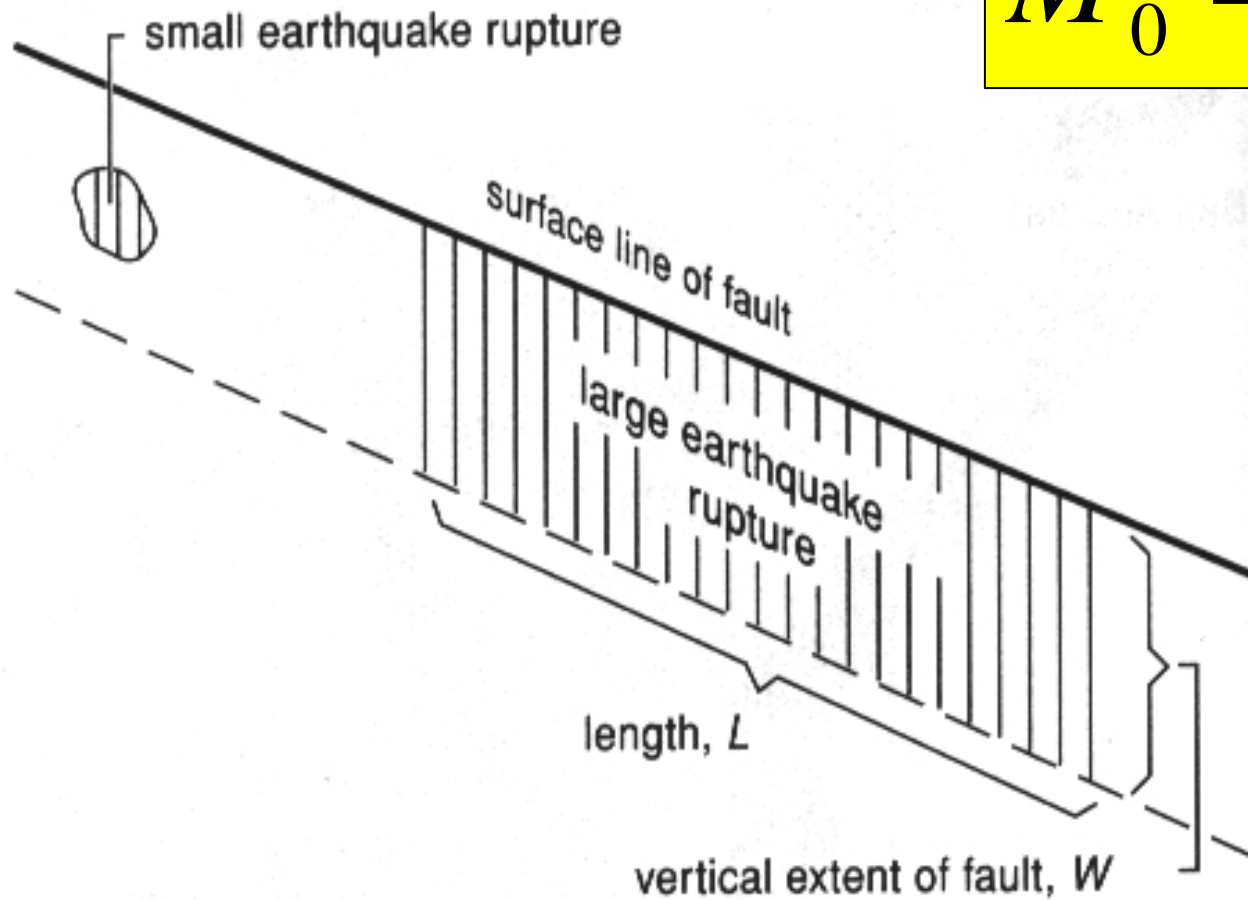




Seismic moment

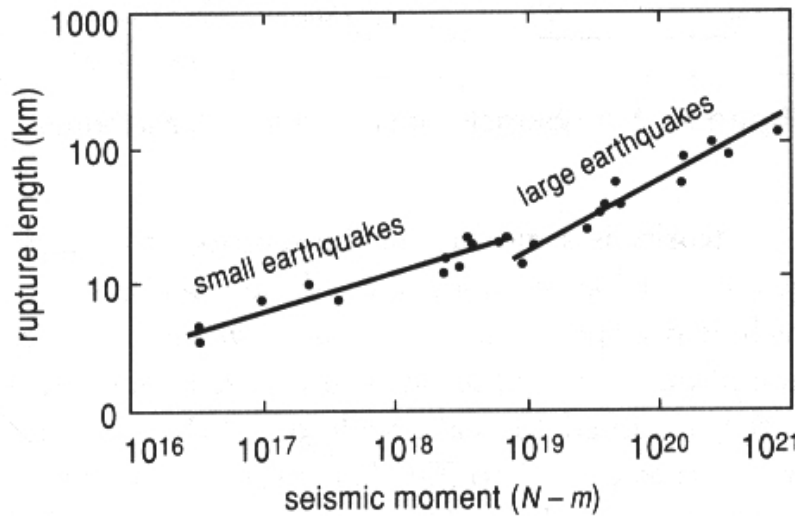
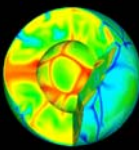


$$M_0 = \mu A d$$



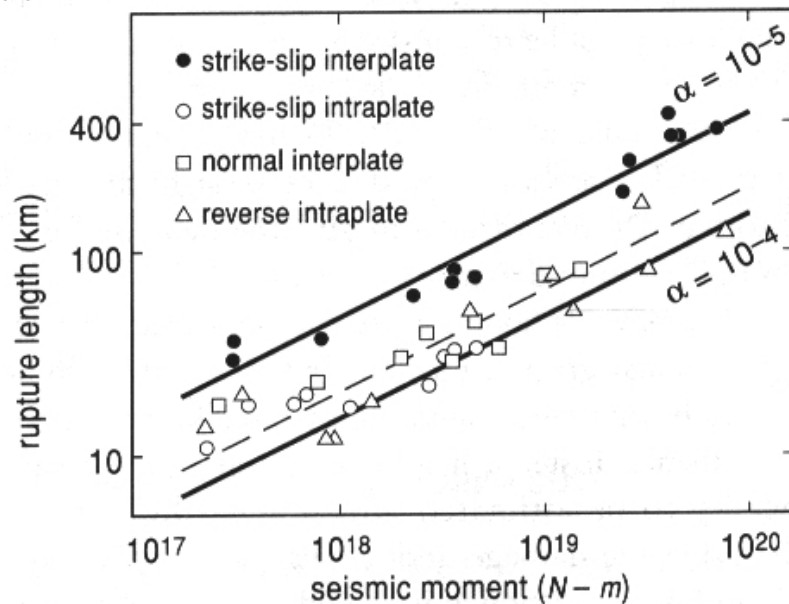


Seismic moment



$$M_0 = \mu A d$$

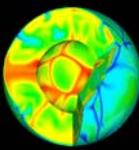
(c)



There are differences in the scaling of large and small earthquakes

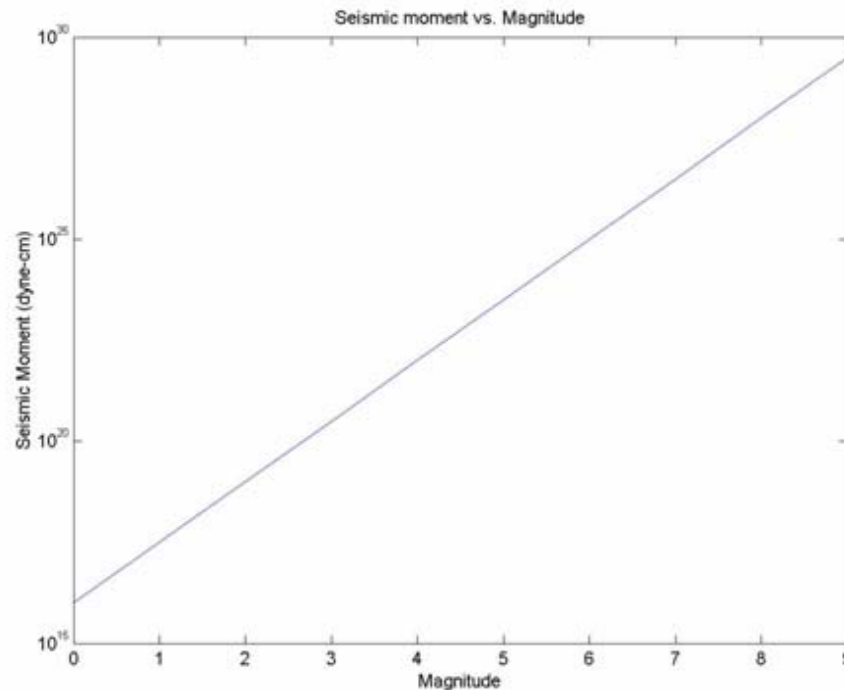


Seismic moment - magnitude



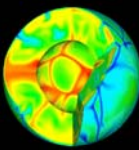
There is a standard way of converting the seismic moment to magnitude M_w :

$$M_w = \frac{2}{3} [\log_{10} M_0 (\text{dyne-cm}) - 16.0]$$





Seismic energy



Richter developed a relationship between magnitude and energy (in ergs)

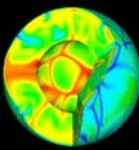
$$\log E_s = 11.8 + 1.5M$$

... The more recent connection to the seismic moment (dyne-cm) (Kanamori) is

$$Energy = Moment / 20000$$



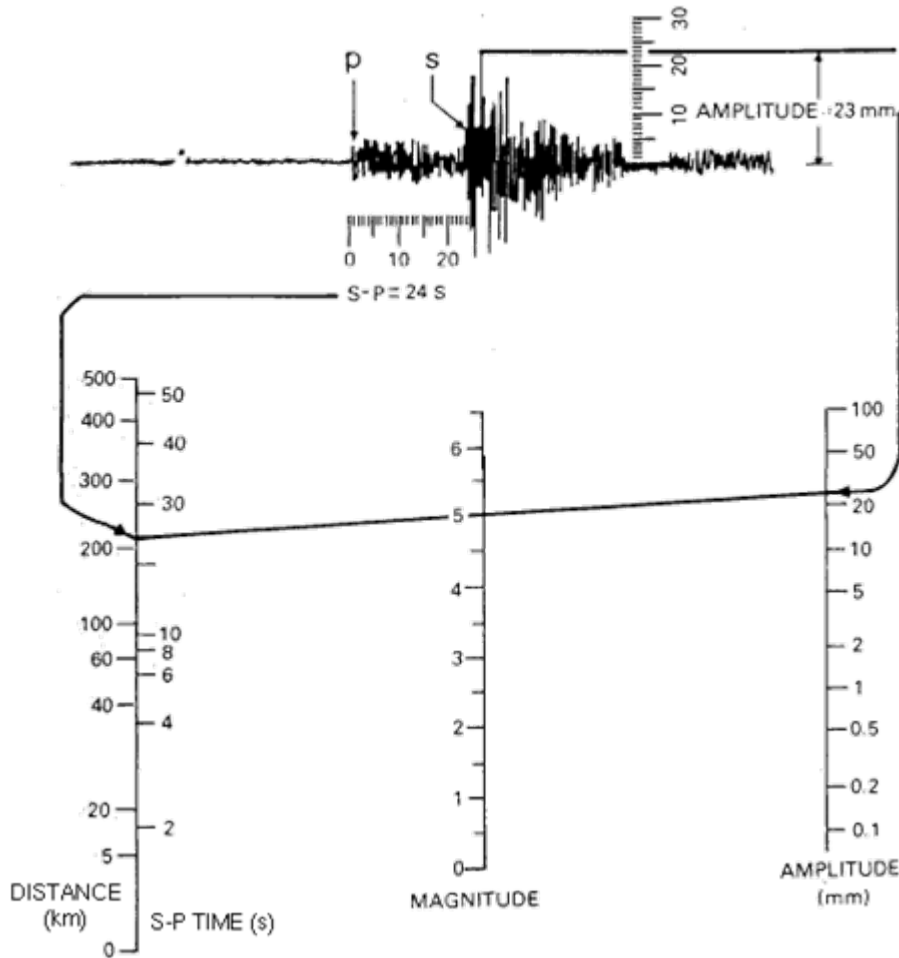
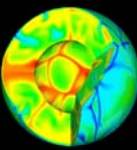
Seismic energy (Examples)



Richter Magnitude	TNT for Seismic Energy Yield	Example (approximate)
-1.5	6 ounces	Breaking a rock on a lab table
1.0	30 pounds	Large Blast at a Construction Site
1.5	320 pounds	
2.0	1 ton	Large Quarry or Mine Blast
2.5	4.6 tons	
3.0	29 tons	
3.5	73 tons	
4.0	1,000 tons	Small Nuclear Weapon
4.5	5,100 tons	Average Tornado (total energy)
5.0	32,000 tons	
5.5	80,000 tons	Little Skull Mtn., NV Quake, 1992
6.0	1 million tons	Double Spring Flat, NV Quake, 1994
6.5	5 million tons	Northridge, CA Quake, 1994
7.0	32 million tons	Hyogo-Ken Nanbu, Japan Quake, 1995; Largest Thermonuclear Weapon
7.5	160 million tons	Landers, CA Quake, 1992
8.0	1 billion tons	San Francisco, CA Quake, 1906
8.5	5 billion tons	Anchorage, AK Quake, 1964
9.0	32 billion tons	Chilean Quake, 1960
10.0	1 trillion tons	(San-Andreas type fault circling Earth)
12.0	160 trillion tons	(Fault Earth in half through center, OR Earth's daily receipt of solar energy)



Richter Scale

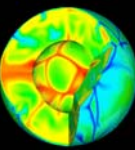


Determination of the magnitude of an earthquake graphically.

$$M_L = \log_{10} A(\text{mm}) + (\text{Distance correction factor})$$



Seismic sources



Far away from the source (far-field) seismic sources are best described as point-like **double couple** forces. The orientation of the initial displacement of P or S waves allows estimation of the orientation of the slip at depth.

The determination of this **focal mechanism** (in addition to the determination of earthquake location) is one of the routine task in observational seismology. The quality of the solutions depends on the density and geometry of the seismic station network.

The size of earthquakes is described by **magnitude** and the **seismic moment**. The seismic moment depends on the rigidity, the fault area and fault slip in a linear way. Fault scarps at the surface allow us to estimate the size of earthquakes in historic times.